

12

ABSTRACT

13 This review combines fluid inclusion data from (*HP*-)*UHP* rocks with experimental
14 research and thermodynamic models to investigate the chemical and physical properties of fluids
15 released during deep subduction, their solvent and element transport capacity, and the
16 subsequent implications for the element recycling in the mantle wedge. An impressive number of
17 fluid inclusion studies indicate three main populations of fluid inclusions in *HP* and *UHP*
18 metamorphic rocks: i) aqueous and/or non-polar gaseous fluid inclusions (FI), ii) multiphase
19 solid inclusions (MSI), and iii) melt inclusions (MI). Chemical data from preserved fluid
20 inclusions in rocks match with and implement “model” fluids by experiments and
21 thermodynamics, revealing a continuity behind the extreme variations of physico-chemical
22 properties of subduction-zone fluids. From fore-arc to sub-arc depths, fluids released by
23 progressive devolatilization reactions from slab lithologies change from relatively diluted
24 chloride-bearing aqueous solutions ($\pm N_2$), mainly influenced by halide ligands, to (alkali)
25 aluminosilicate-rich aqueous fluids, in which polymerization probably governs the solubility and
26 transport of major (e.g. Si and Al) and trace elements (including C). Fluid inclusion studies point
27 to a reconsideration of the petrological models explaining deep volatile liberation, and their flux
28 into the mantle wedge.

29

30

INTRODUCTION

31

32 The chemical behavior of fluids in deep subduction zones and its implications for element
33 cycling and flux melting in the mantle wedge, which causes the formation of arc magmas, has
34 been a subject of growing interest in the last twenty years (e.g., Manning 2004; Bebout 2007,
35 2013). Experimental and theoretical research represents a major tool for understanding the
36 chemical and physical properties of slab-derived fluids, and for modeling the metamorphic
37 evolution of deep subducting rocks (e.g., Ulmer 2001; Poli and Schmidt 2002; Manning 2004;
38 Sanchez Valle 2013; and references cited).

39 Closely related to theoretical and experimental research are studies of fluid inclusions in
40 metamorphic rocks. High pressure (*HP*) and ultra-high pressure (*UHP*) eclogite-facies rocks of
41 crustal origin represent an excellent natural laboratory for the study of subduction-zone fluids,
42 since they underwent pressure and temperature conditions comparable to those hypothesized to
43 occur in deep subducting slabs (Carswell and Compagnoni 2003). A large number of fluid
44 inclusion studies in eclogites and related rocks have reported a remarkable variety of fluid types:
45 aqueous fluids with variable halide content \pm non-polar gases (e.g., N_2 , CO_2 , CH_4),
46 aluminosilicate-rich aqueous fluids, intermediate between silicate melt and water, and hydrous
47 silicate melts (c.f., Scambelluri and Philippot 2001; Touret and Frezzotti 2003; Ferrando et al.
48 2005; and references cited). Many among these inclusions have lost their original composition,
49 but some still preserve the chemistry of fluid trapped during prograde and peak *UHP*
50 metamorphic conditions, despite subsequent exhumation from extreme depths. In this paper, we
51 highlight how fluids in inclusions can constrain and explain deep subduction processes. After
52 providing a review of current research on fluid inclusions in (*HP*-) *UHP* metamorphic rocks, we
53 express our view on the chemical and physical characteristics of subduction-zone fluids
54 preserved as inclusions, including solvent and element transport capacity, and implications to

55 cycling of volatiles. Further recent discussion on fluid inclusions in UHP metamorphic rocks can
56 be found in Scambelluri and Philippot (2001), Touret and Frezzotti (2003), Ferrando et al.
57 (2005), Zheng et al. (2011), Klemd (2013), and Hermann et al. (2013).

58

59 **FLUID INCLUSIONS IN HP-UHP METAMORPHIC ROCKS**

60

61 Fluid inclusions are tiny volumes of mobile volatile-rich phases trapped in minerals
62 during, or after their growth (Roedder 1984). For this reason, they represent the only possible
63 way to acquire firsthand information on naturally occurring fluids. A digest of current main
64 research on fluid inclusions in UHP eclogite-facies metamorphic suites is provided in Table 1. It
65 summarizes different pieces of information, including inclusions' main textural (stage of
66 trapping, inclusion type, and host minerals) and chemical (daughter minerals, water vs. gas)
67 characteristics, along with the nature of host metamorphic rocks (metamorphic belt, peak P - T
68 conditions, matrix lithologies). A synthetic report of the interpretations proposed by the single
69 authors is included. Most examples are on primary fluid inclusions (i.e., fluid inclusions trapped
70 during growth of the host mineral; Roedder 1984) in peak metamorphic minerals of crustal
71 rocks, that, as a consequence, preserve the composition of the hydrous fluid phases released by
72 the slab at more than 90 km depth, during deep subduction (i.e., P - T stability field of coesite).

73 Three distinct populations of fluid inclusions are observed in HP and UHP
74 metamorphic rocks (Table 1): i) chloride-bearing aqueous and/or non-polar gases fluid
75 inclusions (FI), ii) multiphase solid inclusions (MSI), and iii) melt inclusions (MI). We will
76 deliberately be avoiding the term '*polyphase* inclusion' often reported in literature as an
77 alternative to MSI (see, Table 1), since this definition applies to any inclusion containing

78 more than two phases (e.g., a fluid inclusion containing liquid H₂O, liquid CO₂, and gaseous
79 CO₂ is defined as a polyphase inclusion) and may cause confusion.

80

81 **Chloride-bearing aqueous, and/or non-polar gaseous fluid inclusions (FI)**

82 Chloride-bearing aqueous fluid inclusions (FI) are observed in *HP* and *UHP* eclogite
83 (Svensen et al. 1999, 2001; Xiao et al. 2000, 2002; Fu et al. 2001; Shen et al. 2003; Ferrando et
84 al. 2005a; Mukherjee and Sachan 2009), eclogitic veins and segregations (Philippot and
85 Selverstone 1991; Selverstone et al. 1992), quartzite and whiteschist (Philippot et al. 1995;
86 Ferrando et al. 2009), Mn-rich garnet nodules (Frezzotti et al. 2011), and garnet pyroxenite (Fu
87 et al. 2003a; Table 1). FI are small (< 40-50 μm in diameter) and show irregular, rounded, or
88 negative crystal shapes. Although some biphasic (liquid + vapor) liquid-rich FI are reported (e.g.,
89 Ferrando et al. 2005b), most of them are three-phase (solid + liquid + vapor) inclusions with
90 about 40-90% of the total volume consisting of liquid + vapor (Figs. 1b - d).

91 As evident from Table 1, one (e.g., halite) to three chlorides are generally observed (Fig.
92 1c), often associated with carbonates (e.g., calcite, Mg-calcite, dolomite, hydrous carbonates),
93 and/or minor sulfates (mainly gypsum), sulfides (mainly pyrite), and oxides (rutile, ilmenite).
94 More rarely, phosphate (apatite, monazite) and silicate (amphibole, mica, epidote, quartz,
95 feldspars, kyanite, omphacite, ellenbergerite, glaucophane; Fig. 1d) daughter minerals have been
96 described. Interestingly, FI in garnet nodules from Lago di Cignana (Italian western Alps)
97 contain several carbon phases (carbonate, diamond, graphite, carbonaceous material) and an
98 aqueous fluid bearing HCO₃⁻, CO₃²⁻, and SO₄²⁻ ions, and silica monomers in solution, but no
99 chlorides (Fig. 1b).

100 Non-polar gaseous fluid inclusions are observed in eclogite (Fu et al. 2001; Xu et al.
101 2006; Mukherjee and Sachan 2009) and garnet pyroxenite (Fu et al. 2003a; Table 1). These tiny

102 inclusions (< 10 μm in diameter) are single-phase pure (N_2 , or CH_4) or binary ($\text{N}_2\text{-CH}_4$, rarely
103 $\text{CO}_2\text{-N}_2$) mixtures, gray to dark in color due to their low refractive index. Commonly, they are
104 associated with chloride-bearing aqueous fluid inclusions (Fu et al. 2001). Mixed $\text{N}_2\text{-CO}_2\text{-CH}_4\text{-}$
105 $\text{H}_2\text{O} \pm$ solids inclusions (Fig. 1a) are reported from HP eclogite (Fu et al. 2003b; Xiao et al.
106 2001, 2002), garnet pyroxenite (Fu et al. 2003a), quartzite (Frezzotti et al. 2007), and whiteschist
107 (Philippot et al. 1995; Table 1). The gas bubble volume is variable from < 20 to > 80 % of the
108 inclusion. These inclusions occur isolated (Frezzotti et al. 2007), or in association with chloride-
109 bearing aqueous fluid inclusions (Fu et al. 2003a). Halite is the most common mineral phase in
110 these inclusions, although gypsum, anhydrite, and pyrite have been locally reported (Table 1).

111

112 **Multiphase solid inclusions (MSI)**

113 Multiphase solid inclusions (MSI; Table 1) are peculiar primary inclusions present in
114 peak mineral assemblages of UHP rocks, recognized in eclogite (Ferrando et al. 2005 a, 2005b)
115 and eclogitic veins (Zhang et al. 2008), felsic gneiss (Stöckhert et al. 2001, 2009;
116 Dobrzhinetskaya et al. 2003a, 2003b, 2012), quartzite (Frezzotti et al. 2007; Ferrando et al.
117 2005b), Grt-Cpx-Qtz rocks (Hwang et al. 2003), whiteschist (Philippot et al. 1995; Ferrando et
118 al. 2009), marble (Hwang et al. 2005, 2006), garnet pyroxenite (van Roermund et al. 2002;
119 Carswell and van Roermund 2005; Vrijmoed et al. 2008; Malaspina et al. 2010), garnet peridotite
120 (Malaspina et al. 2006, 2009), and Ol-Opx rocks (Scambelluri et al. 2001).

121 MSI show variable sizes (from 5 to ca. 100 μm in diameter), and polygonal or negative
122 crystal shapes (Figs. 2 and 3). At room temperature, they consist of aggregates of 4 to 10 different
123 solid phases of relatively large sizes, with subordinate fluid, or empty cavities (< 40% of the total
124 volume of the inclusions; Figs. 2a and 2b). Most of the solid phases are hydrous (alkali)
125 aluminosilicate minerals (white micas, talc, chlorite, phlogopite, biotite, amphiboles, epidote *s.l.*,

126 ellenbergerite, staurolite), though anhydrous silicates (quartz, feldspars, zircon, kyanite,
127 pyroxenes, titanite, olivine, kalsilite) are also observed (Table 1). Carbonates (mainly calcite,
128 Mg-calcite, dolomite, magnesite) constitute common, subordinate phases. Minor sulfates (mainly
129 gypsum and anhydrite), or sulfides (mainly pyrite), phosphates (mainly apatite and monazite),
130 and oxides (mainly rutile and spinel *s.l.*) can also be present (Table 1). Reduced carbon phases
131 (diamond, graphite, disordered carbonaceous material) are observed in some localities (Stöckhert
132 et al. 2001; van Roermund et al. 2002; Dobrzhinetskaya et al. 2003a, 2003b, 2012; Carswell and
133 van Roermund 2005; Vrijmoed et al. 2008; Stöckhert et al. 2009; Malaspina et al. 2010; Table
134 1) Chlorides are rarely present (Philippot et al. 1995; Hwang et al. 2005, 2006; Ferrando et al.
135 2009; Table 1).

136 In MSI, the relatively large solid phases, in particular lamellar minerals, show intimate
137 intergrowths and preferred orientations, often resulting from periodic crystallization (Figs. 2a,
138 2c, 2d, and 3). Textural relationships of daughter minerals are different from those observed in
139 crystallized MI in plutonic and in *HT* metamorphic rocks (i.e., “stone cavities” in granites, and
140 “nanogranites” in migmatites; Sorby 1858; Frezzotti 1992; Cesare et al. 2009). In MSI,
141 crystallization of daughter phases appears to have started in the center of the inclusion cavity and
142 proceeded to the walls, exponentially increasing mineral surface. Crystallization seems to have
143 occurred from a homogeneous relatively low-viscosity melt or fluid, by processes such as
144 diffusion mass transport and growth (e.g., periodic precipitation; Liesgang phenomena). Locally,
145 granoblastic textures, and rare developments of triple joints among equant and prismatic
146 minerals are also observed (Figs. 2b).

147 In contrast to the previous class of FI, it is difficult to optically detect liquid water in MSI
148 (Figs. 2 and 3) due to: i) the high amount of solid phases in the inclusions, ii) the high density of
149 the trapped fluids, and iii) the possibility that water passively diffused from inclusions during the
150 retrograde evolution of the host rocks (see following paragraphs). Microthermometric analyses

151 collected in MSI with a relatively high amount of fluid indicate a solute-rich aqueous solution
152 (e.g., Zhang et al. 2005). Non-polar gas species, such as N₂ and/or CO₂ have not been detected
153 by Raman microspectroscopy (Table 1).

154

155 **Melt inclusions (MI)**

156 Melt inclusions (MI) can occur in quartz-feldspatic rocks and garnet-bearing gneiss
157 (Hwang et al. 2001, 2006; Mposkos et al. 2009), metapelite (Lang and Gilotti 2007), Grt-Cpx
158 rocks (Hwang et al. 2004, 2005), marble (Korsakov and Hermann 2006), eclogite (Shen et al.
159 2003; Gao et al. 2012), garnet peridotite (Naemura et al. 2009; Table 1). They form by
160 decompression melting of eclogite-facies matrix rocks during early retrogression, or by local *in-*
161 *situ* dehydration-melting of hydrous mineral phase inclusions within prograde and peak minerals
162 (see also, Klemd 2013).

163 In general, MI in UHP metamorphic rocks are recrystallized, similarly to what observed
164 in plutonic or in *HT* metamorphic rocks, although, most unexpectedly, minor volumes of glass
165 have been reported in a few cases (e.g., Hwang et al. 2004; Hwang et al. 2006; Table 1). During
166 slow cooling in the Earth's crust, crystallization of silicate phases inside MI typically initiates on
167 the cavity wall, and the nucleation rate is greater than the growth rate, resulting in the formation
168 of a number of randomly-oriented fine-grained, anhedral to subhedral crystals of similar size
169 (Fig. 4; Frezzotti 2001). In UHP rocks, MI are generally small (ca. 10-50 μm in diameter), with
170 polygonal or negative crystal shapes, generally filled by tiny aggregates of feldspars, micas, and
171 quartz, or, more rarely, of carbonates, and carbon phases ± silicates (Table 1). In particular, these
172 inclusions show magmatic textural relationships, such as an order of crystallization, or graphic
173 intergrowths (e.g., Korsakov and Hermann 2006; Mposkos et al. 2009), similar to (crystallized)
174 MI in plutonic and *HT* metamorphic rocks. Fine-grained mineral assemblages are also observed

175 to partly or totally replace former phengite lamellae because of their *in-situ* dehydration melting
176 (e.g., Gao et al. 2012; Liu et al. 2013). Finally, irregular pockets filled by relatively coarse-
177 grained reactants and products of dehydration melting reactions involving phengite are also
178 reported in UHP rocks (Lang and Gilotti 2007).

179

180 **POST-ENTRAPMENT MODIFICATIONS OF HP-UHP FLUID INCLUSIONS AND** 181 **MULTIPHASE SOLID INCLUSIONS**

182

183 UHP FI and MSI are usually trapped during the growth of the host mineral (primary or
184 early inclusions; Roedder 1984; Touret 2001). They occur isolated or as clusters and, if
185 abundant, they are regularly distributed (Fig. 5a) marking the growth zones of the host mineral
186 (Fig. 5b). The fidelity of the FI and MSI fluid record is based on two fundamental assumptions.
187 The first one is that their chemistry must be representative of the composition of the fluid phases
188 present in the system at trapping conditions (Roedder, 1984). Because of the constant
189 proportions of crystallized minerals (Fig. 2 and 3), most FI and MSI appear to have trapped an
190 homogeneous UHP fluid phase. Exceptions, are represented by those inclusions which contain
191 incidentally-trapped minerals. These last ones are recognizable for their nonsystematic
192 occurrence and, typically, for having larger dimensions than daughter minerals (Fig. 5c).
193 Similarly, in MI, presence of magmatic mineral assemblages, replacing a hydrous phase during
194 its *in-situ* dehydration melting, and/or of reactants and products of a dehydration melting
195 reaction are a strong evidence for incipient anatexis of UHP rocks, but are not representative for
196 the composition of the melt generated during this event.

197 The second assumption requires that, after trapping an homogeneous fluid phase,
198 inclusions behaved as closed systems, i.e. they evolved independently from the host minerals

199 (Roedder 1984; Touret 2001). Post-trapping modifications can affect some inclusions and spare
200 some others nearby (Fig. 6). This depends on the nature of the host mineral, and on the size,
201 shape and location (i.e., dislocations in the host) of the inclusions in the host mineral (Viti and
202 Frezzotti 2000; 2001; Touret 2001; Frezzotti and Ferrando 2007; Stöckhert et al. 2009).

203 As illustrated in Fig. 7, crystallization of daughter minerals (stage 2 of Fig. 7) during
204 decompression does not necessarily imply changes in the chemical composition of the former
205 fluid. In some cases, also the retrograde chemical interaction between the inclusion and the host
206 mineral does not affect the chemical composition of the former fluid. For example, the negative
207 crystal shape shown by most inclusions (Figs. 2, 3 and 6) develops just after the entrapment by
208 dissolution and precipitation of the host mineral on the inclusion walls (stage 2 of Fig. 7). This is
209 an equilibrium process that is not considered to modify the former fluid composition (e.g.,
210 Roedder 1984; van den Kerkhof and Hein 2001; Frezzotti 2001). Similarly, retrograde
211 metamorphic reactions inside MSI (e.g., pseudomorphosis, hydration reactions; Figs. 6 or 8a,
212 respectively; stage 4 of Fig. 7) do not change the chemistry of the former fluid (Frezzotti et al.
213 2012a).

214 In contrast, *non-equilibrium* growth of the host mineral on the inclusion walls (typically
215 with slightly different composition) can modify the chemical composition of both the trapped
216 fluid and the host mineral that crystallizes on the inclusions walls (e.g., Heinrich and Gottschalk
217 1995; Frezzotti 2001). This evolution is commonly observed in MSI (and MI) which were
218 subject to high-*T* conditions during the retrograde evolution (stage 3 of Fig. 7). Retrograde
219 metamorphic reactions between the host mineral and the inclusions on cooling (e.g., hydration
220 reactions) and diffusion of elements from the inclusions, can modify significantly the
221 composition (Fig. 8a and d), and changes in the oxygen fugacity of the inclusion fluid may also
222 occur. Chemically re-equilibrated MSI (Fig. 6 and 8c-f) are easy to recognize, since they show
223 irregular contours typically forming re-entrant angles, and can contain newly formed minerals

224 (i.e., step-daughter minerals; e.g., Svensen et al. 1999; Ferrando et al. 2005b; Stöckhert et al.
225 2001, 2009).

226 Due to the strong decompression during retrograde evolution, decrepitation (i.e., burst by
227 explosion) represents another very common post-trapping process in both FI and MSI. This
228 process modifies the former density of the fluid. Decrepitation occurs when the host mineral is
229 no more able to accommodate the pressure difference in and out of the inclusion ($\Delta P = P_{in} - P_{out}$
230 $\gg 0$; Fig. 7) (e.g., Touret 2001; Franz et al. 2001; Stöckhert et al. 2009). FI in eclogite-facies
231 rocks can re-equilibrate their density at different P - T conditions several times during
232 decompression (Fig. 7; e.g., Touret 2001; Touret and Frezzotti 2003; Frezzotti and Ferrando
233 2007; Stöckhert et al. 2009).

234 In MSI, early UHP-HP retrograde decrepitation (stage 3 of Fig. 7) can be identified by
235 irregular contours and short and large fractures (offshoots) departing from the inclusion corners
236 (Figs. 6, 8a, 8e, 9a, 9c, 9d), similarly to fluid inclusions in mantle xenoliths (Fig. 9b). On the
237 contrary, decrepitation at lower P (stage 5 of Fig. 7), liberating a fluid phase when all minerals
238 have already crystallized (stage 4 of Fig. 7), generally forms: i) star-shaped contours (Figs. 6, 9d-
239 e), and/or ii) haloes of very small FI distributed around the inclusion (Fig. 9c), and/or iii) short
240 trails of tiny secondary FI lining from the decrepitated inclusions (Fig. 8c-d). These petrographic
241 features are similar to those observed in volatile-rich MI in plutonic rocks, undergoing volatile
242 oversaturation and magma degassing during slow cooling after trapping as inclusions (Fig. 9f).

243

244

ELUSIVE AND FUGITIVE WATER

245

246 A careful petrographic study is necessary to select those UHP inclusions potentially
247 representative for the chemical composition of the metamorphic fluids. The determination of the

248 water content of the former fluids is more problematic. In most MSI, water represents an elusive
249 and fugitive component, although empty spaces between hydrous daughter minerals (Fig. 8a)
250 testify for the (former) presence of an aqueous fluid component. A valuable way to study water
251 speciation inside MSI is by spectroscopic techniques. One possible analytical approach is
252 described in Frezzotti et al. (2010). In Fig. 10, synchrotron FT-IR maps show the speciation of
253 water in MSI both as H₂O molecules (Fig. 10b) (i.e., liquid phase; absorption bands in the 3000–
254 3500 cm⁻¹ region, Fig. 10e), and as OH⁻ groups (Fig. 10c) (i.e., in daughter mineral phases;
255 absorption bands in the 3500–3800 cm⁻¹ region, Fig. 10e). Comparing Figs. 10a and 9b, the
256 systematic absorption increase in the 3000–3500 cm⁻¹ region in correspondence of MSI reveals
257 that, in contrast with the optical observation, they may contain some liquid water. Quantitative
258 analyses, however, are not possible.

259 An additional concern is that the low-water content of MSI could result from passive H₂O
260 diffusion during the retrograde *P-T* evolution. MSI can exchange H₂O with the host minerals if
261 the internal $P_{\text{H}_2\text{O}}$ in the inclusion is higher than the external $P_{\text{H}_2\text{O}}$ (or $X_{\text{H}_2\text{O}}$) of the fluids
262 circulating in the rocks (Sterner and Bodnar, 1991), a very common case during retrograde
263 decompression at relatively high temperatures (Fig. 7). If decrepitation does not occur, the
264 inclusion fluid $P_{\text{H}_2\text{O}}$ can passively re-equilibrate with the external fluid conditions. Identifying
265 MSI dehydration, if present at all, would also require spectroscopic investigations. The example
266 in Fig. 10d shows a synchrotron FT-IR map of hydrogen distribution in pyrope (Frezzotti et al.
267 2012a and b). Exponential enrichment of bonded hydrogen in garnet approaching MSI reveals
268 the change from molecular water in MSI to OH⁻ bonds to the anhydrous host. Hydrogen
269 gradients, recorded by symmetrically different colored haloes, represent actual water content
270 variations in garnet, with the exception of the MSI area where the extremely high water
271 concentration in garnet in part results from a MSI contribution (yellow to pink haloes in Fig.
272 10d). Fluid inclusion “dehydration” processes were previously proposed to interpret the origin of
273 some pure CO₂ FI in peridotite xenoliths, formed at depths of about 60 - 80 km and rapidly

298 and Sachan 2009), aluminosilicate-rich aqueous fluid phases (e.g., Ferrando et al. 2005 and
299 references cited), and, more rarely, hydrous melts.

300 Early studies by Philippot et al. (1995) first described a relevant variation in the
301 chemistry of fluids passing from prograde *HP* to peak *UHP* metamorphic conditions in
302 coesite-pyropo-bearing whiteschists from the Dora-Maira massif (Italian Western Alps). They
303 noted that large quantities of dehydration fluids released during prograde metamorphism -
304 predating pyrope and coesite *UHP* assemblage - were preserved as low-solute aqueous FI in
305 relic kyanite within pyrope. Further evidence for dehydration fluids at *UHP* conditions,
306 attending the growth of garnet, was given by MSI in pyrope, which contain phyllosilicates,
307 Mg-phosphate, chlorides, and opaques, but no optically-visible aqueous fluid. Over time, fluid
308 inclusion research combined with petrological investigations further demonstrated that water-
309 rich fluids with variable halide contents ($\pm N_2$ and CH_4/CO_2) generally dominate in *HP*
310 prograde, low-*T* *UHP* peak, and early retrograde metamorphic mineral assemblages, whereas
311 complex aluminosilicate-rich aqueous fluid phases are typically preserved in *UHP* peak
312 metamorphic mineral assemblages at *T* above about 600-650°C.

313 In the following sections, we provide a synthesis of the current knowledge on the
314 chemical composition of fluids in deep subduction zones, based on fluid inclusion literature
315 summarized in Table 1. We concentrate on those eclogite-facies rocks of crustal origin
316 equilibrated at *P-T* conditions relevant for slabs at sub-arc depths (i.e., coesite facies; 2.5-4
317 GPa).

318 *Aqueous fluids* $\pm N_2$

319 In some eclogite-facies rocks from Caledonian, Hercynian, and Alpine *LT-HP*, and -
320 *UHP* metamorphic terranes, the prograde, peak, and early retrograde fluid composition within
321 different rock types is uniform in terms of H_2O , which represents the main component, but

322 variable in terms of solutes, and non-polar components (N₂ contaminated by CO₂, or CH₄), as
323 previously summarized by Scambelluri and Philippot (2001), Touret and Frezzotti (2003),
324 Heinrich (2007), and Klemd (2013).

325 FI studies in eclogites from the Alps indicate that prograde aqueous fluids contain
326 mainly NaCl and MgCl₂, with only subordinate concentrations of CaCl₂ and KCl. Calculated
327 salinities vary from about 3 to 50 in NaCl equivalent wt. % (e.g., Philippot and Selverstone
328 1991; Selverstone et al. 1992; Philippot et al. 1995; Scambelluri et al. 2001). To explain the
329 Na-, and Mg-dominated nature of such *HP* chloride-bearing solutions, two processes have
330 been proposed: recycling of sea-water (Scambelluri et al. 1997), or hydrothermal alteration of
331 the oceanic lithosphere (Philippot et al. 1998). More recently, Sharp and Barnes (2004)
332 presented a model for the generation of mobile high-MgCl₂ aqueous plumes at mantle depths,
333 via breakdown of subducted serpentinites.

334 In the Dabie-Shan and Sulu UHP eclogite-facies continental rocks, prograde and early
335 retrograde metamorphic fluids preserved within FI are generally CaCl₂-dominated (Xiao et al.
336 2000, 2001; Fu et al. 2001, 2002, 2003; Zhang et al. 2005; Frezzotti et al. 2007), and not
337 NaCl-rich as in oceanic rocks. Xiao et al. (2000) and Fu et al. (2001, 2003) reported also the
338 presence of N₂, ± CO₂, or CH₄-rich *HP* fluids. In matrix rocks, the scale of fluid chemical
339 heterogeneity is extremely variable. For example, in metasediments and eclogites from the
340 Dabie-Shan UHP terrane, aqueous fluids have different salinities and, most importantly,
341 different cations in solution, even at the scale of the crystal (Fu et al. 2002). In these fluids,
342 the nature of cations is related to the host mineral phase (e.g. Ca-dominated chlorides in
343 epidote, and Na-, K-chlorides in amphibole). Fluid chemical variability indicates very limited
344 fluid transport before trapping as inclusions, and generally supports an internal origin. Based
345 on the very low δ¹⁸O of minerals containing FI in eclogite-facies continental rocks from
346 Dabie-Sulu, Fu et al. (2002, 2003), Ferrando et al. (2005a), and Frezzotti et al. (2007)

347 proposed an ultimate origin from past meteoric water brought to mantle depths (e.g., Zhang et
348 al. 2011, and references cited).

349 Besides chlorine and alkalis, additional minor solutes in aqueous fluids are Si, Al, Ti,
350 C, and S species. The aluminosilicate content in aqueous fluids increases with increasing
351 pressure, while chlorine shows the opposite trend. For example, prograde fluids formed at
352 about 2.5 GPa and 600°C in the Dora-Maira whiteschists contain Cl, Mg, alkalis, with
353 subordinate Si and Al (i.e., Na- and Mg-chlorides and paragonite as daughter phases in FI;
354 Philippot et al. 1995; Ferrando et al. 2009). At 3.5 GPa and similar temperatures (600°C),
355 peak aqueous fluids in UHP metasediments from Lago di Cignana, also in W Alps, contain
356 dominantly Si, Al, Ti, C, Ca, and alkalis (e.g., quartz, rutile, paragonite, carbonate and
357 diamond as daughter phases in FI; Frezzotti et al. 2011; 2014), while Cl appears as a very
358 subordinate component.

359 From Table 1, calculated solute contents in FI range from less than 3 up to 50 wt. % in
360 NaCl equivalent. Experiments and thermodynamic modeling suggest that aqueous fluids
361 formed at depths below about 90 km are diluted solutions (< 20 wt. % in NaCl equivalent;
362 e.g., Manning 1998; Kessel et al. 2005a; Hermann and Spandler 2008; Spandler and Picard
363 2013; and references cited). Considering that many HP aqueous fluids contain less than 30 wt.
364 % solutes (cf., Table 1), most researchers have interpreted the inclusion fluid record to
365 confirm a dominantly diluted nature of HP aqueous subduction-zone fluids. Fluid-fluid
366 immiscibility processes (Newton and Manning 2010), or passive solute enrichment due to
367 water loss from inclusions, could have brought to the extreme chloride enrichment observed
368 in some FI.

369 HP-UHP FI highlight also the possible presence of non-polar fluid species, such as N₂,
370 CO₂, and CH₄, although this last one is more commonly formed during early decompression
371 (cf., Table 1; Xiao et al. 2000, 2002; Fu et al. 2001, 2002, 2003a, 2003b; Yang et al. 2001). N₂

372 is by far the most abundant non-polar gas in prograde and peak metamorphic fluids (cf.,
373 review by Touret 2001; Frezzotti and Touret 2003; Klemd, 2013). Up to 20 mole % N₂ were
374 measured in *HP* aqueous fluids in eclogites from western Norway, where eclogitization is
375 triggered by channeled fluids in shear zones (Austreheim 1987; Andersen et al. 1989; 1990;
376 1993). Although the thermodynamic properties of the C-O-H-N system are not fully known, it
377 is possible to trace the transition from *HP* to *UHP* conditions of aqueous fluids containing
378 nitrogen. Since the binary system H₂O - N₂ system shows a miscibility gap at $P > 2\text{--}2.2$ GPa
379 at temperatures $\leq 700^\circ\text{C}$ ($X_{\text{N}_2} > 0.2$; Haefner et al. 2002), H₂O-N₂ fluids formed in the *P-T*
380 stability field of coesite are expected to be immiscible, while full H₂O-N₂ miscibility should
381 be expected at lower pressures.

382 N₂ fluids can form by release of ammonium (NH₄⁺) substituting for K⁺ in micas,
383 during prograde dehydration with increasing metamorphism (Hallam and Eugster 1976; Duit
384 et al. 1986; Haendel et al. 1986; Bebout and Fogel 1992; Bebout, 2013). The speciation of
385 nitrogen in C-O-H-N fluids depends on pressure, temperature, fluid composition, and is
386 probably controlled by redox conditions (Andersen et al. 1993; Pöter et al. 2004). It has been
387 shown, that, during cold subduction, almost the complete budget of nitrogen can remain in
388 micas up to 2.5 GPa (Busigny et al. 2003). In addition, boron nitride (BN) and osbornite
389 nitrides (TiN) inclusions in coesite have been reported at higher pressures (Dobrzhinetskaya
390 et al. 2008).

391

392 *(Alkali) aluminosilicate-rich aqueous fluids*

393 MSI in *UHP* subduction rocks ($P > 3$ GPa), formed at peak metamorphic temperatures
394 above about 600°C, typically consist of several hydrous (alkali) aluminosilicates, such as
395 paragonite, muscovite, chlorite, and phlogopite, and variable amounts of anhydrous and
396 hydrous sulfates, carbonates, phosphates, chlorides, and (former) H₂O (Table 1; Figs. 2, 3,
397 and 6). Over the last decade, the growing interest on these volatile-rich microsystems induced

398 to an increase of research and discussion to understand the composition and the properties of
399 the fluid media that were originally trapped. From the compilation of data on daughter
400 minerals in MSI, it is evident that Si, Al, and alkalis, but not halides, are the major
401 constituents of UHP water-rich fluid phases preserved in MSI. Element solubility appears
402 distinct from an aqueous fluid phase at crustal depths (cf., Yardley 2013). Deep subduction
403 fluids contain far too much Al_2O_3 and SiO_2 , which are known to have very limited solubility
404 in water at crustal conditions, irrespective of salinity. Nonetheless, these water-rich phases
405 have major element contents that do not resemble any silicate melts of magmatic composition,
406 generated by volatile-saturated partial melting of metasedimentary or eclogitic rocks at high
407 metamorphic grades.

408 A reconstruction of the major element composition of UHP water-rich phases has been
409 attempted by Ferrando et al. (2005, 2009), and Frezzotti et al. (2007) averaging the vol. % of
410 daughter minerals in MSI (EMPA analyses; for the procedure see Ferrando et al. 2005). The
411 method suffers of uncertainties, in part due to the processes that may have affected inclusions
412 during retrograde decompression, including chemical interaction with the host mineral phase
413 (Heinrich and Gottschalk 1995; Svensen et al. 1999; Franz et al. 2001). Nevertheless, semi-
414 quantitative analyses are sufficient for petrological investigations and provide constrains on
415 the major element composition of water-rich phase released during deep subduction.

416 There is an evident link between the nature of chemical species present in UHP fluids
417 and the chemistry of the matrix metamorphic rocks which undergo progressive
418 devolatilization reactions. A nice example is represented by the composition of inclusion
419 fluids in UHP eclogite-facies metamorphic series from the Sulu terrane, which consist of both
420 sedimentary and igneous lithologies attaining the same P - T conditions (3.5 GPa, 750 °C). In
421 UHP quartzite (phengite and epidote *s.l.* not stable) reconstructed peak fluid compositions
422 indicate major SiO_2 (24 wt.%), Al_2O_3 (30 wt.%), CaO (9 wt.%), K_2O (5 wt.%), Na_2O (3

423 wt.%), and SO_3 (11 wt.%), with subordinate TiO_2 , Fe_2O_3 , FeO , MgO , BaO , P_2O_5 , Cl , F , and
424 $(\text{CO}_3)^{2-}$. In eclogites, fluids have similar SiO_2 (26 wt.%), Al_2O_3 (20 wt.%), and CaO (7 wt.%),
425 but considerably higher MgO (4 wt.%), $\text{FeO}_{(\text{tot})}$ ($\text{FeO} = 6$ wt.% $\text{Fe}_2\text{O}_3 = 4$ wt.%), and P_2O_5 ,
426 and almost no alkalis, carbonates, and sulfates. More evidence supporting that UHP aqueous
427 phases can dissolve rock components can be found in the chemistry of fluids formed during
428 the growth of the large pyropes in Dora-Maira metasomatic whiteschists (phlogopite and talc
429 not stable; phengite stable; 4 GPa, 700 °C), which contain high MgO (25.5 wt%), SiO_2 (29
430 wt%), and Al_2O_3 (23 wt%), while extremely subordinate Fe_2O_3 (4.7 wt%), S (3.2 wt%), Na_2O
431 (1.3 wt%), CaO (0.6 wt%), P_2O_5 (0.4wt%), and K_2O (0.2 wt%).

432 A minimum water content of about 20 wt. % has been calculated considering the
433 amount of water bonded in daughter hydrous minerals in MSI, and hypothesizing that the void
434 volume between different daughter phases originally contained H_2O with a density of 1.3
435 g/cm^3 (cf., Ferrando et al. 2005 and references cited; Fig. 2, 3, 8, and 9). The original water
436 content of deep subduction fluids, however, should have been considerably higher, perhaps in
437 the order of 40-60 wt. %, in part lost by passive H_2O diffusion from inclusions during
438 retrogression. Chlorine contents are low, below a few unit wt. %, and CO_2 , N_2 and CH_4 have
439 not been detected (e.g., Raman spectroscopy analyses).

440 Extending the examples discussed above with literature data in Table 1 shows that
441 progressive devolatilization reactions in rocks control the chemistry of aqueous fluids released
442 during deep subduction, depending on rock composition and P - T subduction path, as
443 predicted by theoretical and experimental petrology (e.g., Schmidt and Poli 1998, 2003;
444 Hermann 2002a; Hacker et al. 2003; Hermann et al. 2006). As illustrated in Figure 11, the
445 breakdown of phengite governs the alkali enrichment observed in alkali aluminosilicate
446 aqueous fluid phases in Sulu metasediments, while the Mg-enriched nature of Dora-Maira

447 UHP fluids derives from phlogopite and talc breakdown in these metasomatic rocks of crustal
448 origin.

449 In summary, the emerging feature from FI and MSI studies (Table 1) is that at *HP*
450 conditions, subduction zone fluids are relatively diluted aqueous solutions (\pm non-polar
451 gases), containing chlorides, alkalis, Si, and Al, with properties similar to crustal aqueous
452 fluids. At greater sub-arc depths, fluids released by crustal lithologies contain increasing
453 amounts of alkali aluminosilicate components \pm carbonates phosphates, and sulfates/sulfides,
454 while halides do not appear anymore more as the principal ligands. Thus, although we must
455 acknowledge the limitations of quantitative major element analyses in FI and MSI, observed
456 chemical variations provide key insights into element solubility in deep-subduction aqueous
457 fluids, which results from major structural variations, as it will be discussed in the next
458 section.

459

460 **Experimental approach: Structure and solvent capacity of *HP/UHP* fluids**

461 Knowledge of deep subduction-zone fluids behavior implicitly demands for
462 thermodynamic models. When considering the chemical composition of FI and MSI, it is
463 clear that water is the relevant volatile at sub-arc depths. However, water-rich phases
464 containing a great silicate concentration (i.e., MSI) have properties which are intermediate
465 between those generally attributed to aqueous fluids and to hydrous silicate melts, an obvious
466 avenue for experiments and theoretical models.

467 Experimental and theoretical studies on quartz solubility in water have shown that at
468 high pressure, the silica content of aqueous fluids rises considerably, and the silica-saturated
469 fluid phase and the water-saturated melt progressively approach each other (Fig. 12; Newton
470 & Manning 2008; Mysen 2010; Dolejš and Manning 2010; Hunt and Manning 2012). In the

471 SiO₂-H₂O system, immiscibility between melt and fluid vanishes at the upper critical endpoint
472 (point C in Fig. 12), as aqueous fluids form a single continuous solution with silicate melts
473 (Kennedy et al. 1962; Anderson and Burnham 1965; Paillat et al. 1992; Ryabchikov 1993;
474 Bureau and Keppler 1999; Newton and Manning, 2008; Stalder et al. 2000, 2001). The *P-T*
475 location of the upper critical endpoint for the SiO₂-H₂O system was determined at 1 GPa and
476 1080°C by Kennedy et al. (1962) and Anderson and Burnham (1965). In natural systems, the
477 *P-T* conditions of the upper critical endpoint depend on the chemistry of the silicate system.
478 For example, they correspond to 2 GPa and 700 °C for the NaAl-Si₃O₈-H₂O system (Paillat et
479 al. 1992; Stalder et al. 2000). For more complex systems, such as basalt-H₂O, Ca-bearing
480 granite-H₂O, haplogranite-H₂O, jadeite-H₂O, nepheline-H₂O, supercritical *P-T* conditions
481 have been proposed at pressures variable from 3.4 to 6.0 GPa, with currently no general
482 consensus (for discussion cf., Eggler and Rosenhauer 1978; Ryabchikov 1993; Bureau and
483 Keppler 1999; Mysen and Shang 2003; Mibe et al. 2011).

484 Based on thermodynamics, the nature of UHP fluids in subduction zones has been
485 discussed at great length in the literature, and several definitions have been applied:
486 intermediate fluids (Manning 2004; Frezzotti et al. 2007; Ferrando et al. 2009), supercritical
487 liquids (Schmidt et al. 2004; Kessel et al. 2005b; Mysen 2012; Spandler and Pirard 2013),
488 supercritical fluids (Ferrando et al. 2005a; Malaspina et al. 2006; Hack et al. 2007a; Liebscher
489 2010; Zhang et al. 2011; Zheng et al. 2011; Kawamoto et al. 2012), supercritical hydrous
490 melts and transitional fluids (Hermann et al. 2006, 2013). Recently, Sanchez-Valle (2013)
491 proposed to adopt the term *fluid* for any water-rich phases released in subduction zones at
492 sub-arc depths, since: “*the meaningful distinction of liquid, vapor, gases at ambient*
493 *conditions becomes obsolete at high pressures and temperatures where transitions between*
494 *the various fluid types are often gradational*”.

495 At high pressure and temperature conditions, the (supercritical) fluid vs. melt *vexata*
496 *quaestio* is more a semantic than a petrological controversy. As summarized by Hack et al.
497 (2007b) most subduction geothermal gradients are likely to reach P - T conditions close or
498 above the upper critical endpoints of most crustal lithologies. Thus, during deep subduction,
499 the hydrous *solidus* should be seen as a convenient temperature boundary at which the
500 concentration of silicate components dissolved into aqueous fluids progressively increases,
501 without the abrupt chemical changes resulting at crustal P - T conditions (cf., Fig. 11).

502 FI and MSI demonstrate beyond reasonable doubt the relevant and gradual increase of
503 the (alkali) aluminosilicate solute load, without phase change, predicted in model fluids by
504 thermodynamics at increasing P (Fig. 12; e.g., Bureau and Keppler 1999; Hunt and Manning
505 2012). For example, at depths greater than about 90-100 km ($P \geq 3.2$ GPa) with transport to
506 UHP conditions with little concomitant heating ($T=600^\circ\text{C}$), water dominates over alkali
507 aluminosilicate solutes in released fluids (Fig. 12; FI in UHP rocks from Alps). At higher P
508 and sub-magmatic temperatures ($P = 3.4$ - 4 GPa; $T = 700$ - 800°C), approaching the hydrous
509 *solidus* for sedimentary and granitic rocks, aqueous fluids record an increase of solubility of
510 Si, Al, and alkalis by at least a factor of ten, reaching about 50-60 wt. % solutes (Fig. 12; MSI
511 in UHP rocks from Alps and Dabie Shan-Sulu). Fluids released at extremely high P and T
512 (about 5-8 GPa, and 900°C ; Kokchetav Massif; Dobrzhinetskaya et al. 2003a; Hwang et al.
513 2006; Korsakov and Hermann 2006; Hermann et al. 2013) consist mostly of aluminosilicate
514 components, appearing as hydrous silicate melts. This last case, however, applies to very deep
515 subduction, probably not directly reflecting on arc magma generation.

516 Enhanced silicate solubility approaching critical “mixing” of silicate-aqueous fluids is
517 due to solute structure change from (hydrated) ions to monomers and higher polymers
518 (Newton and Manning 2002; 2003; Hunt and Manning 2012). Silica polymerization was
519 recognized in aqueous fluids by *in-situ* spectroscopic studies both from experiments at high

520 pressures (Zotov and Keppler 2002), and in FI in UHP metamorphic rocks (Fig. 13; Lago di
521 Cignana metasediments). Although polymerized aqueous fluids can reach a state of hydration
522 which is similar to that in hydrous silicate melts, their physical properties are different,
523 perhaps more akin to that of colloidal silica (e.g., Iler 1979; Hunt and Manning 2012). In the
524 albite-H₂O system (Fig. 14), UHP fluids containing up to about 60 wt. % solutes show
525 considerably lower viscosities than hydrous melts, and form a linear trend towards pure water
526 values (Hack and Thompson 2011). Higher polymeric complexes, with characteristics similar
527 to hydrous melts, only form at much greater solute concentrations (well above 80 wt. %; Hack
528 and Thompson 2011).

529 Based on these observations, it appears not convenient to distinguish UHP aqueous
530 fluid phases in terms of various categories based on the solute (e.g., SiO₂) load (e.g.,
531 supercritical fluids or supercritical melts). Across the *P-T* range of most subduction zones,
532 silicate solubility can increase multifold in polymerized fluids, as well as the amount of
533 carbonate and sulfide/sulfate complexes, resulting in variable fluid densities and viscosities.
534 Polymerization at high pressure softens the transition from *sub-solidus* to *super-solidus*
535 conditions that governs the volatile behavior at shallower crustal conditions. These properties
536 directly reflect on the rate of mass transfer into the mantle wedge, and suggest that deep
537 aqueous fluids can represent effective metasomatic agents in the mantle wedge.

538

539 **Element recycling by deep subduction fluids**

540 *Trace elements*

541 To what extent the physico-chemical properties of UHP fluids, in particular ligand
542 species and concentration, influence elements fractionation between aqueous fluids and
543 matrix rocks during deep subduction directly reflects on metasomatism of the mantle wedge.

544 Element transfer by solute-poor aqueous fluids at depths of less than 90 km is generally
545 considered very limited, and gives way to transfer via hydrous melts or polymerized aqueous
546 fluids at sub-arc depths (see e.g., Manning 2004; Kawamoto 2006; Spandler et al. 2007;
547 Hermann and Spandler 2008; and references cited). At depths greater than about 90 km,
548 sedimentary, and oceanic rocks in the slab are expected to have different amounts of the same
549 minerals, and at T of about 750-800°C, they would initiate melting if water-saturated
550 (Schmidt et al. 2004). According to several authors (e.g., Johnson and Plank 1999; Rudnick et
551 al. 2000; Hermann and Rubatto 2009), incipient hydrous melting of slab rocks represents the
552 most important source of incompatible elements into the overlying mantle wedge. The LILE-
553 enriched, HFSE (Ti, Zr, Hf, Nb, Ta) depleted incompatible element pattern of arc-magma
554 would result from dissolution of mineral phases, such as phengite (LILE, and Be; e.g.,
555 Schmidt and Poli 1998, 2003; Hermann et al. 2006; Hermann and Spandler 2008), epidote
556 (REE, Sr, Th, Pb; e.g., Schmidt and Poli 1998, 2003; Hermann 2002a; Hermann et al. 2006),
557 lawsonite (REE, Sr, Pb, Th, U; e.g., Martin et al. 2011), allanite/monazite (La, Ce, Th; e.g.,
558 Hermann 2002b; Hermann and Rubatto 2009), leaving in the residue refractory zircon (Zr,
559 Hf), and rutile (Ti, Nb, Ta).

560 Kessel et al. (2005b) investigated the fluid/rock and melt/rock partition coefficients of
561 trace elements for water-saturated K-free synthetic basaltic systems from 4 to 6 GPa and from
562 700° to 1200°C. Their experimental data suggested that trace element fractionation resembles
563 subduction patterns in arc magmas, showing enrichments in LREE, LILE, Th, U, and Pb, and
564 depletion in HFSE. Differences in the partitioning behavior of trace elements diminish with
565 increasing depth and temperature, and at high P - T conditions, the trace element load of
566 aqueous fluids and hydrous melts becomes increasingly alike, and varies continuously with
567 temperature.

568 Data on the trace element concentrations of natural UHP fluids have been reported
569 only by a small number of examples, mainly because of the difficulties in analyzing MSI. In
570 these studies (Fig. 15), quantitative element concentrations in MSI were obtained by Laser
571 Ablation-Inductively Coupled Plasma-Mass Spectrometer (LA-ICP-MS) in rehomogenized
572 (for the method, see Malaspina et al. 2006), and in unheated inclusions (for the method, see
573 Halter et al. 2002), using Ca and/or Si as internal standards. Malaspina et al. (2006, 2009)
574 reported the first trace element data for MSI in metasomatic garnet meta-orthopyroxenite (P =
575 4 GPa, T = 750 °C) from the Maowu Complex in the Dabie Shan. UHP fluid trace element
576 patterns show fractionations similar to arc magmas, with LILE (Cs, Pb, Rb, Sr, Ba), LREE, U,
577 Pb, Th (high U/Th) enrichments, and HFSE (Ti, Zr, Hf, Nb, Ta) depletion (Fig. 15a). They
578 suggested that metasomatic fluids leading to important orthopyroxene and garnet
579 crystallization in peridotites were solute-rich aqueous fluids derived from surrounding crustal
580 lithologies.

581 In deeply-subducted crustal rocks, the geochemical signature of UHP fluids was
582 investigated by Ferrando et al. (2009) in the Dora-Maira whiteschists, formed by metasomatic
583 processes during prograde HP evolution. Magnesium-rich aluminosilicate aqueous fluids,
584 released during the growth of peak garnet (about 4 GPa and 750°C), are enriched in trace
585 elements compared to the whiteschists (Fig. 15b), and contain fractionated incompatible trace
586 element patterns with positive spikes for LILE, U, and Pb (up to about 100 to 1000 times the
587 mantle; Fig. 15b), and negative HFSE anomalies (Ti, Zr, Nb, Ta). Although some LILE reach
588 “granitic” melt concentrations, element fractionation in fluids does not strictly reflect either an
589 aqueous fluid (i.e., low Ba/Cs, Rb/Cs), or a silicate melt (low LREE, Sr, and Th enrichment
590 trend). Measured LILE enrichments, over LREE and HFSE, are in agreement with a main
591 contribution from the breakdown of phlogopite. Most of the released Rb, Ba and Th, however,
592 remain in phengite, which represents the stable mica at peak metamorphic conditions (Fig.

593 15b). Accessory monazite, Mg-dumortierite, zircon, and rutile in matrix whiteschists control
594 the HFSE and LREE rock budget (Fig. 15b).

595 Geochemical data like those of the two examples discussed above are extremely
596 valuable, since preserve an experimental flavor, reflecting the chemical control imposed by
597 the metamorphic evolution of rocks. There is an interesting parallel between experimentally-
598 derived fluid-mineral partitioning data which predict enhanced incompatible trace element
599 solubility in UHP fluids at, or close to supercritical conditions (Kessel et al. 2005b), and the
600 capacity of natural fluids to selectively extract LILE and LREE during prograde
601 metamorphism (from HP to UHP conditions). In deep subduction fluids, incompatible
602 element uptake is enhanced by mineral dissolution, but governed by metamorphic reactions.

603 Comparison between examples from the literature further allows us to evaluate the
604 extent of LILE/HFSE fractionation by deep slab fluids. Rutile is a common daughter phase in
605 many FI and MSI (Table 1), pointing to mobilization of HFSE in polymerized fluids.
606 Interestingly, relevant TiO₂ concentrations are observed also in dilute UHP aqueous fluids
607 formed at relatively low *T*, provided that they contain alkalis (cf., Table 1). The efficiency of
608 HFSE uptake appears mainly controlled by fluid properties at high *P*, and directly associated
609 with the matrix slab-rock chemical composition. A similar geochemical behavior is different
610 to that proposed for mass transfer by hydrous silicate melts, where the concentration of HFSE
611 has been recognized to be fundamentally a function of the degree of melting (cf., Hermann
612 and Rubatto 2009, and references cited): in other words, it is temperature dependent.
613 Accessory minerals controlling the HFSE budget of slab-rocks have been experimentally
614 found to be stable up to 850-900 °C.

615 Thus, alkali aluminosilicate fluids in subduction zones can mobilize and transport
616 HFSE, also at *sub-solidus* thermal conditions. This is in accordance with experimental and
617 thermodynamic research (e.g., Antignano and Manning 2008; Manning et al. 2008; Wilke et

618 al. 2012; Sanchez-Valle 2013), and with extensive field evidence (e.g., Selverstone et al.
619 1992; Rubatto and Hermann 2003; Garrido et al. 2005; Gao et al. 2007). However, HFSE
620 should be lost by UHP fluids during transport into the mantle wedge, before reaching the
621 hotter levels where subduction magmas are generated. According to Garrido et al. (2005)
622 chlorite-bearing peridotites overlying subducting slab could represent the potential HFSE
623 filters removing these elements from slab fluids, without changing their LILE budget. Only
624 after metasomatic re-equilibration, fluids would acquire the LILE-enriched and HFSE-
625 depleted trace element pattern characterizing arc magma.

626

627 *Carbon*

628 A major issue of subduction-zone geochemistry is to model and quantify the fluxes of
629 volatiles released into the mantle wedge, and returned to the exosphere via arc magmatism.
630 Carbon isotopic composition in volcanic gases of arc magmas requires that most of the Earth's
631 degassed CO₂ is recycled (e.g., Alt and Teagle 1999), suggesting that fluids released from the
632 subducting slab metasomatically enrich in carbon the mantle wedge. Modeling the fluxes of
633 carbon ingassing is, however, difficult, since release of carbon during subduction is modeled as a
634 discontinuous process.

635 In slab rocks, carbon is present in a variety of forms, including carbonate minerals and
636 organic matter, derived from altered oceanic crust and marine sediments. Part of the carbon
637 originally present in crustal lithologies is liberated as CO₂, during low-grade metamorphic
638 decarbonation reactions (e.g., Kerrick and Connolly 2001), or fixed as graphite (Foustoukos
639 2012; Galvez et al. 2013). At greater sub-arc depths, coupled thermal and thermodynamic
640 models strongly suggest that carbonate minerals do not react with silicate phases liberating CO₂,
641 and are stable along most subduction geothermal gradients (Connolly 2005; Gorman et al. 2006;

642 Molina and Poli 2008). In the absence of carbonate melting, most subducted carbon is predicted
643 to be transported deep in the mantle (Dasgupta et al. 2004; Poli et al. 2009).

644 Fluid inclusions in eclogites-facies rocks have been generally used to support the models,
645 as they contain little to no detectable CO₂, in contrast to fluid inclusions in many other
646 metamorphic environments (e.g., granulites; Touret 2001). Limiting the carbon inventory within
647 FI and MSI to the absence of CO₂, however, can result in an underestimation of the carbon
648 budget in UHP fluids. Carbonates have been reported in FI and MSI from several UHP terranes,
649 often associated with microdiamonds (cf., Table 1; and Dobrzhinetskaya 2012). For example,
650 Dobrzhinetskaya et al. (2007) identified H₂O molecules and carbonate ions in nanometric MSI in
651 Erzgebirge diamonds; Korsakov and Hermann (2006) described carbonate-rich MSI in diamond-
652 bearing calc-silicate rocks from Kokchetav massif. Crystallization from supercritical C-O-H
653 fluids represents the current preferred mechanism by most researchers to account for subduction
654 (i.e., metamorphic) diamond formation (cf., Dobrzhinetskaya 2012; and references cited).
655 Precipitation of diamond from C-O-H fluids containing carbonate ions, among other species,
656 implies that high f_{O_2} are prevailing in fluid phases (e.g., CCO buffer, Dobrzhinetskaya et al.
657 2001; de Corte et al. 2002). Yet, most diamond-bearing UHP metamorphic rocks record P - T
658 conditions (4-8 GPa, and 950-1000°C) that greatly exceed those of arc magma genesis, and
659 hence involve transport of carbon that cannot follow a simple return trajectory to the surface
660 through arc magmatism.

661 Spectroscopic study of carbon speciation in FI from oceanic-crust metasediments from
662 Lago di Cignana in western Alps shed light on the mechanisms of carbon liberation and transport
663 by deep subduction fluids at P - T conditions directly relevant to slab-mantle-arc transfer ($P \geq 3.2$
664 GPa; $T = 600^\circ\text{C}$). Raman analyses of C species in aqueous FI formed about 100 km depth,
665 revealed appreciable amounts of oxidized carbon dissolved as CO₃²⁻(aq) and HCO₃⁻(aq), along
666 with hydrous and hydrated carbonates and diamond, and no detectable CO₂ (Fig. 13b; Frezzotti

667 et al. 2011; 2014). These results strongly suggest that dissolution of carbonate minerals in
668 aqueous fluids as carbonate ions, in addition or alternatively to decarbonation, is a relevant
669 process for carbon transfer into the mantle wedge during subduction. Further, FI data indicate a
670 major role of water in the transport of carbon at P - T conditions that are directly relevant to slab-
671 mantle-arc transfer, with implications for the Earth's carbon cycle.

672 Recent geochemical, experimental, and theoretical research suggest similar trends for the
673 release and transport of oxidized carbon during deep subduction (cf., Manning et al. 2013; and
674 references cited). Ague and Nicolescu (2014) proposed that carbonate dissolution, accompanied
675 by silicate precipitation, represents a relevant mechanism for the release of oxidized carbon from
676 subduction zones, based on carbonate $\delta^{18}\text{O}$ and $\delta^{13}\text{C}$ systematics, combined with rock and fluid
677 inclusion studies in metamorphic rocks from the Cycladic complex. By computing the dielectric
678 constant of water at upper mantle P - T conditions, Pan et al. (2013) predicted the solubility of
679 carbonate minerals in the subducting lithosphere during dehydration reactions. Similar results
680 were obtained by spectroscopy experiments by Sanchez-Valle et al. (2013), which identified
681 carbonate ions as the dominant C-species in oxidized UHP aqueous fluids at P - T conditions
682 (650°C and 4 GPa) similar to those of Lago di Cignana rocks. The results summarized above
683 illustrate that investigations on the deep Earth's carbon cycle have entered a new phase where
684 observations in natural rock systems, thermodynamic models, and experimental research can be
685 merged. New research modifies the magnitude of carbon fluxes in the mantle in subduction
686 zones, and points to a strong need for a better quantification of the properties of carbon species
687 and minerals at depth.

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IMPLICATIONS

692

693 Present review shows that fluid inclusions (FI and MSI) are almost ubiquitous in eclogite-
694 facies crustal lithologies of both continental and oceanic origin. Their composition appears to be
695 controlled by the progressive devolatilization reactions occurring in rocks at *HP* and *UHP*
696 conditions, thus preserving a firsthand information on deep subduction fluids.

697 Water is the dominant volatile species in subduction-zone fluids, which contain variable
698 concentration of solutes, and of other volatiles. At *HP* conditions down to about 90 km depth
699 (i.e., fore-arc), moderate concentrations of chloride salts, alkalis, Si, and Al, \pm non-polar gases
700 are present in aqueous solutions, with properties similar to crustal fluids (e.g., halide ligands); at
701 greater mantle depths (i.e., sub-arc), water-rich phases contain gradually increasing amounts of
702 (alkali) aluminosilicate components (e.g., Si, Al, Ca, Fe, alkalis, Ti, Zr, $(\text{SO}_4)^{2-}$, $(\text{CO}_3)^{2-}$, and Cl).
703 Inclusion fluids record the solute structure change from (hydrated) ions to monomers and higher
704 polymers in aqueous fluids at increasing pressure during subduction.

705 Semi-quantitative chemical data on major and trace element enrichments in FI and MSI,
706 together with experimental and thermodynamic research, unambiguously attest for relevant and
707 selective element mobilization in *UHP* aqueous fluids, enhanced by mineral dissolution, and
708 ruled by the composition of matrix rocks. Increased element solubility and transport by the
709 aqueous phase is related to silica polymerization, as the pressure and temperature conditions
710 approach the critical endpoint.

711 It must be emphasized that research on fluid inclusions suggests that deep subduction
712 fluids have more complex chemical compositions than those of model fluids considered by
713 experiments. The finding of substantial amounts of dissolved oxidized carbon in aqueous fluids
714 at sub-arc depths implies a reconsideration of the petrological models supporting liberation of C

715 into the mantle wedge by metamorphic decarbonation reactions only, and to consider dissolution
716 and transport of carbon by aqueous fluids a relevant process in deep subduction zones.

717 Despite the amount of data so far obtained, open questions do remain. The chemical
718 behavior of other ligands, such as sulfur, is still very poorly constrained, although it is generally
719 acknowledged that slab-fluid transfer into the mantle wedge causes both oxidation and S
720 addition. Fluid inclusions could probably provide a most robust insight to the behavior of sulfur
721 species, by combining chemical and structural analyses. Analytical work should be undertaken to
722 reveal if the high solute load observed in some inclusions (for example MSI) is a primary
723 feature, or alternatively results in part from passive dehydration trends. A better quantification of
724 inclusions post-trapping chemical modifications, including element diffusion, and exchanges
725 with the host mineral, is also essential.

726 We believe that fluid inclusion research in subduction zones will provide exciting new
727 results in the coming years. Although hampered by many analytical difficulties, the study of
728 fluids in rocks provides the added value of revealing the continuity existing behind the apparent
729 extreme variations of the physico-chemical properties of aqueous fluids progressively released
730 during deep subduction.

731

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1225

1226 CAPTIONS TO FIGURES

- 1227 Figure 1: Photomicrographs of primary fluid inclusions (FI) in eclogite-facies HP-
1228 UHP rocks. (a) Aqueo-carbonic FI elongated parallel to the c-axis of the host OH-rich topaz.
1229 Liquid and gaseous CO₂ corresponds to about 30 % of the inclusion total volume. The two

1230 solids are a chloride (s) and gypsum. OH-rich topaz – kyanite quartzite from Sulu, China
1231 (sample RPC546, Plane Polarised Light, PPL). Modified from Frezzotti et al. (2007). (b)
1232 Diamond-bearing aqueous FI in garnet [liquid (L) + vapor (V) > 60% of the total volume of
1233 the inclusions]. Garnetite nodules from Lago di Cignana, Italian western Alps (sample C2BA,
1234 PPL). (c) Chloride-bearing aqueous FI in epidote. L + V represent about 80% of the inclusion
1235 total volume. Two isotropic salts (s) are also present. Phengite-kyanite-epidote eclogite from
1236 Sulu, China (sample RPC778, PPL). Modified from Ferrando et al. (2005a). (d) Chloride-
1237 bearing aqueous FIs elongated parallel to the c-axis of the host kyanite. L+V represent about
1238 40% of the inclusion total volume. Paragonite and ellenbergerite are the solid phases. Pyrope
1239 whiteschists from Dora-Maira, Italian western Alps (sample DM1598, PPL). Modified from
1240 Ferrando et al. (2009).

1241 Figure 2: Photomicrographs of preserved primary multiphase solid inclusions (MSI).
1242 (a) MSI in kyanite showing negative crystal shape. It contains the typical association of
1243 paragonite + muscovite + anhydrite + “alunite”-type sulfate + pyrite + small volume of an
1244 aqueous fluid. OH-rich topaz – kyanite quartzite from Sulu, China (sample RPC547, PPL). (b)
1245 MSI in garnet showing negative crystal shape. They are filled by paragonite + phlogopite +
1246 diamond + quartz + apatite + rutile without apparent fluid. Garnet gneiss, Saxonian
1247 Erzgebirge (sample RAR10, PPL). (c) Back-scattered electron (BSE) image of a freshly-
1248 broken surface of a pyrope including a MSI. Note the preferred orientation of Mg-chlorite and
1249 Na-K-phlogopite. Pyrope whiteschists from Dora-Maira, Italian western Alps (sample
1250 DM69). Modified from Ferrando et al. (2009). (d) MSI in garnet showing negative crystal
1251 shape. In this case, the oriented intergrowth described in Fig. 2c is recognizable also under the
1252 microscope. Pyrope whiteschists from Dora-Maira, Italian western Alps (sample DM1035)

1253 Figure 3: Photomicrographs (PPL: a, d; crossed polarizer: b, e) and back-scattered
1254 images (c, f) of two preserved primary MSI in pyrope from a Dora-Maira whiteschist. Both

1255 MSI have negative crystal shapes and are filled by of Mg-chlorite, Na-K-phlogopite, apatite
1256 and pyrite showing periodic crystallization, without any visible fluid phase. Pyrope
1257 whiteschists from Dora-Maira, Italian western Alps (a, b, and c: MSI in sample DM1035; d, e,
1258 and f: MSI in sample DM1042). Modified from Frezzotti and Ferrando (2007).

1259 Figure 4: Photomicrographs (PPL) of a primary melt inclusion in quartz from Mt.
1260 Genis granite (Sardinia, Italy). (a) At magmatic conditions (900°C), the inclusion consists of
1261 immiscible silicate-melt and hypersaline fluid. (b) At 20°C, the melt phase is crystallized in a
1262 fine-grained aggregate of randomly oriented anhedral to subhedral crystals, coexisting with
1263 chlorides, and liquid H₂O. Modified from Frezzotti (2001).

1264 Figure 5: Photomicrographs showing petrographic features of MSI. (a) Primary MSI in
1265 kyanite. Inclusions have the same dimensions and are oriented along the c axis of the host
1266 mineral. OH-rich topaz – kyanite quartzite from Sulu, China (sample RPC 547, PPL).
1267 Modified from Frezzotti and Ferrando (2007). (b) Distribution of primary MSI marking the
1268 growth zones of the host garnet. Pyrope whiteschists from Dora-Maira, Italian western Alps
1269 (sample DM1616, PPL). (c) BSE image of a MSI in pyrope, containing a large incidentally-
1270 trapped rutile. Pyrope whiteschists from Dora-Maira, Italian western Alps (sample DM69).
1271 Modified from Ferrando et al. (2009).

1272 Figure 6: Collages of BSE images of MSI on an EMP Ca-distribution map of the host
1273 garnet (courtesy of B. Stöckhert), showing abundance and the primary distribution of the
1274 MSI. The MSI are filled by silicates (quartz + white micas, locally retrogressed) + apatite +
1275 rutile ± diamond/graphite. Some inclusions still preserve a negative crystal shape. Other
1276 inclusions show reentrant angles at grain edges between inclusion minerals and host garnet
1277 and development of offshoots from the corners. Other inclusions show a star-shaped contour
1278 with short and thin microfractures lining from the corners. The distribution of the inclusion

1279 types is not systematic. Quartzofeldspathic rock, Saxonian Erzgebirge (modified from
1280 Stöckhert et al. 2009).

1281 Figure 7: Outline of the most important post-trapping modifications that can occur in a
1282 MSI during exhumation. Stage 1: Fluid trapping. Stage 2: Periodic crystallization of daughter
1283 minerals and dissolution and precipitation of the host mineral on the inclusion walls. Stage 3:
1284 Early decrepitation event, occurring at the UHP-HP transition, that can induce i) formation of
1285 short and large fractures (offshoots) departing from the inclusion corners, ii) crystallization of
1286 step-daughter minerals and partial re-crystallization of the host mineral producing irregular
1287 contours and re-entrant angles in the MSI, and/or iii) diffusion of water into the host mineral
1288 (see, Fig. 10). Stage 4: Daughter mineral crystallization leaves a residual fluid that may
1289 produce retrograde metamorphic reactions inside MSI. Stage 5: Late decrepitation event at
1290 LP, characterized by the formation of microfractures containing trails of small fluid
1291 inclusions, haloes of small fluid inclusions distributed around the inclusion, and/or star-
1292 shaped contours in MSI. The BSE image shows a MSI that experienced most of the described
1293 post-trapping modifications (see, Fig. 8a, and its caption).

1294 Figure 8: Photomicrographs showing post-trapping modifications in MSI. (a) BSE
1295 image of a decrepitated MSI in kyanite, showing minerals and cavities (black). The inclusion
1296 is characterized by the typical oriented association of daughter-minerals. Post-crystallization
1297 hydration reactions occurred on corundum to form diaspore. Decrepitation features (offshoots
1298 and microfractures) are also recognizable. OH-rich topaz – kyanite quartzite from Sulu, China
1299 (sample RPC 547). Modified from Ferrando et al. (2005b). (b) Fan-shaped chemical re-
1300 equilibration of the host garnet around a MSI. Pyrope whiteschists from Dora-Maira, Italian
1301 western Alps (sample DM14, PPL). PPL (c) and BSE (d) images of a MSI in garnet that
1302 experienced different post-trapping modifications. A non-equilibrium growth of the host
1303 mineral on the inclusion wall produced irregular and re-entrant contours. Subsequent

1304 decrepitation produced microfractures lining from the corners and loss of fluid (FI) and Ca-
1305 enrichment in the host garnet. Ky-Phe-Ep eclogite from Sulu, China (sample RPC 778).
1306 Modified from Ferrando et al. (2005b). (e) MSI with irregular contours due to a non-
1307 equilibrium growth of the host mineral on the inclusion wall. Offshoots due to partial
1308 decrepitation are also present (thick arrows). OH-rich topaz – kyanite quartzite from Sulu,
1309 China (sample RPC547, PPL). (f) Aqueo-carbonic fluid inclusion in quartz. Irregular contours
1310 and reentrant angles are produced by the non-equilibrium growth of the host mineral on the
1311 inclusion wall. Alpine quartz vein (sample AV3, PPL).

1312 Figure 9: Photomicrographs showing post-trapping modifications in FI and MSI. (a)
1313 Haloes of very small fluid inclusions distributed around single decrepitated MSI within
1314 pyrope. Note the irregular habit and the presence of some offshoots from the corners (thick
1315 arrows). Pyrope whiteschists from Dora-Maira, Italian western Alps (sample DM1598. PPL).
1316 Modified from Frezzotti and Ferrando (2007); (b) Typical example of post-trapping
1317 modifications in fluid inclusions from mantle clinopyroxene. The presence of offshoots (thick
1318 arrows) and of haloes of very small fluid inclusions is evident. Mantle xenolith from Mt. Iblei
1319 (Italy, PPL). (c) Post-entrapment re-equilibrations of a MSI in garnet. The large fracture
1320 forms a sort of offshoot and a halo of tiny fluid inclusions is also present. UHP eclogite from
1321 Western Gneiss Region (sample V2297, PPL). (d) Primary decrepitated MSI in pyrope
1322 showing irregular, star-shaped, contours. Both offshoots (thick arrows) and thin
1323 microfractures (thin arrows) are present. Pyrope whiteschists from Dora-Maira, Italian
1324 western Alps (sample DM69, PPL). (e) Decrepitated diamond-rich MSI in garnet. Thin
1325 microfractures (thin arrows) line from the star-shaped habit. Garnet gneiss, Saxonian
1326 Erzgebirge (sample RAR11, PPL). (f) Post-entrapment modification in silicate-melt
1327 inclusions in a quartz phenocrysts. The inclusions have thin microfractures (colorless arrows)
1328 lining from the inclusion apexes. Swarms of tiny fluid inclusions mantle the microfractures

1329 and form grayish halos around the inclusion cavity. Microtonalite porphyry (Calabona,
1330 Sardinia, PPL). Modified from Frezzotti (2001).

1331 Figure 10: Synchrotron FT-IR maps revealing the distribution and the concentrations
1332 of water in a Dora-Maira pyrope including some primary MSI (sample DM14). (a) Mapped
1333 garnet area (PPL). (b) Qualitative distribution map of liquid H₂O within MSI inclusions,
1334 revealed by absorption bands in the 3000–3500 cm⁻¹ region. (c) Qualitative distribution map
1335 of OH⁻ (bands in the 3500–3800 cm⁻¹ region), showing the hydrogen diffusion from the
1336 inclusions into the garnet. (d) Absorbance FT-IR map in the 3000–3800 cm⁻¹ region and
1337 relative calculated water contents in garnet (ppm). (e) FT-IR spectrum showing the absorption
1338 bands for molecular H₂O, and for OH⁻.

1339 Figure 11: *P–T* diagram showing metamorphic evolution of UHP units discussed in
1340 the text. *P–T* path for the Lago di Cignana Unit is simplified from Groppo et al. (2009). *P–T*
1341 path for the Dora-Maria (Brossasco-Isasca Unit) is inferred combining data from different
1342 lithologies (whiteschist, marble, eclogite, calc-silicate rocks: Rubatto and Hermann 2001;
1343 Hermann 2003; Castelli et al. 2007; Ferrando et al. 2009). *P–T* path for quartzite and eclogite
1344 from South Sulu is simplified from Ferrando et al. (2005a) and Frezzotti et al. (2007). *P–T*
1345 path for the high-*T* (dashed line) UHP zones of Dabie-Sulu is simplified from Zheng et al.
1346 (2011). The colored squares on the *P–T* paths represent the trapping conditions of the fluid
1347 inclusions reported in Table 1. The compilation of melting reactions in mafic rocks (thick
1348 gray lines) and pelites (dashed gray lines) is revised after Schmidt and Poli (2003) and Zheng
1349 et al. (2011). The gray-shaded zones representing the *solidi* for H₂O-saturated crustal rocks
1350 and the phengite dehydration melting are from Bebout (2007; 2013). Mineral abbreviation
1351 after Whitney and Evans (2010) and Kretz (1983).

1352 Figure 12: *P–T* diagram, modified from Hack and Thompson (2011), showing the
1353 silica solubility isopleths (in wt. % SiO₂) in the H₂O-SiO₂ system, the prograde subduction-

1354 zone gradients of 5°C/km (cold subduction) and 10°C/km (typical subduction), and the P-T
1355 paths of metamorphic rocks in Fig. 11. C is the upper critical endpoint for the H₂O-SiO₂
1356 system.

1357 Fig. 13: Raman spectra from aqueous fluid inclusions associated with diamonds from
1358 Lago di Cignana metasediments. Modified from Frezzotti et al. (2011). a) Silica in solution
1359 within a fluid inclusion. The 773 cm⁻¹ peak corresponds to the Si–O symmetric stretch of the
1360 Si(OH)₄ monomer, and the peak at 1,017 cm⁻¹ to the deprotonated monomers (SiO(OH)₃⁻,
1361 SiO₂(OH)₂²⁻, and so on). The aqueous fluid also contains SO₄²⁻, with Mg-calcite (MgCc) and
1362 quartz (Qtz). b) HCO₃⁻, and CO₃²⁻ in solution within a fluid inclusion, containing Qtz.
1363 Asterisks indicate host-garnet Raman bands.

1364 Fig. 14: Plot of viscosity data in the NaAlSi₃O₈ + H₂O system: measured values are
1365 compared with those obtained from models for fluid and melt (modified from Hack and
1366 Thompson 2011).

1367 Fig. 15: *In-situ* LA-ICP-MS analyses on multiphase solid inclusions. Normalizing
1368 values to the primordial mantle after McDonough and Sun (1995). (a) Trace-element
1369 concentrations in inclusions + host garnet compared with those in host garnet from mantle
1370 wedge garnet peridotites from Sulu (modified from Malaspina et al 2009; courtesy of N.
1371 Malaspina). (b) Trace-element concentrations in inclusions + host peak pyrope compared with
1372 those in host peak pyrope from continental UHP whiteschists from Dora-Maira. The trace-
1373 element patterns of the hosting metagranite, and of peak phengite, and prograde ellenbergerite
1374 and dumortierite are also reported. Modified from Ferrando et al. (2009).

Table 1

Table 1. Collection of published data on deep fluid inclusions

Metamorphic Belt	Lithology	OCCURRENCES				SOLID PHASES										FLUID PHASES			NATURE OF FLUID PHASE
		Peak P-T conditions	Stage of trapping ^a	Host mineral ^b	Name of the inclusions	silicates (hydrous)	silicates	carbonates	sulphates	sulphides	phosphates	oxides	chlorides	C phases	glass	water	gas	empty cavities	
<i>Oceanic crust</i>																			
Piemonte Zone (Monviso)	eclogitic vein ¹	2.6 GPa 550°C ²	near peak	HP Omp	fluid inclusions		albite sphene	calcite dolomite	anhydrite gypsum barite	pyrite	monazite	baddeleyite rutile Fe-oxides	halite sylvite			yes			complex aqueous brines with Na, K, Ca, Mg, Fe, Si, Al, Zr, Ti, P, Ba, Ce, La, S, CO ₃ ²⁻
Piemonte Zone (Lago di Cignana)	Mn-nodules ³	≥ 3.2 GPa 600°C ⁴	peak	UHP Grt	multiphase solid inclusions	paragonite	quartz	calcite Mg-calcite rhodochrosite dawsonite dypingite	anhydrite pentahydrate					diamond carbonaceous material		with SO ₄ ²⁻ , HCO ₃ ⁻ , CO ₃ ²⁻ , silica monomers		aqueous fluid enriched in Ca-Mg-Mn-Fe carbonate components ± Si ± Al ± Ti ± Na ± K ± S	
Tauern Window	eclogitic segregation ⁵	2.0 GPa 625°C ⁵	peak	HP Omp, Ep, Ap and Dol	fluid inclusions	white mica zoisite glaucofane	kyanite omphacite	carbonates	Ca-sulphates		apatite	Fe-Ti oxides rutile	chlorides salt-hydrated species		yes			brines with CO ₂ ± N ₂	
Tso Morari	eclogite ⁶	ca. 3.9 GPa ca. 750°C ⁶	prograde	Qtz	high-salinity brines								halite		yes			high-salinity brines	
	eclogite ⁶	ca. 3.9 GPa ca. 750°C ⁶	peak	Qtz, UHP Omp	gas-rich inclusions												N ₂ CH ₄	H ₂ -CH ₄ fluids	
Dabie-Sulu	eclogite ⁷	1.9 GPa 600-730°C ⁷	peak	HP Ep	aqueous inclusions										yes			highly saline brine-dominated fluids	
<i>Continental crust</i>																			
Western Gneiss Region	eclogite ^{8,9}	1.5-1.7 GPa 600°C ^{8,9}	peak	HP Omp and Grt	multiphase brine fluid inclusions	amphibole epidote	K-feldspar quartz albite	calcite	gypsum	sphalerite pentlandite chalcopyrite galena opaque minerals		rutile ilmenite	halite Pb-Br-chlorides		brine			Na-K-Ca-brines with Br, Pb, I, F, Li, SO ₄	
	Prp-quartzite ^{10,11}	4.0-4.3 GPa 730°C ¹²	prograde	HP-UHP Grt	fluid-free inclusions	talc Mg-chlorite		magnesite				Mg-phosphate	opaque minerals	chloride				aqueo-carbonic brines with P, Mg, Na, K	
Dora-Maira	Prp-quartzite ^{10,11}	4.0-4.3 GPa 730°C ¹²	retrograde	palizade Qtz	salt-bearing inclusions							pyrite	halite sylvite		yes			immiscibility between saline and carbonic fluids	
	Prp-quartzite ^{10,11}	4.0-4.3 GPa 730°C ¹²	retrograde	palizade Qtz	CO ₂ -rich inclusions										locally yes	CO ₂		immiscibility between saline and carbonic fluids	
	Prp-whiteschist ¹³	4.0-4.3 GPa 730°C ¹²	prograde	HP-UHP Ky	fluid inclusions	paragonite ellenbergerite								salts	yes			NaCl-MgCl ₂ -rich brines with Si and Al	
Rhodope	Prp-whiteschist ¹³	4.0-4.3 GPa 730°C ¹²	peak	UHP Grt	multiphase-solid inclusions	Mg-chlorite Na-K-phlogopite talc		magnesite		Zn-pyrite	Cl-apatite		very rare		yes			intermediate Al-Si aqueous solutions with Mg, Fe, Na, K, Ca, P, Cl, S, CO ₃ ²⁻ , LILE, U, Th	
	Grt-Ky-Bt gneiss ¹⁴	ca 4.5 GPa ca 1000°C ¹⁴	prograde	UHP Grt	multiphase inclusions	muscovite	quartz	Mg-siderite calcite rhodochrosite						graphite		CO ₂		high density hydrous-carbonaceous melt rich in C, Fe, Mg, Si, Al, K	
	Grt-Ky-Bt gneiss ¹⁴	ca 4.5 GPa ca 1000°C ¹⁴	retrograde	UHP Grt	multiphase inclusions	biotite muscovite	quartz kyanite zircon				apatite	rutile						melt	

(a) all the inclusions considered in this work were trapped at UHP conditions or at the HP-UHP or UHP-HP transition

(b) mineral abbreviations after Kretz (1983) and Whitney & Evans (2010)

⁽¹⁾ Philippot & Selverstone (1991), ⁽²⁾ Angiboust et al. (2012), ⁽³⁾ Frezzotti et al. (2011), ⁽⁴⁾ Groppo et al. (2009),⁽⁵⁾ Selverstone et al. (1992), ⁽⁶⁾ Mukherjee & Sachan (2009), ⁽⁷⁾ Fu et al. (2002),⁽⁸⁾ Svensen et al. (1999), ⁽⁹⁾ Svensen et al. (2001), ⁽¹⁰⁾ Philippot (1993), ⁽¹¹⁾ Philippot et al. (1995),⁽¹²⁾ Castelli et al. (2007), ⁽¹³⁾ Ferrando et al. (2009), ⁽¹⁴⁾ Mposkos et al. (2009),

Table 1

Table 1. continue

Metamorphic Belt	Lithology	Peak P-T conditions	Trapping conditions ^a	Host mineral ^b	Name of the inclusions	silicates (hydrous)	silicates	carbonates	sulfates	sulfides	phosphates	oxides	chlorides	C phases	glass	water	gas	empty cavities	
<i>Continental crust</i>																			
Erzgebirge	quartz-feldspathic rocks ¹⁵	7-8 GPa 1100°C ¹⁶	peak	UHP Grt	multiphase inclusions	phlogopite paragonite	quartz zircon			ZnS	apatite	rutile		diamond				silicate melt	
	Grt-gneiss ^{17, 18}	7-8 GPa 1100°C ¹⁶	peak	UHP Grt	polyphase inclusions	phlogopite paragonite phengite	quartz				apatite	rutile		diamond graphite				supercritical COH fluid rich in K, Na, SiO ₂	
	quartzo-feldspathic rocks ¹⁹	7-8 GPa 1100°C ¹⁶	peak	Dia	nanometric inclusions		SiO ₂ K-feldspar kyanite				archerite			diamond as host mineral			yes	COH-rich multicomponent fluid rich in Si, Al, K, P, Ti	
	Grt-bearing gneiss ²⁰	7-8 GPa 1100°C ¹⁶	peak	Dia	nanometric inclusions									diamond as host mineral	P/K-rich silica glass				P/K-rich silica melt
	Grt-Phg-Qtz feldspathic gneiss ¹⁶	7-8 GPa 1100°C ¹⁶	peak	Zrn	nanometric inclusions		quartz K-feldspar	CaCO ₃ BaCO ₃				TiO ₂ Fe _x O _y	KCl	diamond	quenched matter	yes			silicic-hydrous-saline high-density fluid, i.e. a C-O-H fluid mixed with a hydrous-silicic fluid rich in Al, K and a hydrous-saline fluid rich in Cl, K, Na
	felsic gneiss ²¹	4.3 - 6.0 GPa ca. 1000°C ²²	peak	UHP Grt, Zrn	multiple-inclusion pockets	phengite phlogopite chlorite biotite	quartz K-feldspar albite titanite					apatite	rutile		diamond				COH-rich multicomponent supercritical fluid
	felsic gneiss ²¹	4.3 - 6.0 GPa ca. 1000°C ²²	peak	Dia	nanometric inclusions	phengite phlogopite chlorite biotite	SiO ₂ Fe-rich pyroxene zircon	MgCO ₃	BaSO ₂			apatite	Ti-oxide Cr-oxide Fe oxide Th-oxide		diamond as host mineral			yes	COH-rich multicomponent supercritical fluid
Grt-Cpx-Qtz rocks ²³	4.3 - 6.0 GPa ca. 1000°C ²²	peak	UHP Grt	multiphase inclusions	phlogopite chlorite paragonite		calcite			metal sulfides				diamond					
Grt-clinopyroxenite ²⁴	4.3 - 6.0 GPa ca. 1000°C ²²	peak	Dia	fluid inclusions			carbonates							diamond as host mineral		yes		C-O-H hydrous/carbonatite-like fluid	
Kokchetav	Grt-Cpx rocks ²⁵	4.3 - 6.0 GPa ca. 1000°C ²²	retrograde	UHP Grt and Cpx	multiphase inclusions	phlogopite phengite talc	kokchetavite K-feldspar quartz cristobalite titanite zircon	calcite							siliceous glass				K-rich melt
	Grt-Qtz-Px rocks ^{20, 26}	4.3 - 6.0 GPa ca. 1000°C ²²	peak	Dia	nanometre-size inclusions									diamond as host mineral	K-rich, Si-poor glass?	K-rich, Si-poor aqueous solution?			K-rich, Si-poor fluid or melt
	Dol-marble ^{20, 26}	4.3 - 6.0 GPa ca. 1000°C ²²	peak	Dia	nanometre-size inclusions		silicates	carbonate	sulphates	sulphides	apatite phosphates	oxides	chlorides	diamond as host mineral		brine	yes		SiO ₂ -poor, ultrapotassic COH fluid with P, Cl, S
	Grt-bearing gneiss ²⁰	4.3 - 6.0 GPa ca. 1000°C ²²	peak	Dia	nanometric inclusions									diamond as host mineral	P/K-rich silica glass	brine			P/K-rich silica melt
	Mg-Cc marble ²²	4.3 - 6.0 GPa ca. 1000°C ²²	peak	UHP Grt and K-Cpx	polycrystalline & monocrystalline inclusions			Mg-calcite						diamond					carbonate melt
	Mg-Cc marble ²²	4.3 - 6.0 GPa ca. 1000°C ²²	peak	UHP Grt, Cpx and Ttn	polyphase inclusions	biotite	K-feldspar Cpx garnet titanite												K-rich silicate melt
	Mg-Cc marble ²²	4.3 - 6.0 GPa ca. 1000°C ²²	peak	UHP Grt, Cpx and Ttn	polyphase inclusions	biotite zoisite	K-feldspar titanite	Mg-calcite											K-rich silicate melt
Dol marble ²²	4.3 - 6.0 GPa ca. 1000°C ²²	peak	UHP Grt	polyphase inclusions	Ti-phengite		Mg-calcite				allanite							carbonate melt	

^(a) all the inclusions considered in this work were trapped at UHP conditions or at the HP-UHP or UHP-HP transition

^(b) mineral abbreviations after Kretz (1983) and Whitney & Evans (2010)

⁽¹⁵⁾ Hwang et al. (2001), ⁽¹⁶⁾ Dobrzhinetskaya et al. (2012), ⁽¹⁷⁾ Stöckhert et al. (2001), ⁽¹⁸⁾ Stöckhert et al. (2009),

⁽¹⁹⁾ Dobrzhinetskaya et al. (2003b), ⁽²⁰⁾ Hwang et al. (2006), ⁽²¹⁾ Dobrzhinetskaya et al. (2003a), ⁽²²⁾ Korsakov and Hermann (2006)

⁽²³⁾ Hwang et al. (2003), ⁽²⁴⁾ De Corte et al. (1998), ⁽²⁵⁾ Hwang et al. (2004), ⁽²⁶⁾ Hwang et al. (2005)

Table 1

Table 1. continue

Metamorphic Belt	Lithology	Peak P-T conditions	Trapping conditions ^a	Host mineral ^b	Name of the inclusions	silicates (hydrous)	silicates	carbonates	sulphates	sulphides	phosphates	oxides	chlorides	C phases	glass	water	gas	empty cavities		
<i>Continental crust</i>																				
	eclogite ²⁷	2.8 - 3.3 GPa 750-800°C ²⁷	retrograde	UHP Grt	multiphase solid inclusions	epidote	k-feldspar quartz plagioclase		barite									immiscibility between aqueous fluid and hydrous silicate melt		
	eclogite ²⁸	3.5 - 4.0 GPa 650-700°C ²⁸	retrograde	UHP Grt	polyphase inclusions	amphibole muscovite phengite	k-feldspar quartz plagioclase											yes	subsolidus breakdown products of phengite	
	eclogite ²⁸	3.5 - 4.0 GPa 650-700°C ²⁸	retrograde	UHP Grt	polyphase inclusions		k-feldspar quartz											yes	hydrous felsic melt produced during incongruent phengite melting	
	Coe-bearing eclogite ²⁹	3.5 - 4.0 GPa 650-700°C ²⁸	peak	UHP Omph and Ky	high-salinity aqueous inclusions								halite			yes		NaCl-dominated solutions		
	Coe-eclogite ³⁰	3.5 - 4.0 GPa 650-700°C ²⁸	prograde-peak	Qtz in UHP Omph, UHP Ep and Grt	gaseous inclusions												N ₂	N ₂ -fluids		
	Coe-eclogite ³⁰	3.5 - 4.0 GPa 650-700°C ²⁸	prograde-peak	Qtz in UHP Omph, UHP Ep and Grt	high-salinity aqueous inclusions								halite			yes		high-salinity aqueous fluids		
	Coe-eclogite ³¹	3.5 - 4.0 GPa 650-700°C ²⁸	peak	Qtz, UHP Omph and Ky	high-salinity NaCl-dominated aqueous inclusions											yes		NaCl-dominated fluids		
Dabie-Sulu	HP and UHP eclogite and Grt-clinopyroxenite ³²	≥ 2.7 GPa 630-890°C ³²	prograde-peak	UHP Grt, Qtz in UHP Grt	N ₂ -rich inclusions												H ₂ ± CO ₂	N ₂ -rich fluids		
	HP and UHP eclogite and Grt-clinopyroxenite ³²	≥ 2.7 GPa 630-890°C ³²	prograde-peak	UHP Grt, Qtz and Ap in UHP Grt	high-salinity brine inclusions		unidentified						halite			yes		high-salinity brines		
	UHP eclogite ³²	3.0 - 4.5 GPa 700-850°C ³³	prograde-peak	Qtz in UHP Grt	CH ₄ -rich inclusions													CH ₄ ± N ₂	CH ₄ -rich fluids	
	eclogite ³⁴	3.0 - 4.5 GPa 700-850°C ³³	retrograde	UHP Grt and Omph	polyphase inclusions		k-feldspar quartz albite												hydrous Na-K-Al-Si melt produced during dehydration melting of mica	
	OH-rich Toz - Ky quartzite ^{35, 36}	3.0 - 4.5 GPa 700-850°C ³³	peak	UHP Ky	multiphase-solid inclusions	paragonite muscovite chlorite	SiO ₂ zircon	calcite	anhydrite alunite-type sulphate barite	pyrite		corundum							yes	supercritical silicate-rich aqueous fluid / intermediate alkali-alumino-silicate aqueous solutions
	OH-rich Toz - Ky quartzite ^{35, 36}	3.0 - 4.5 GPa 700-850°C ³³	retrograde	HP Toz	aqueo-carbonic fluid inclusions				gypsum anhydrite				halite			yes	CO ₂		CaCl ₂ -rich brines	
	Ky-Phe-Ep eclogite ^{35, 37}	3.0 - 4.5 GPa 700-850°C ³³	peak	UHP Grt and Ky	multiphase-solid inclusions	paragonite amphibole Zn-staurolite	plagioclase		alunite-type sulphate	pyrite	apatite	rutile magnetite Zn-Mg-Al-Fe-Ti spinel							yes	supercritical silicate-rich fluid aqueous
	Ky-Phe-Ep eclogite ^{35, 37}	3.0 - 4.5 GPa 700-850°C ³³	retrograde	UHP Ep and Ky	bi-phase aqueous inclusions			Mg-calcite								yes			Na-Ca brine	
Ky-Phe-Ep eclogite ^{35, 37}	3.0 - 4.5 GPa 700-850°C ³³	retrograde	HP Ep, Qtz	three-phase aqueous inclusions								salts			yes			Na-brine		

^(a) all the inclusions considered in this work were trapped at UHP conditions or at the HP-UHP or UHP-HP transition

^(b) mineral abbreviations after Kretz (1983) and Whitney & Evans (2010)

⁽²⁷⁾ Gao et al. (2012), ⁽²⁸⁾ Liu et al. (2013), ⁽²⁹⁾ Xiao et al. (2000), ⁽³⁰⁾ Fu et al. (2001), ⁽³¹⁾ Xiao et al. (2002), ⁽³²⁾ Fu et al. (2003a),

⁽³³⁾ Zhang et al. (2008), ⁽³⁴⁾ Zeng et al. (2009), ⁽³⁵⁾ Ferrando et al. (2005b), ⁽³⁶⁾ Frezzotti et al. (2007), ⁽³⁷⁾ Ferrando et al. (2005a)

Table 1

Table 1. continue

Metamorphic Belt	Lithology	Peak P-T conditions	Trapping conditions ^a	Host mineral ^b	Name of the inclusions	silicates (hydrous)	silicates	carbonates	sulphates	sulphides	phosphates	oxides	chlorides	C phases	glass	water	gas	empty cavities
<i>Continental crust</i>																		
Dabie-Sulu	Qtz vein in eclogite ³⁸	3.0 - 4.5 GPa 700-850°C ³³	retrograde	Qtz	gas-rich inclusions												N ₂ CH ₄	H ₂ -CH ₄ fluids
	eclogite and Ky-quartzite ^{39, 40}	3.0 - 4.5 GPa 700-850°C ³³	peak	UHP Grt, Omp and Ky	high-salinity H ₂ O inclusions	amphibole mica	zircon	carbonate		pyrite		opaque minerals	halite			brine		multicomponent brine
	eclogite and veins ³³	3.0 - 4.5 GPa 700-850°C ³³	peak	UHP Zo and Ky	multiphase solid inclusions	paragonite		carbonate	anhydrite	pyrite		corundum magnetite				yes	yes	silicate-rich supercritical fluid
	eclogite and veins ³³	3.0 - 4.5 GPa 700-850°C ³³	peak	UHP Aln and Zo	multi-solid fluid inclusions	paragonite muscovite	quartz	calcite		anhydrite celestite		apatite				yes		silicate-rich supercritical fluid
	eclogite ⁴¹	3.0 - 4.5 GPa 700-850°C ³³	peak	UHP Omp and Grt	multi-phase aqueous inclusions	amphibole mica		calcite					oxides	halite		yes		high-density brine
Greenland Caledonides	granulitized UHP eclogite ⁴²	3.0 - 4.5 GPa 700-850°C ³³	prograde-peak	UHP Grt	N ₂ ±H ₂ O inclusions								halite		brine	N ₂	H ₂ -fluids coeval with high-salinity brines	
Greenland Caledonides	metapelites ⁴³	3.6 GPa 970°C ⁴³	retrograde	UHP Grt	polyphase inclusions	phengite biotite	quartz kyanite K-feldspar plagioclase						rutile					melt from dehydration melting of phengite
<i>Mantle</i>																		
Betic Cordillera	Ol-Opx rocks ⁴⁴	1.5 - 2.2 GPa 640 - 750°C ⁴⁴	peak	HP Ol	fluid + mineral inclusions	chlorite	olivine				Cl-apatite	magnetite				yes		homogeneous (supercritical?) fluid loaded of dissolved components
Western Gneiss Region	Grt websterite ^{45, 46, 47}	3.9 - 4.3 GPa 850-900°C ⁴⁵	peak	UHP Spl, Opx, Cpx and Grt	multiphase solid inclusions	phlogopite	zircon kalisilite orthopyroxene	magnesite (BaCa)CO ₃ dolomite		Fe-Ni-sulphide Ba-sulphide	Cl-apatite monazite	Cr-spinel rutile periclase		diamond C-phase				COH- and silicate-rich supercritical fluid
	Grt websterite ⁴⁸	5.5 GPa 800°C ⁴⁸	peak	UHP Grt	polyphase solid inclusions	amphibole chlorite	Al-silicate			Fe-Pb-sulphides	Y-phosphate	Al-Fe-Mg oxides Al-spinel		diamond C-phase				supercritical, dense, H-C-N-O-F-P-S-Cl metal-bearing supercritical fluids
Moldanubian Zone	Phl-Ap-bearing Spl-Grt peridotite ⁴⁹	2.3 - 3.5 GPa 850 - 1030°C ⁴⁹	peak	UHP Cr-Spl	multiphase solid inclusions	phlogopite chlorite Mg-hornblende talc		dolomite calcite		Fe-Ni sulphide galena	apatite monazite	U-Th oxide		graphite				carbonatite melt or supercritical fluid
Dabie-Sulu	Grt pyroxenite ^{29, 50}	ca 2.0 GPa 680-860°C ⁵⁰	prograde-peak	UHP Grt	high-salinity aqueous inclusions ± N ₂								halite		yes	H ₂ ± CO ₂	N ₂ -bearing NaCl-rich solutions	
	Grt orthopyroxenite and Grt-websterite ^{51, 52}	3.9 - 4.1 GPa 700 - 800°C ⁵¹	peak	UHP Grt	polyphase inclusions	amphibole chlorite talc mica	albite sphene			sulfide	apatite	spinel						solute-rich aqueous fluids
	Grt-peridotite ⁵²	ca 4.2 GPa ca 760°C ⁵²	peak	UHP Grt	polyphase inclusions	amphibole chlorite talc mica				pyrite		spinel						solute-rich aqueous fluids

(a) all the inclusions considered in this work were trapped at UHP conditions or at the HP-UHP or UHP-HP transition

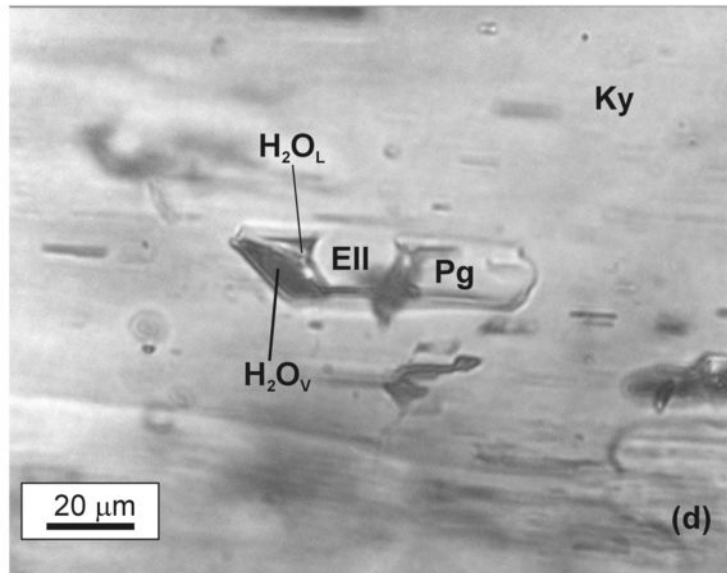
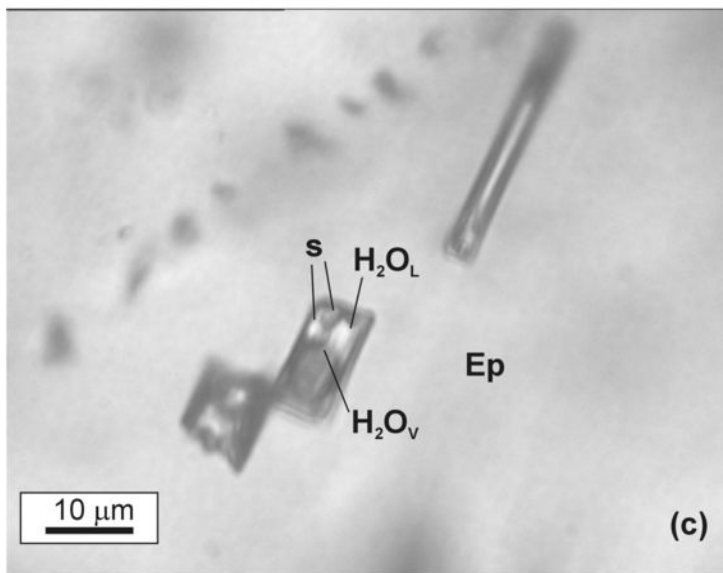
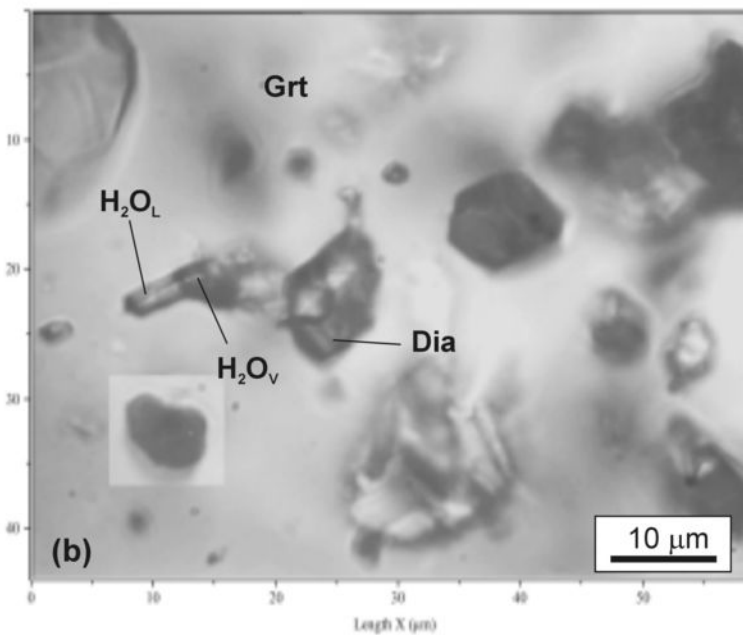
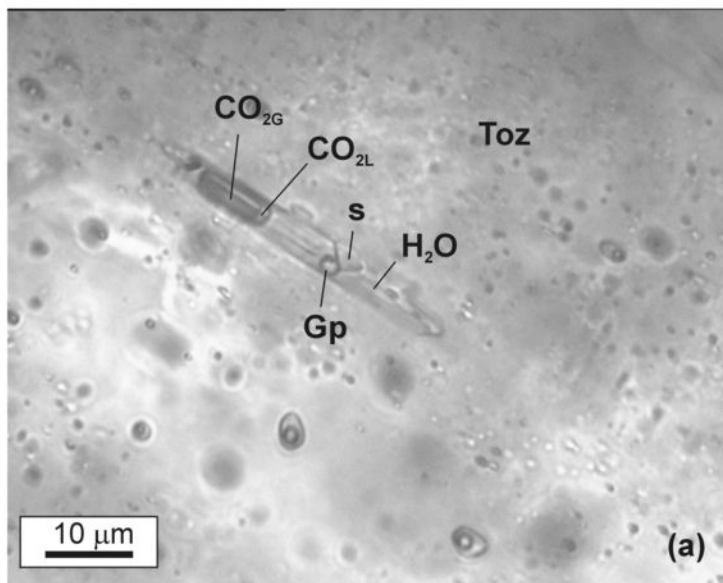
(b) mineral abbreviations after Kretz (1983) and Whitney & Evans (2010)

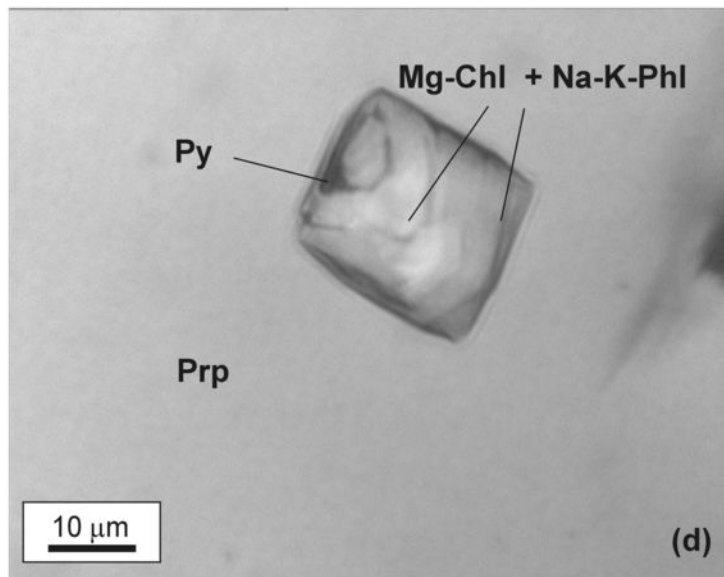
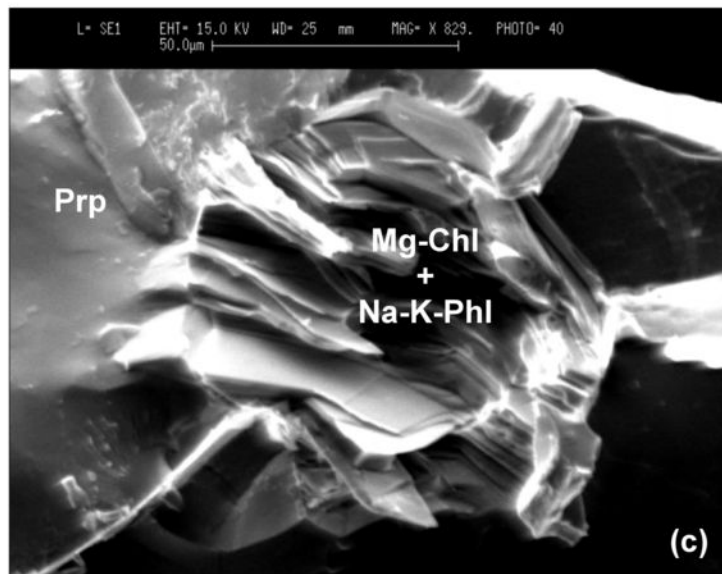
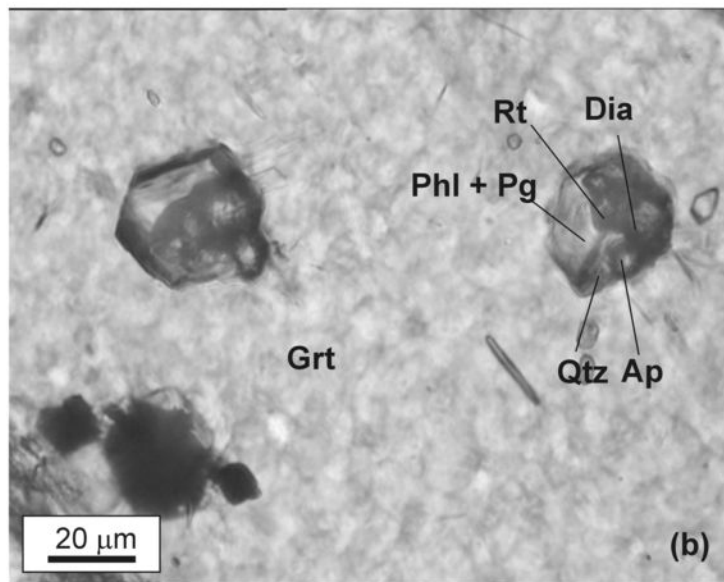
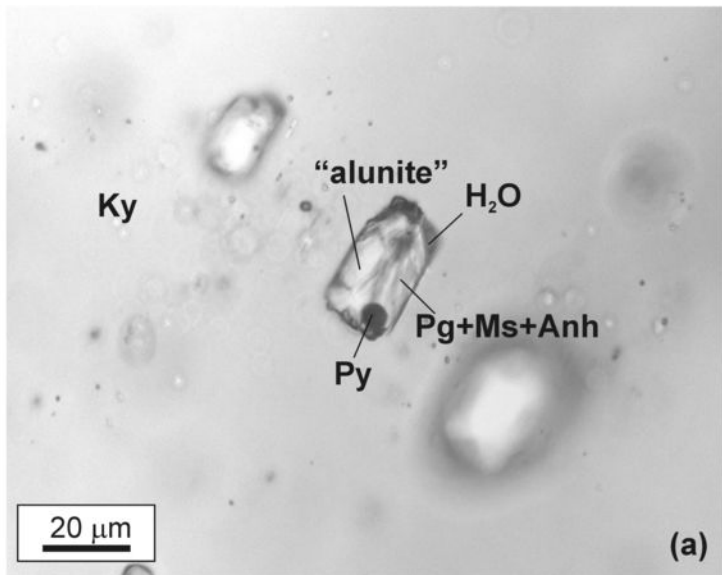
(38) Xu et al. (2006), (39) Zhang et al. (2005), (40) Xiao et al. (2006), (41) Shen et al. (2003), (42) Fu et al. (2003b),

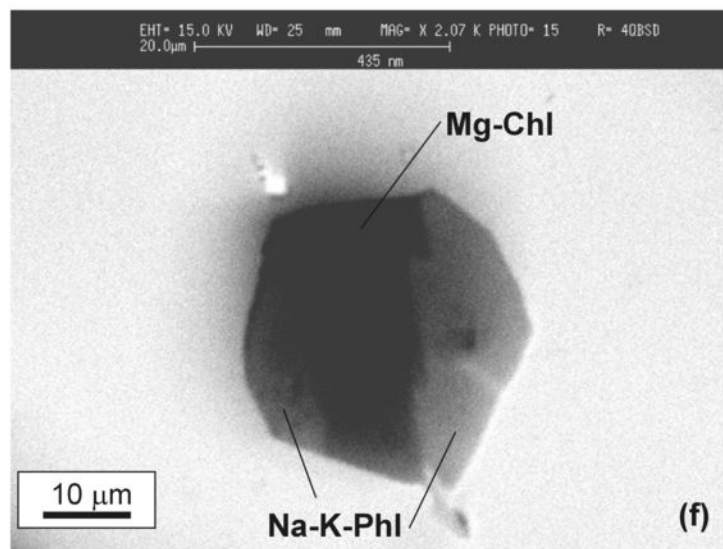
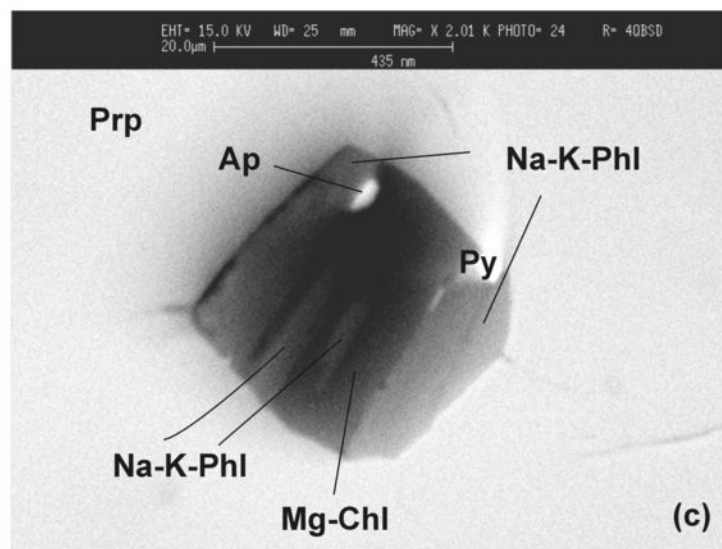
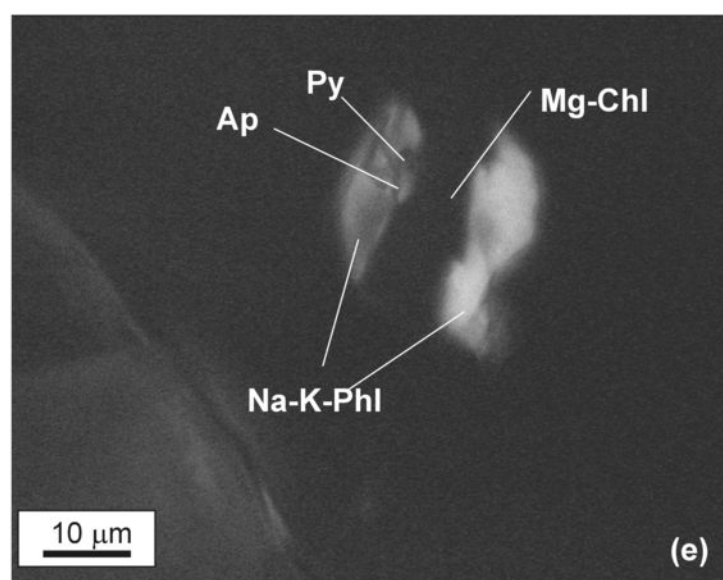
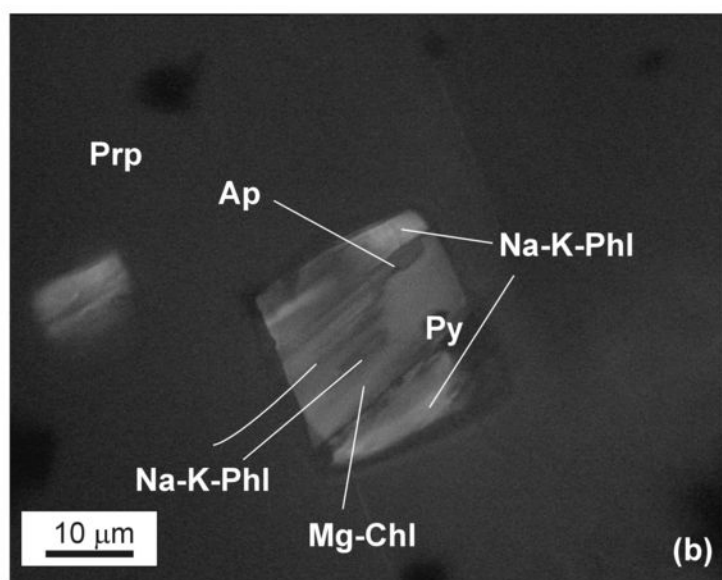
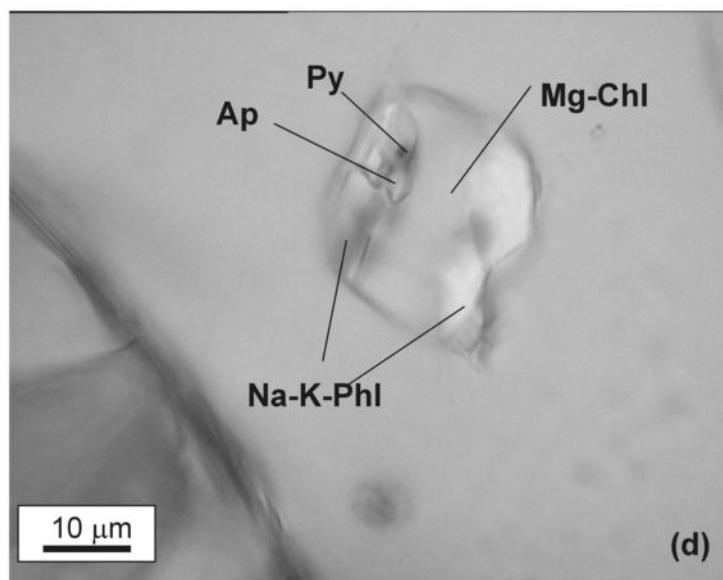
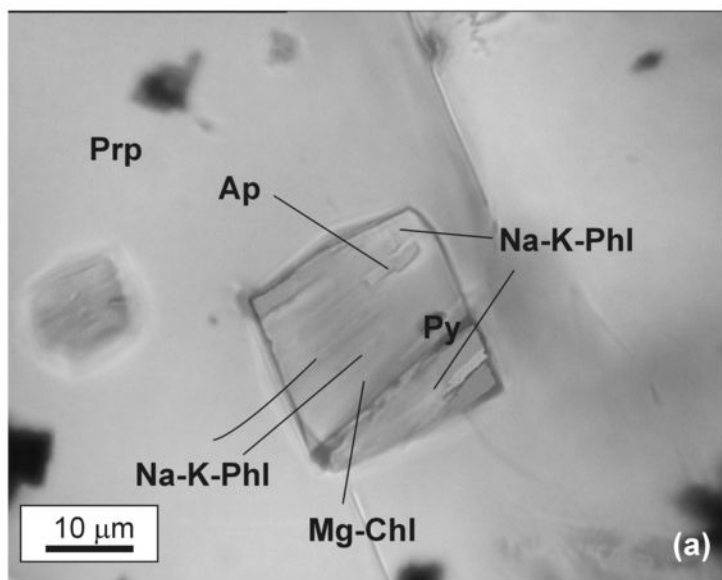
(43) Lang & Gilotti (2007), (44) Scambelluri et al. (2001), (45) Carswell and van Roermund (2005),

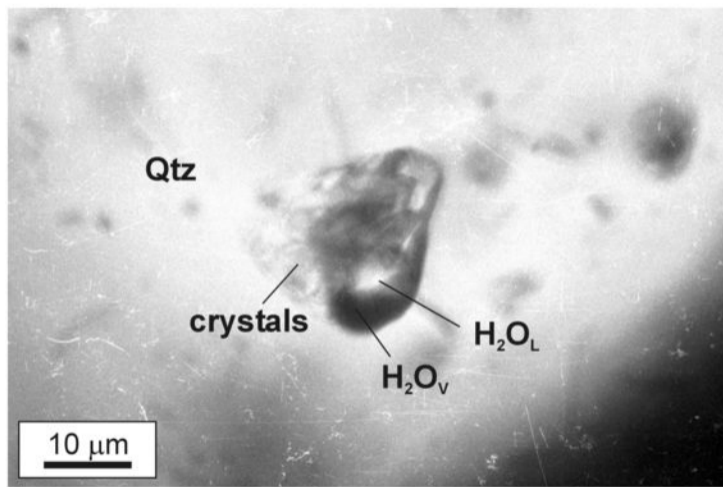
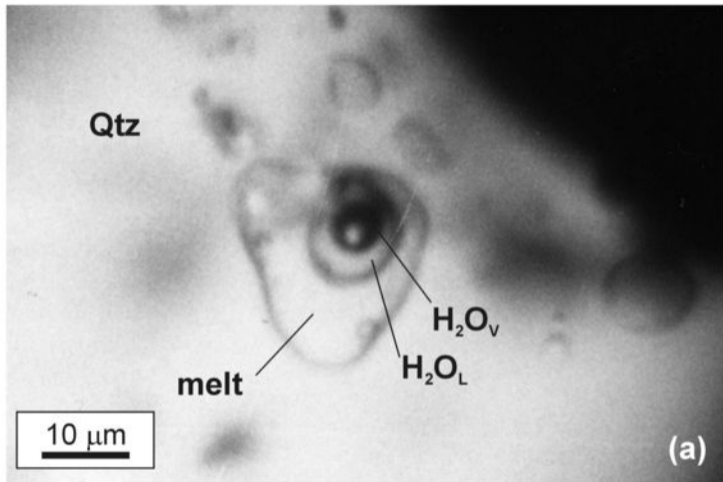
(46) van Roermund et al. (2002), (47) Malaspina et al. (2010), (48) Vrijmoed et al. (2008), (49) Naemura et al. (2009),

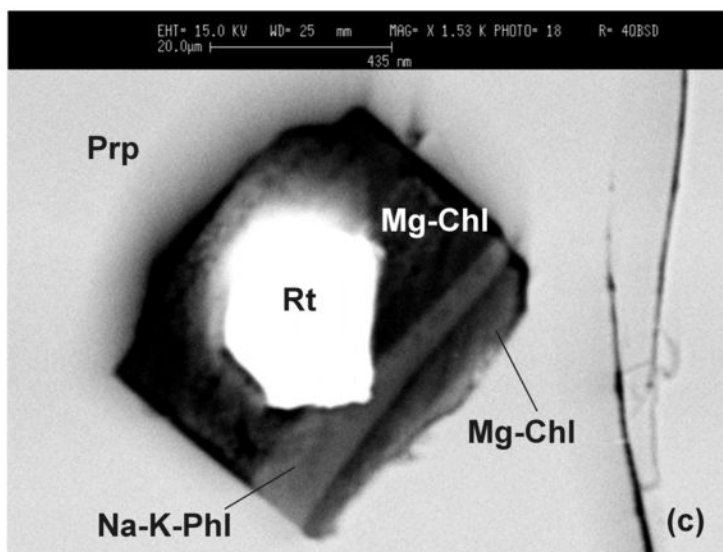
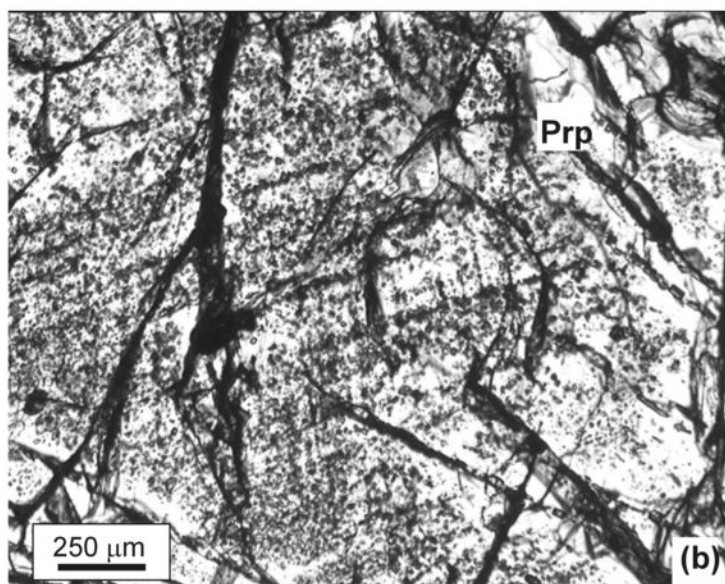
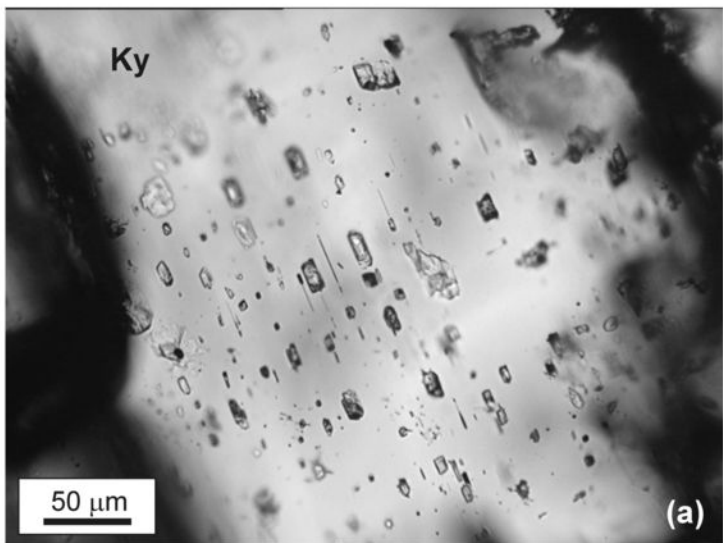
(50) Xiao et al. (2001), (51) Malaspina et al. (2006), (52) Malaspina et al. (2009)











St6103/4-L3

garnet

background: Ca-distribution map



