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ZnMoO₄: A promising bolometer for neutrinoless double beta decay searches

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ABSTRACT

We investigate the performances of two ZnMoO₄ scintillating crystals operated as bolometers, in view of a next generation experiment to search the neutrinoless double beta decay of ¹⁰⁰Mo. We present the results of the α vs β/γ discrimination, obtained through the scintillation light as well as through the study of the shape of the thermal signal alone. The separation obtained at the 2615 keV line of ²⁰⁸Tl is 8σ , using the heat-light scatter plot, while it exceeds 20σ using the shape of the thermal pulse alone. The achieved FWHM energy resolution ranges from 2.4 keV (at 238 keV) to 5.7 keV (at 2615 keV). The internal radioactive contaminations of the ZnMoO₄ crystals were evaluated through a 407 h background measurement. The obtained limit is <32 µBq/kg for ²²⁸Th and ²²⁶Ra. These values were used for a Monte Carlo simulation aimed at evaluating the achievable background level of a possible, future array of enriched Zn¹⁰⁰MoO₄ crystals.

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1. Introduction

Double beta decay (DBD) searches became of critical importance after the discovery of neutrino oscillations. Plenty of experiments are now in the construction phase and many others are in R&D phase [1–4]. The main challenges for all the different experimental techniques are the same [5]: (i) increase the active mass, (ii) decrease the background, and (iii) increase the energy resolution.

Thermal bolometers are ideal detectors for this survey: crystals can be grown with a variety of interesting DBD-emitters and multikg detectors can be operated with excellent energy resolution [6].

The Cuoricino experiment [7], an array of 62 TeO₂ crystal bolometers, demonstrated not only the power of this technique, but also that the main source of background for these detectors arises from surface contaminations of radioactive α -emitters. α particles, emitted from radioactive contaminations located on the surfaces of the detector or of passive elements facing them, can lose part of their energy in a few μ m and deposit the rest in the crystal bolometer. This produces an essentially flat background starting from the

* Corresponding author. *E-mail address:* Stefano.Pirro@mib.infn.it (S. Pirro). Q-value of the decays (several MeV) down to low energies, completely covering, therefore, the region of the $Q_{\beta\beta}$ values. Moreover simulations show that this contribution will largely dominate the expected background of the CUORE experiment [8,9] in the region of interest, since there is no possibility to separate this background from the two DBD electrons. The natural way to discriminate this background is to use a scintillating bolometer [10]. In such a device the simultaneous and independent readout of the heat and of the scintillation light permits to discriminate events due to β/γ , α and neutrons thanks to their different scintillation yield. Moreover, if the crystal is based on a DBD emitter whose transition energy exceeds the 2615 keV γ -line of ²⁰⁸Tl, the environmental background due to natural γ 's will decrease abruptly.

¹⁰⁰Mo is a very interesting ββ-isotope because of its large transition energy $Q_{\beta\beta}$ = 3034.4 keV [11] and a considerable natural isotopic abundance δ = 9.67% [12]. Several inorganic scintillators containing molybdenum were developed in the last years. ZnMoO₄ [13] is an example of crystal recently tested as a cryogenic detector giving very promising results [14].

In this work we present the results obtained with two ZnMoO₄ crystals of superior quality with respect to the sample previously studied. This work is focused on the recent observation [15] that



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the thermal signal induced by α and γ/β particles shows a slightly different time development. This feature seems to be explained [16] by the relatively long scintillation decay time (of the order of hundreds of μ s) observed in some scintillating crystals (e.g. molybdates). This long decay, combined with an high percentage of non-radiative de-excitation of the scintillation channel, will transfer phonons (i.e. heat) to the crystal. This extremely tiny, but measurable, time dependent phonon release has a different absolute value for isoenergetic α and β/γ particles due to their different scintillation yield. The possibility to have a bolometer in which the α background is identified without the need of an additional light detector is particularly appealing since it translates in a simplified set-up.

In Section 2 the experimental set-up is outlined; in Section 3 we describe the α vs β/γ discrimination capability. The evaluation of the internal radioactive contamination of our sample is presented in Section 4, while in Section 5 we present a detailed Monte Carlo simulation of the background performances of a possible future experiment based on a ZnMoO₄ crystal array.

Within the Lucifer Project [17], an array of enriched $Zn^{82}Se$ crystals operated as scintillating bolometers, there is also the option of having part of the detectors made of ¹⁰⁰Mo enriched crystals. The present work, though based on small size crystal samples, shows the performance that could be reached by a small scale DBD decay experiment based on $Zn^{100}MoO_4$ crystals.

2. Experimental details

High quality ZnMoO₄ crystals were developed in the Nikolaev Institute of Inorganic Chemistry (NIIC, Novosibirsk, Russia). Starting material for the crystal growth were High Purity ZnO (produced by Umicore) and MoO₃, synthesized by NIIC. Crystals up to 25 mm in diameter and 60 mm in length were grown by the low-thermalgradient Czochralski technique in a platinum crucible with a size of \emptyset 40 mm \times 100 mm [18].

In this paper we present the result of two separate runs, both carried out in an Oxford 200 ${}^{3}\text{He}/{}^{4}\text{He}$ dilution refrigerator located deep underground in the Laboratori Nazionali del Gran Sasso (average depth \approx 3650 m w.e. [19]).

The first run was dedicated to the study of the discrimination capability of $ZnMoO_4$ bolometers, while the second one was devoted to the evaluation of the internal contaminations of the sample.

In the first run two different ZnMoO₄ crystals were tested in order to study the reproducibility of the background rejection over different samples. The samples were a 27.5 g colourless cylinder (\emptyset 18.5 mm × 22.3 mm) and a 29.9 g slightly orange parallelepiped (28.5 × 18.4 × 13.2 mm³).

The ZnMoO₄ crystals were held by means of four S-shaped PTFE supports fixed to cylindrical Cu columns. They were surrounded (with no direct contact) by a plastic reflecting sheet (3M VM2002). The Light Detector (LD) [20] is constituted by a 36 mm diameter, 1 mm thick pure Ge crystal absorber. The Ge wafer is heated up by the absorbed photons and the temperature variation is proportional to the scintillation signal. The set-up of the detectors is schematized in Fig. 1.

The temperature sensor of the ZnMoO₄ crystals is a $3 \times 3 \times 1 \text{ mm}^3$ Neutron Transmutation Doped (NTD) germanium thermistor, the same used in the Cuoricino experiment. It is thermally coupled to the crystal via 9 glue spots of $\approx 0.6 \text{ mm}$ diameter and $\approx 50 \text{ µm}$ height. The temperature sensor of the LD has a smaller volume ($3 \times 1.5 \times 0.4 \text{ mm}^3$) in order to decrease its heat capacity, increasing therefore its thermal signal. A resistor of $\approx 300 \text{ k}\Omega$, realized with a heavily doped meander on a 3.5 mm³ silicon chip, is attached to each absorber and acts as a heater to stabilize the gain of



Fig. 1. Set-up of the detectors. The ball-bonded Au wires are crimped into "male" Cu tubes (pins) and inserted into ground-insulated "female" Cu tubes. Custom wires from detectors towards cryostat are not drawn. A section of the light detector is not drawn for a better understanding.

the bolometer [21,22]. The details of the electronics and the cryogenic facility can be found elsewhere [23–25].

The heat and light pulses, produced by a particle interacting in the absorber and transduced in a voltage pulse by the NTD thermistors, are amplified and fed into a 18 bit NI-6284 PXI ADC unit. The trigger is software generated on each bolometer and when it fires waveforms 0.6 s long, sampled at 2 kHz, are saved on disk. Moreover, when the trigger of a $ZnMoO_4$ crystal fires, the corresponding waveform from the LD is recorded, irrespective of its trigger.

As the main goal of the measurements was to test the discrimination capability of the detectors between α and β/γ events, a ²³⁸U/²³⁴U α source was faced to the crystals, on the opposite side with respect to the LD. The source was covered with a 6 µm thick Mylar film, in order to "smear" the α 's energies down to the $Q_{\beta\beta}$ energy region. The γ calibration of the ZnMoO₄ crystals is performed through a movable a ²³²Th source inserted between the Dewar housing the cryostat and the external lead shield. The energy calibration of the LD is achieved thanks to a ⁵⁵Fe X-ray source, producing two X-rays at 5.9 and 6.5 keV, faced to the LD (see Fig. 1).

In the second run only the best performing bolometer (the cylindrical crystal) was characterized. The α source was removed and the crystal was surrounded by three layers of 12 µm thick ultrapure polyethylene sheets in order to shield the bolometer against the surface contamination of the copper structure. In this measurement the lack of space prevented us from mounting the LD. However this was not a problem, as the previous run convincingly demonstrated that the pulse shape analysis can provide an extremely good α background rejection without the need for the light detection.

2.1. Data analysis

To maximize the signal to noise ratio, the pulse amplitude is estimated by means of an optimum filter technique [26,27]. The filter transfer function is built from the ideal signal shape s(t) and the noise power spectrum $N(\omega)$. s(t) is estimated by averaging a large number of triggered pulses (so that stochastic noise superimposed to each pulse averages to zero) while $N(\omega)$ is computed averaging the power spectra of randomly acquired waveforms not containing pulses. The amplitude of a signal is estimated as the maximum of the filtered pulse. This procedure is applied for the signal on the ZnMoO₄ bolometer. The amplitude of the light signal, instead, is estimated from the value of the filtered waveform at a fixed time delay with respect to the signal of the ZnMoO₄ bolometer, as described in detail in Ref. [28]. The detector performances are reported in Table 1. The baseline resolution, FWHM_{base}, is governed

Table 1

Technical details for the $ZnMoO_4$ bolometers (cylinder and parallelepiped) and for the LD. The cylindrical $ZnMoO_4$ was measured twice, so we reported the parameters also for the background run (Cylinder^{*}). R_{work} is the working resistance of the thermistors. Signal represents the absolute voltage drop across the termistor for a unitary energy deposition.

Crystal	R _{work} [MΩ]	Signal [µV/MeV]	FWHM _{base} [keV]	$\tau_R [\mathrm{ms}]$	$\tau_D [ms]$
Cylinder	3.7	140	0.6	17	50
Parallel.	4.7	320	1.2	8	33
LD	8.8	1700	0.16	4	11
Cylinder*	2.5	200	0.7	17	48

by the noise fluctuation at the filter output, and does not depend on the absolute pulse amplitude. The rise (τ_R) and decay times (τ_D) of the pulses are computed as the time difference between the 10% and the 90% of the leading edge, and the time difference between the 90% and 30% of the trailing edge, respectively.

After the application of the optimum filter, signal amplitudes are corrected for temperature and gain instabilities of the set-up thanks to a monochromatic power injection in the Si heater taking place every few minutes. The ZnMoO₄ is calibrated using the most intense γ -peaks from the ²³²Th source, while the LD is calibrated using the ⁵⁵Fe X-ray doublet.

The FWHM energy resolution obtained on the cylindrical (parallelepiped) crystal ranges from 2.5 \pm 0.1 (2.4 \pm 0.1) keV at 238 keV to 3.8 \pm 0.9 keV (7.6 \pm 1.3) at 2615 keV. The energy resolution on the 5407 keV α + recoil line (due to a weak internal contamination of ²¹⁰Po) can be evaluated only on the long background run for the cylindrical crystal and gives 5.3 \pm 1.1 keV. The FWHM energy resolution on the LD, evaluated on the ⁵⁵Fe X-ray doublet, is 321 \pm 9 eV. Experimental resolutions are worse than theoretical ones in agreement with the observed performance of macro-bolometers [29].

The light-to-heat energy ratio as a function of the heat energy is shown for the calibration spectrum in Fig. 2. β/γ and α decays give rise to very clear separate distributions. In the upper band, ascribed to β/γ events, the 2615 keV ²⁰⁸Tl γ -line is well visible. The lower band, populated by α decays, shows the continuum rate induced by the degraded α source as well as the ²¹⁰Po doublet.

The Light Yield (LY), defined as the ratio between the measured light (in keV) and the nominal energy of the event (in MeV), was measured for the most intense γ -lines giving 1.10 ± 0.03 keV/MeV



Fig. 2. The light-to-heat energy ratio as a function of the heat energy obtained with the cylindrical crystal in the first run, during a 62 h ²³²Th calibration. The upper band (ascribed to β/γ events) and lower band (populated by α decays) are clearly separated. The 2615 keV ²⁰⁸Tl γ -line is well visible in the β/γ band as well as a the continuum rate induced by the degraded α source and the 5407 keV ²¹⁰Po doublet in the α band. The discrimination power is reported in Section 3.



Fig. 3. Comparison of the β/γ and α pulse shape. Top figure: thermal pulse obtained by averaging events in the 2615 keV γ peak (blue line) and events belonging to the α band with an equivalent energy (red line): their difference is not appreciable. Bottom figure: percentage difference $s(t)_{\alpha} - s(t)_{\beta/\gamma}$. Differences at a level of a few per mille are visible both in the rise and decay of the thermal pulse. (For interpretation of the references to colour in this figure legend, the reader is referred to the web version of this article.)

and 0.78 ± 0.02 keV/MeV for the cylinder and for the parallelepiped, respectively. These values are constant from 0.2 to 2.6 MeV and are not corrected for the light collection efficiency. The LY of the cylindrical crystal is well in agreement with the one reported in [14,30], while the parallelepiped shows a lower LY.

The Quenching Factor (QF), defined as the ratio of the $LY_{\alpha}/LY_{\beta/\gamma}$ for the same energy release, was evaluated on the 5407 keV α -line and results 0.18 ± 0.02 for both crystals.

3. α vs β/γ discrimination

As reported in [14,15],Molybdate crystals can provide α vs β/γ discrimination by making use of the thermal information only. In Fig. 3 the ideal signal shape s(t) for the two event classes is shown together with the percentage difference $s(t)_{\alpha} - s(t)_{\beta/\gamma}$. Pulse shapes are obtained by averaging pulses (obtained in the same calibration measurement of Fig. 2) in the energy range 2610–2620 keV. In order to take into account possible (small) jitter times of the triggered events, the average pulse is computed aligning the pulses at their maximum.

Differences at a level of a few per mille are visible both in the rise and decay of the thermal pulse.¹

To fully exploit this feature several shape variables have been studied: τ_R , τ_D , TVL, TVR, χ^2_{OT} . The Test Value Left (TVL) and Right (TVR) are defined as:

$$\Gamma VL = \frac{1}{A \cdot w_l} \sqrt{\sum_{i=i_M - w_l}^{i_M} (y_i - As_l)^2};$$
(1)

$$\text{TVR} = \frac{1}{A \cdot w_r} \sqrt{\sum_{i=i_M}^{i_M + w_r} (y_i - As_i)^2}.$$
 (2)

¹ We will refer to pulses from the cylindrical bolometer throughout the rest of the text. However, the parallelepiped bolometer showed consistent results.



Fig. 4. Top Figure: TVR as a function of the energy. The upper band is populated by α particles (events in the 2300–3200 keV energy range are shown in red) while β/γ 's contribute to the lower band (events in the 2300–3200 keV energy range are shown in black). Bottom Figure: TVR histogram for the α sample in red and for the β/γ in black. The mean values and the standard deviations, as estimated from a Gaussian fit, are reported. Using Eq. 3 a theoretical discrimination power of \approx 20 is obtained. (For interpretation of the references to colour in this figure legend, the reader is referred to the web version of this article.)

All quantities in the above equations refer to the optimum filtered pulses: y_i is the pulse, A and i_M its amplitude and maximum position, s_i the ideal signal pulse scaled to unitary amplitude and aligned to y_i , $w_{l(r)}$ the left (right) width at half maximum of s_i . The maximum of the ideal pulse is aligned with the maximum of the filtered pulse, fractional delays are handled with a linear interpolation of the data samples.

Finally, χ^2_{OT} was computed as the sum of the squared fit residuals normalized to the noise RMS, using s_i as a fit function of the central part of the pulse. This algorithm is described in detail in [31].

Bolometers are nonlinear detectors [32], i.e. the shape of the signal slightly depends on the amount of released energy. This implies that also the shape parameters (evaluated with respect to an energy independent pulse, s_i) will show such a dependence. To remove such a dependency variables were linearized in the 2300–3200 keV energy range. This procedure, moreover, allowed to enlarge the data sample on which the pulse shape analysis was performed.

Fig. 4 shows the TVR variable as a function of the energy. As for the case of Fig. 2, β/γ and α events are distinctly separated. Events in the linearized energy region and belonging to the upper (α) band are shown in red while events, in the same energy interval, belonging to the lower (β/γ) band are shown in black. The theoretical discrimination power (DP) between the two distributions is usually

- -	1.1	1	2	
d	DI	e	2	

The theoretical discrimination power (DP) as defined in Eq. 3 for all the shape variables.

Shape variable	DP
$ au_R$	4.9
$ au_D$	4.6
TVL	4.2
TVR	20.0
χ^2_{OT}	6.2

quantified as the difference between the average values of the two distributions normalized to the square root of the quadratic sum of their widths:

$$DP = \frac{\mu_{\beta/\gamma} - \mu_{\alpha}}{\sqrt{\sigma_{\beta/\gamma}^2 + \sigma_{\alpha}^2}}.$$
(3)

The above definition does not imply any probabilistic meaning, it was used for the sake of comparison with similar quantities reported in published papers [14,15]. The actual discrimination power, in fact, should be directly measured using event populations with the required high statistics, since real event populations can exhibit non-Gaussian tails which reduce the theoretical discrimination power.

The TVR distribution (see bottom panel in Fig. 4) exhibits a theoretical discrimination power of \approx 20. This is a very impressive result. For comparison, the Light Yield DP for the same events (Fig. 2) is \approx 8. The separation achievable with the shape algorithm alone is therefore much more powerful than the one obtained using the information of the light detector and allows to reject α events to any desirable level.

Results for all the shape parameters are reported in Table 2. Considering the outstanding TVR performance compared to other variables, it was not worth combining them into a multivariate analysis. For completeness, TVR and scintillation were combined [33] into a Fisher linear discriminant analysis [34,35], resulting in a discrimination power of \approx 24. The two variables showed only a modest correlation: +23%(-25%) for the $\beta/\gamma(\alpha)$ events.

4. Internal contaminations

As pointed out in Section 2, the internal contaminations were evaluated in a second, dedicated run. In this run only the cylindri-



Fig. 5. Energy spectrum obtained in the second run with the cylindrical crystal corresponding to 407 h of measurements. The spectrum includes also a 232 Th calibration. The α events, recognized through the TVR parameter, are in red. (For interpretation of the references to colour in this figure legend, the reader is referred to the web version of this article.)

 Table 3

 Evaluated internal radioactive contaminations for the cylindrical crystal. Limits are at 68% CL.

Chain	Nuclide	Activity µBq/kg
²³² Th	²³² Th ²²⁸ Th	<32 <32
²³⁸ U	²³⁸ U ²³⁴ U ²³⁰ Th ²²⁶ Ra ²¹⁰ Po	350 ± 90 <110 <68 <32 1660 ± 200

cal crystal was measured. The set-up of the detector was identical to the one sketched in Fig. 1, but this time the LD was not present. In order to avoid surface radioactivity, the Cu facing the crystal was covered by several layers of polyethylene, the same used in CUORE Crystal Validation Runs [36]. The detector was run in similar temperature conditions of the previous run (see Table 1). The obtained energy spectrum, consisting of a sum of background and ²³² Th calibration spectra (totalling 407 h), is presented in Fig. 5.

The α induced background shows clearly three peaks. Two of them are identified as induced by ²¹⁰Po (internal, 5407 keV, and on surface, 5304 keV) decay, while the third one is due to an internal contamination of ²³⁸U. No other structure is visible on the spectrum, resulting in a flat continuum. In particular no events are recorded above 6 MeV, implying no decays induced by the most dangerous radionuclides for DBD searches, namely the ones that undergo β/γ decay with Q-values larger than $Q_{\beta\beta}$ (see Section 5.2.1). The evaluated internal radioactive contaminations are presented in Table 3.

It is remarkable that this crystal, grown without any dedicated effort in cleanliness issues, shows a radiopurity that is significantly better with respect to other molybdates, in particular CaMoO₄, already proposed [37,38] for future DBD experiments.

5. Background simulation

5.1. General consideration

The study of background issues for an experiment aiming at reaching high sensitivity on 0vDBD search can be divided into different types of sources depending on their position. In the following we will analyze in detail the "near sources" (crystal and copper mounting structure contaminations). "Far sources" (radioactive contaminations in the experimental apparatus, mainly the refrigerator and its external shields) and "environmental sources" (contribution present at the experimental site: mainly muons, neutrons and γ rays) are instead closely dependent on the experimental apparatus and environment in which the set-up is installed. A detailed description concerning these categories of sources in Cuoricino, the inspirer of this configuration can be found in [39].

Background induced by far sources is produced by γ -rays since β 's and α 's can be easily shielded. The choice of an isotope with a Q-value exceeding 2615 keV (by far the most intense -high energy-natural γ -line emitted by the decay of ²⁰⁸Tl, belonging to the ²³²Th decay chain) allows to greatly reduce this background. Above this energy there are only extremely rare high energy γ 's from ²¹⁴Bi: the total Branching Ratio (BR) in the energy window from 2615 up to 3270 keV is 0.15%.

In this frame it is mandatory to take into account also coincidence events that can occur in one crystal. The most relevant case is represented by the decay of ²⁰⁸Tl. In this decay, for example, there is 85% probability that the 2615 keV γ is emitted in cascade with the 583 keV γ . In fact, as we will evaluate in Section 5.2.2, this mechanism represents the largest source of background above



Fig. 6. Geometry used for the simulation of the ZnMoO4 array: a tower-like detector of 40 cylindrical crystals (h = 60 mm, $\emptyset = 60 \text{ mm}$), where each crystal is faced by one Ge LD (h = 0.5 mm, $\emptyset = 60 \text{ mm}$).

2615 keV. On the other hand it is clear that the probability of simultaneous interaction in the same crystal strongly depends on the relative distance from the detector. Moreover the lower energy γ is easily shielded compared to the higher energy one, so that this type of background can be completely neglected, in the case of "far sources", with the installation of a proper shielding.

Finally the last contribution that must be considered, for what concerns environmental sources, is the one coming from neutron and muon interactions in the detector. For these we consider the results reported in [39,40] by the CUORE collaboration, for a modular TeO₂ bolometric detector operated underground, with a geometrical configuration similar to the one proposed here (see Fig. 6). The change in the molecular compound does not heavily affect the background rates. Indeed, as in the case of TeO₂ crystals, none of the isotopes of ZnMoO₄ have high neutron cross sections, the main parameter that could affect the background. External neutrons yield an integrated contribution in the 3–4 MeV region < 10^{-5} c/(keV kg y) [39], while external muons - operating the array in anticoincidence - yield a contribution of < 10^{-4} c/(keV kg y) in the 2–4 MeV region [40].

5.2. Near sources

The region of interest (ROI) for the evaluation of the background is a few FWHM wide and centered at the 0vDBD Q-value. For ZnMoO₄ crystals the ROI has therefore been chosen between 3014 and 3054 keV. All the background levels quoted in the following have been defined running dedicated Monte Carlo simulations based on GEANT4 code. A tower-like detector (see Fig. 6) was simulated, made out of 40 detectors (ZnMoO₄ crystal + Ge LD) arranged in a modular configuration of 10 floors with 4 crystals each. The ZnMoO₄ crystals here considered are cylinders (h = 60 mm, $\varnothing = 60 \text{ mm}$) faced to Ge slabs of h = 0.5 mm, $\varnothing = 60 \text{ mm}$. Despite the results obtained in Section 3 we conservatively decided to simulate the tower structure including the light detectors. The possibility to have a second, almost independent, way to discriminate α particles cannot be completely discarded. As stated in Section 3 the possible presence of outliers (due to non-Gaussian tails) in the distributions of Fig. 4 (not observed in this work) cannot be a priori discarded at this level of R&D. In any case the presence of a thin Ge crystal does not affect appreciably the result of the simulations.

The total mass of the ZnMoO₄ detectors is 29.3 kg, which corresponds to 1.20 kg of ¹⁰⁰Mo in the case of non enriched crystals and to a mass of 11.22 kg in the case of enriched crystals at the 90% level. The design of the detectors is done in such a way as to reduce the amount of inert material facing the detector and at the same time to overcome the problem of differential thermal contractions between the detector and structure. The most "massive" materials employed for the detectors assembly are copper (skeleton structure) and PTFE (detector holders), and the mass ratio between these two components is about 100:2. The contribution of PTFE is neglected due to the reduced mass and relatively high radio-purity of this material [41]. Thus the only relevant near sources of background are radioactive contaminations of the copper structure and of the scintillating crystals themselves. In both cases not only γ 's, but also β 's contribute to the counting rate in the 0 ν DBD region.

The contamination of the Ge wafers, if any, will not be a problem: (I) high purity Ge slab will be used as LD's with negligible radioactive internal contaminations, (II) radioactive decay in the LD produce thermal pulses with a different shape than events produced by scintillation light. These events can therefore be easily recognized and rejected.

In the following sections, all the possible near sources will be considered separately.

5.2.1. Crystal bulk contaminations

In the case of crystal bulk contaminations the main contribution to the background rate in the ROI arises from the β/γ emissions of the natural radioactive chains. Indeed the kinetic energy of α 's emitted by U and Th daughters are always far above the ROI. The β emitters that give an important contribution to the background are ²⁰⁸Tl (²³²Th chain), ²¹⁴Bi and ²¹⁰Tl (²³⁸U chain).

²⁰⁸Tl has a Q-value of 5 MeV and decays through a β channel with a half-life of 3 min. It gives rise to a continuous spectrum which, in combination with γ 's emitted in the re-assessment of the daughter nucleus, can produce events falling in the ROI. However, the background induced by this isotope can be effectively reduced by means of delayed coincidences between the ²⁰⁸Tl signal and the α emitted by its precursor, ²¹²Bi (Q-value = 6.2 MeV). The coincidence pattern to be studied would be an event in the 0*v*DBD region preceded by an α decay of 6.2 MeV. Provided that the α counting rate is not too high, the dead time introduced by this method is low or even negligible.

In the ²³⁸U decay chain the only contribution above 2.7 MeV comes from the β decay of ²¹⁴Bi and ²¹⁰Tl. ²¹⁴Bi decays with a BR of 99.98% (*Q*-value = 3.27 MeV) to ²¹⁴Po which α decays with a very short half-life (163 µs) and a *Q*-value of 7.8 MeV. Since the rise time in our bolometers is of the order of ms, the β events add up to α -events producing events with energies higher than 7.8 MeV (the so called Bi-Po's). These events therefore do not produce background into the ROI. In the residual 0.02% of cases, ²¹⁴Bi α -decays to ²¹⁰Tl that is a β emitter with a *Q*-value of 5.5 MeV and an half-life of 1.3 min. The background induced by this β emitter can be easily rejected with a cut on the delayed coincidences between the ²¹⁰Tl signal and the ²¹⁴Bi α , as it happens for ²¹²Bi.

For the sake of completeness we neglect, at this level, internal contaminations of "exotic" long lived, high energy (Q > 3 MeV, τ > 0.3 years) β emitters, the most dangerous being ⁴²Ar (⁴²K), ¹⁰⁶Ru (¹⁰⁶Rh) and ¹²⁶Sn (¹²⁶Sb).

5.2.2. Copper bulk contamination

Copper is the most abundant material in the experimental setup near the detectors. The most important contribution to the counting rate in the ROI might be given by ²¹⁴Bi high energy γ 's, and by ²⁰⁸Tl γ 's at 583 and 2615 keV that, being emitted in cascade, can sum up to 3.2 MeV.

A proper selection of high quality copper would be enough to guarantee a negligible contribution from these sources. Indeed, simulating a bulk contamination of 19 μ Bq/kg in ²²⁸Th and 16 μ Bq/kg in ²²⁶Ra (limits on NOSV copper [42]) in the detector copper structure, the overall contribution in the ROI is 2.40 ± 0.04 \cdot 10⁻⁴ c/(keV kg y).

Considering the same bulk contaminations for the copper of the 10 mK shielding (the inner thermal shielding, 2 mm thick) the total background expected is $1.40 \pm 0.05 \cdot 10^{-4}$ c/(keV kg y). The background due to the other, more distant, thermal shielding and the other far parts of the cryostat is lower and therefore, negligible.

5.2.3. Surface contaminations

An important contribution to the background of the experiment may arise from surface contamination, specially from degraded α coming out mostly from surfaces of inert materials facing the detector [8,9]. Two procedures can be applied for its reduction: α rejection based on pulse shape analysis (or the scintillation signals) and the use of anticoincidences between the detectors of the array. This last possibility is, however, useful only for crystals surface contaminations. It does not allow in any way to recognize the α background due to surface contamination of inert material facing the detectors. Taking into account the discrimination power obtained both with pulse shape analysis and scintillation signals described in Section 3, this background source can be completely neglected.

5.2.4. β/γ Pile-up

Finally, one has to take into account that for slow detectors, like bolometers, even a feeble β/γ emission with Q < 3 MeV can produce an unwanted background due to the difficulty to recognize and reject pile-up on the rise time of the thermal pulse. Indeed independent signals can randomly add generating a background that extends even beyond the ROI.

This background depends mainly on 3 factors: the time within which is impossible to distinguish two separate events, the rate of β/γ interactions and their energy distribution. With respect to the rate due to β/γ events arising from near sources or internal contamination, it can be easily evaluated given the contaminations of Sections 5.2.1, 5.2.2. This rate is of the order of 0.1 mHz. Even considering the far sources (that are negligible in the ROI but contribute in the 0–2615 keV region) the total rate is well below 1 mHz.

In fact it turns out that the main contribution arises from the 2vDBD of ¹⁰⁰Mo, due to its relatively "fast" decay time. The expected rate in a 90% enriched crystal (h = 60 mm, $\emptyset = 60 \text{ mm}$) is 5.45 mHz. Unfortunately, the energy spectra of the 2vDBD exhibit a maximum at rather high energy ($\simeq 1.25 \text{ MeV}$) with a "broad" distribution, so that the pile-up energy spectrum distribution falls very close to the 0v-ROI.

The pile-up is recognized through the pulse shape analysis of the optimally filtered pulses. The time window in which we are not able to distinguish two pulses depends mostly on two parameters: the noise of the baseline and the relative pulse height of the two pulses. Our simulation (involving a dedicated pulse shape and noise generator) shows that the time in which two pulses cannot be recognized is definitely smaller than $\tau_R/10$. Here we conservatively assume as 5 ms the time in which we are not able to distinguish two different pulses. Giving this as input, we performed a Monte Carlo simulation of the 2vDBD of ¹⁰⁰Mo. The irreducible

Table 4

Background projection for a ZnMoO₄ tower-like experiment, as proposed in Fig. 6. Contaminations in Cu refer to the mounting structure. The background reported for internal contaminations in U and Th chains takes into account the background suppression obtained with delayed coincidences at 5 τ (i.e. \approx 9 min for the U chain and \approx 22 min for the Th chain). The dead time therefore introduced is 0.01% for U chain and 2.0% for Th chain. The background due to the 2*v*DBD of ¹⁰⁰Mo reported here corresponds to 90% enriched crystals. Only statistical errors are quoted.

Source	Position	Background [c/(keV kg y)]
U chain Th chain 2vDBD pile-up U + Th chains U + Th chains	Crystals bulk Crystals bulk Crystals bulk Cu frame bulk Cu shield bulk	$ \begin{array}{l} < 1.16 \pm 0.1 \cdot 10^{-5} \\ < 2.18 \pm 0.04 \cdot 10^{-4} \\ 1.96 \pm 0.36 \cdot 10^{-3} \\ < 2.40 \pm 0.04 \cdot 10^{-4} \\ < 1.40 \pm 0.05 \cdot 10^{-4} \end{array} $

2vDBD pile-up evaluated in the ROI is $(1.96 \pm 0.36) \cdot 10^{-3}$ c/ (keV kg y).

5.3. Total background evaluation.

In this section we report the background level that could be reached taking into account the above mentioned sources. When quoting background rates, we assume to operate the detectors in anticoincidence, thus recording only single hit events. Background contributions due to sources described in Sections 5.2.1, 5.2.2, 5.2.4 are reported in the last column of Table 4.

The background reported for internal contaminations in U and Th chains takes into account the suppression obtained with delayed coincidences at 5τ . The dead time therefore introduced is 0.01% for U chain and 2.0% for Th chain. The inefficiency introduced by those events where the α decay occurs near the crystal surface and the α particle escapes the crystal is of about 0.1%. This evaluation is based on Monte Carlo simulations. In this case we assume to loose any possibility of delayed coincidence, and we compute the background as the sum of the fraction of the untagged β events due to the α escape and the fraction of the untagged β events because they are out of the maximum allowed delayed coincidence window.

Contribution due to the other "near sources" (i.e. the Cu frames and the Cu 10 mK shield bulk contaminations) are evaluated considering both the ²¹⁴Bi high energy γ -lines and the ²⁰⁸Tl γ 's cascade. The induced background spectra due to ²³²Th and ²³⁸U in the Cu frames is reported in Fig. 7.



Fig. 7. Simulated background spectra induced by ²³² Th (red) and ²³⁸U (green) contamination in the Cu of the frames. The one induced by the 10 mK Cu shield is similar in shape, but a factor \simeq 2 smaller. The background induced by the unrecognized 2vDBD pile-up (blue) clearly dominates. (For interpretation of the references to colour in this figure legend, the reader is referred to the web version of this article.)

5.4. Experimental sensitivity

The maximum sensitivity reachable by a 0*v*DBD experiment corresponds to the "zero background" condition. This occurs when $(B \cdot M \cdot T \cdot \Delta) \simeq 1$ [43], where *B* is the background level per unit mass, energy, and time, *M* is the detector mass, *T* is the measuring time and Δ is the FWHM energy resolution. For a "standard" live-time for a 0*v*DBD experiment (i.e. 5 years), an energy resolution of 5 keV and a detector mass of ~30 kg the "zero background" condition will be reached with a background of $\approx 1.5 \cdot 10^{-3}$ c/(keV kg y).

Considering the different sources of background discussed in the previous sections and summarized in Table 4, with the detector array described above, for an enrichment of 90%, in 5 years of data taking we are able to fulfill the "zero background" condition.

For such an experiment, for a neutrino mass of $\langle m_{\beta\beta} \rangle = 100$ meV we will expect between 1.7 and 13.9 counts of 0*v*DBD depending on the Nuclear Matrix Element [44–47].

6. Conclusions

We successfully tested two ZnMoO₄ crystals as bolometers. The separation achievable on the shape of the thermal signal alone is much more powerful than the one obtained using the information of the light detector and allows to reject α events to any desirable level. Furthermore we would like to point out a very important consideration in favour of the pulse shape discrimination. The light collection generally depends on the size of the scintillating crystal, due to self absorption mechanisms. Since the LY of the ZnMoO₄ is rather small, this could imply a decrease in the achievable discrimination power obtainable through the light detection, moving from a few tenth grams crystals to a few hundreds grams crystals. On the contrary, this mechanism will enhance the pulse shape discrimination since the self-absorbed light signal will convert into heat, summing up with the non-radiative de-excitations that makes possible the α vs β/γ discrimination.

Moreover, even without any kind of material selection, the internal contaminations in ²²⁸Th and ²²⁶Ra are already at an extremely low level, never obtained in any crystal compound based on Molybdenum. The projection of these results to a small-size array foresees an experiment that could operate close to the "zero background" condition.

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