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A note on the complete rotational invariance OF BIRADIAL SOLUTIONS TO SEMILINEAR ELLIPTIC EQUATIONS

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# A NOTE ON THE COMPLETE ROTATIONAL INVARIANCE OF BIRADIAL SOLUTIONS TO SEMILINEAR ELLIPTIC EQUATIONS 

L. ABATANGELO AND S. TERRACINI

Abstract. We investigate symmetry properties of solutions to equations of the form

$$
-\Delta u=\frac{a}{|x|^{2}} u+f(|x|, u)
$$

in $\mathbb{R}^{N}$ for $N \geq 4$, with at most critical nonlinearities. By using geometric arguments, we prove that solutions with low Morse index (namely 0 or 1 ) and which are biradial (i.e. are invariant under the action of a toric group of rotations), are in fact completely radial. A similar result holds for the semilinear Laplace-Beltrami equations on the sphere. Furthermore, we show that the condition on the Morse index is sharp. Finally we apply the result in order to estimate best constants of Sobolev type inequalities with different symmetry constraints.

## 1. Introduction and statement of the result

Let $x=(\xi, \zeta) \in \mathbb{R}^{k} \times \mathbb{R}^{N-k}$, with $k, N-k \geq 2$. A function $u: \mathbb{R}^{N} \rightarrow \mathbb{R}$ is termed biradial if it is invariant under the action of the subgroup $S O(k) \times S O(N-k)$ of the group of rotations, namely, if there exists $\varphi: \mathbb{R}^{+} \times \mathbb{R}^{+} \rightarrow \mathbb{R}$ such that $u(\xi, \zeta)=\varphi(|\xi|,|\zeta|)$. Consider the equation

$$
\begin{equation*}
-\Delta u=\frac{a}{|x|^{2}} u+f(|x|, u) \quad \text { in } \mathbb{R}^{N} \backslash\{0\}, \tag{1}
\end{equation*}
$$

in this paper, we wonder under what circumstances it is possible to assert that a biradial solution to (1) is actually radially symmetric.

This problem arises from [12], where the following symmetry breaking result is given for the critical nonlinearity $f(|x|, u)=u^{(N+2) /(N-2)}$ : if $a<0$ and $|a|$ is sufficiently large, there are at least two distinct positive solutions, one being radially symmetric and the second not. These solution are obtained by minimization of the associated Rayleigh quotient over functions possessing either the full radial symmetry or a discrete group of symmetries, namely, for given $k \in \mathbb{Z}$, functions which are invariant under the $\mathbb{Z}_{k} \times S O(N-2)$-action on $D^{1,2}\left(\mathbb{R}^{N}\right)$ given by

$$
u(\xi, \zeta) \mapsto v(\xi, \zeta)=u(R \xi, T \zeta),
$$

$T$ being any rotation of $\mathbb{R}^{N-2}$ and $R$ a fixed rotation of order $k$. Once proved that the infimum taken over the $\mathbb{Z}_{k} \times S O(N-2)$-invariant functions is achieved, by comparing its value with the infimum taken over the radial functions, one deduces the occurrence of symmetry breaking (see also [1]).

In order to obtain multiplicity of solutions, the first attempt is to increase the order $k$ of the symmetry group and, eventually, to let it diverge to infinity, finding in the limit a minimizer of the Rayleigh quotient over the biradial functions. Now, will all these solutions be distinct and different from the radial one? When examining this question, we need to take

[^0]into account the construction due to Ding of an infinity of nontrivial biradial solutions to the Lane-Emden equation with critical nonlinearity (cfr [4]). In that case it is well known that there is a unique family of radially symmetric solutions, which are the global minimizers of the Rayleigh quotients, while in Ding's construction the nontrivial biradial solutions have a Morse index larger than 2.

We recall the following definition:
Definition 1.1. The (plain, radial, biradial) Morse index of a solution $u$ is the dimension of the maximal subspace of the space of (all, radial, biradial) functions of $\mathcal{C}_{0}^{\infty}\left(\mathbb{R}^{N} \backslash\{0\}\right)$ on which the quadratic form associated to the linearized equation at $u$ is negative definite.

We stress it is rather a geometric definition, so it is independent from any spectral theory about the differential operator we are dealing with.

The recent literature indicates that, for general semilinear equations, solutions having low Morse index do likely possess extra symmetries. Following these ideas and questions, we investigated in particular the biradial solutions with a low Morse index, and we are able to prove the following
Theorem 1.2. Let $u \in D^{1,2}\left(\mathbb{R}^{N}\right)$ be a biradial solution to

$$
\begin{equation*}
-\Delta u=\frac{a}{|x|^{2}} u+f(|x|, u) \tag{2}
\end{equation*}
$$

with $a>-\left(\frac{N-2}{2}\right)^{2}$ and $f: \mathbb{R}^{N} \times \mathbb{R} \rightarrow \mathbb{R}$ being a Carathéodory function, $C^{1}$ with respect to $z$, such that it satisfies the growth restriction

$$
\left|f_{y}^{\prime}(|x|, y)\right| \leq C\left(1+|y|^{2^{*}-2}\right)
$$

for a.e. $x \in \mathbb{R}^{N}$ and for all $y \in \mathbb{C}$.
If the solution $u$ has biradial Morse index $m(u) \leq 1$, then $u$ is radially symmetric.
An analogous result also holds for bounded domains having rotational symmetry, and for elliptic equations on the sphere. The following result holds in any dimension $N \geq 3$ :
Theorem 1.3. Let $f \in \mathcal{C}^{1}(\mathbb{R} ; \mathbb{R})$ : if $u \in \mathcal{C}^{2}\left(\mathbb{S}^{N}\right)$ is a biradial solution to

$$
-\Delta_{\mathbb{S}^{N}} u=f(u)
$$

with $N \geq 3$, and it has biradial Morse index $m(v) \leq 1$, then $u$ is constant on the sphere $\mathbb{S}^{N}$.
The paper is organized as follows: the next section is devoted to introduce the main tools and facts which will play a key role within the proof; in section 3 we present the proofs of Theorems 1.2 and 1.3 splitting it according to solutions' Morse index. In section 4 we give applications to the estimate of the best constants in some Sobolev type embeddings with symmetries. Finally section 5 is devoted to the discussion of the sharpness of the Theorems with respect to the Morse index.

## 2. Preliminaries

Here we start the proof of Theorem 1.2. For the sake of simplicity, we will work in dimension $N=4$. We devote the last part of the proof to discuss the validity of the result in higher dimensions.

Let us consider the following three orthogonal vector fields in $\mathbb{R}^{4}$ :

$$
X_{1}=\left[\begin{array}{r}
x_{2} \\
-x_{1} \\
x_{4} \\
-x_{3}
\end{array}\right], \quad X_{2}=\left[\begin{array}{r}
x_{4} \\
x_{3} \\
-x_{2} \\
-x_{1}
\end{array}\right], \quad X_{3}=\left[\begin{array}{r}
-x_{3} \\
x_{4} \\
x_{1} \\
-x_{2}
\end{array}\right]
$$

The related derivatives

$$
w_{i}=\nabla u \cdot X_{i}, \quad i=1,2,3
$$

represent the infinitesimal variations of the function $u$ along the flows of the vector fields $X_{i}$ respectively. As the equation is invariant under the action of such flows, these directional derivatives are solutions to the linearized equation

$$
\begin{equation*}
-\Delta w-\frac{a}{|x|^{2}} w=f_{y}^{\prime}(|x|, u) w \tag{3}
\end{equation*}
$$

We can associate the singular differential operator

$$
\begin{equation*}
L_{u} w=-\Delta w-\frac{a}{|x|^{2}} w-f_{y}^{\prime}(|x|, u) w \tag{4}
\end{equation*}
$$

Remark 2.1. The vector space of $\left\{X_{1}, X_{2}, X_{3}\right\}$ generates the whole group of infinitesimal rotations on the sphere of $\mathbb{R}^{4}$, which can be structured as a 3-dimensional manifold. In order to prove Theorem 1.2 it will be sufficient to show that every $w_{i} \equiv 0$.

Obviously, we have $w_{1} \equiv 0$ because the vector field $X_{1}$ generates the rotations under which the function $u$ is invariant for. Let us fix polar coordinates

$$
\left\{\begin{array} { l } 
{ x _ { 1 } = r _ { 1 } \operatorname { c o s } \theta _ { 1 } }  \tag{5}\\
{ x _ { 2 } = r _ { 1 } \operatorname { s i n } \theta _ { 1 } }
\end{array} \quad \left\{\begin{array}{l}
x_{3}=r_{2} \cos \theta_{2} \\
x_{4}=r_{2} \sin \theta_{2}
\end{array}\right.\right.
$$

we have $r_{1}=\sqrt{x_{1}^{2}+x_{2}^{2}}, r_{2}=\sqrt{x_{3}^{2}+x_{4}^{2}}$ and $\theta_{1}=\arctan \frac{x_{2}}{x_{1}}, \theta_{2}=\arctan \frac{x_{4}}{x_{3}}$.
Therefore, since $u$ is biradial, we have

$$
w_{i}=\nabla u \cdot X_{i}=w\left(r_{1}, r_{2}\right) z_{i}\left(\theta_{1}, \theta_{2}\right), \quad i=2,3
$$

where

$$
w\left(r_{1}, r_{2}\right)=\frac{\partial u}{\partial r_{1}} r_{2}-\frac{\partial u}{\partial r_{2}} r_{1}, \quad z_{2}=\sin \left(\theta_{1}+\theta_{2}\right), \quad z_{3}=-\cos \left(\theta_{1}+\theta_{2}\right)
$$

Remark 2.2. According to Remark 2.1, to our aim it will be sufficient to prove that $w \equiv 0$.
We now focus our attention on a few fundamental properties of the functions $w_{i}$. At first, as the $z_{i}$ 's are spherical harmonics and depend on the angles $\theta_{1}$ and $\theta_{2}$ only, we have

$$
\begin{equation*}
-\Delta z_{i}=\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right) z_{i} \quad \text { for } i=2,3 \tag{6}
\end{equation*}
$$

Joining this with the linearized equation (3) solved by the $w_{i}$ 's, we obtain the equation for $w$.
Proposition 2.3. The function $w$ is a solution to the following equation

$$
\begin{equation*}
-\Delta w-\frac{a}{|x|^{2}} w-f_{y}^{\prime}(|x|, u) w+\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right) w=0 \tag{7}
\end{equation*}
$$

Proof. It holds that

$$
f_{y}^{\prime}(|x|, u) w_{i}=-\Delta\left(w z_{i}\right)-\frac{a}{|x|^{2}} w z_{i}=-\Delta w z_{i}-\nabla w \cdot \nabla z_{i}-w \Delta z_{i}-\frac{a}{|x|^{2}} w z_{i}
$$

Since $\nabla w \cdot \nabla z_{i}=0$, thanks to (6), this becomes

$$
-\Delta w z_{i}-\frac{a}{|x|^{2}} w z_{i}=f_{y}^{\prime}(|x|, u) w z_{i}+w \Delta z_{i}=f_{y}^{\prime}(|x|, u) w z_{i}-\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right) w z_{i}
$$

that is

$$
z_{i}\left\{-\Delta w-\frac{a}{|x|^{2}} w-f_{y}^{\prime}(|x|, u) w+\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right)\right\}=0
$$

Last, multiplying by $z_{i}$ and summing for $i=1,2$ we obtain the desired equation.

## 3. Proofs

We will split the argument according to the Morse index of solution $u$ : we denote it by $m(u)$.

In order to complete our proof, we need a couple of preliminary results: the first one is about the asymptotics of the solution and is contained in [7].

Lemma 3.1. ([7]) Under the assumptions of Theorem 1.2, let $u$ be any solution to (1). Then the following asymptotics hold

$$
\begin{array}{ll}
u(x) \sim|x|^{\gamma} \psi\left(\frac{x}{|x|}\right) & \text { for }|x| \ll 1  \tag{8}\\
u(x) \sim|x|^{\delta} \psi\left(\frac{x}{|x|}\right) & \text { for }|x| \gg 1
\end{array}
$$

where $\gamma=\gamma(a, N)=-\frac{N-2}{2}+\sqrt{\left(\frac{N-2}{2}\right)^{2}+\mu}, \delta=\delta(a, N)=-\frac{N-2}{2}-\sqrt{\left(\frac{N-2}{2}\right)^{2}+\mu}$ and $\mu=\mu(a, N)$ is one of the eigenvalues of $-\Delta_{\mathbb{S}^{N-1}}-a$ on $\mathbb{S}^{N-1}$, and $\psi$ one of its related eigenfunctions.

This turns out to be the key for proving the following result.
Lemma 3.2. The function $\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right) w^{2}$ is $L^{1}$-integrable on $\mathbb{R}^{N}$.
Proof. Since $w\left(r_{1}, r_{2}\right)=\frac{\partial u}{\partial r_{1}} r_{2}-\frac{\partial u}{\partial r_{2}} r_{1}$, we first observe that by regularity of $u$ outside the origin and its radial symmetry, the functions

$$
\frac{1}{r_{i}} \frac{\partial u}{\partial r_{i}}
$$

$i=1,2$, are continuous outside the origin. Next we remark that

$$
\begin{aligned}
\frac{1}{4}\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right) w^{2} \leq\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right) & \left\{\left(\frac{\partial u}{\partial r_{1}}\right)^{2} r_{2}^{2}+\left(\frac{\partial u}{\partial r_{2}}\right)^{2} r_{1}^{2}\right\} \\
= & \left(\frac{r_{2}}{r_{1}}\right)^{2}\left(\frac{\partial u}{\partial r_{1}}\right)^{2}+\left(\frac{r_{1}}{r_{2}}\right)^{2}\left(\frac{\partial u}{\partial r_{2}}\right)^{2}+\left(\frac{\partial u}{\partial r_{1}}\right)^{2}+\left(\frac{\partial u}{\partial r_{2}}\right)^{2}
\end{aligned}
$$

The integrability of the last two terms is a straightforward consequence of $u \in D^{1,2}\left(\mathbb{R}^{N}\right)$. In order to study the other two terms, let us focus our attention in a ball around the origin, namely $B_{1}(0)$, so that $r_{1}^{2}+r_{2}^{2} \leq 1$. Then

$$
\left(\frac{\partial u}{\partial r_{1}}\right)^{2}\left(\frac{r_{2}}{r_{1}}\right)^{2} \leq \frac{1}{r_{1}^{2}}\left(\frac{\partial u}{\partial r_{1}}\right)^{2}-\left(\frac{\partial u}{\partial r_{1}}\right)^{2}
$$

so that the question of integrability is restricted to the first term. From Lemma 3.1, Equation (8) we know $u \sim r^{\gamma} \psi\left(r_{1}, r_{2}\right)=\left(r_{1}^{2}+r_{2}^{2}\right)^{\gamma / 2} \psi\left(r_{1}, r_{2}\right)$, from which

$$
\frac{\partial u}{\partial r_{1}} \sim \psi\left(r_{1}, r_{2}\right) \gamma\left(r_{1}^{2}+r_{2}^{2}\right)^{\gamma / 2-1} r_{1}+\left(r_{1}^{2}+r_{2}^{2}\right)^{\gamma / 2} \frac{\partial \psi}{\partial r_{1}}
$$

So we are lead to consider the integrability of $\int_{B_{1}(0)} \frac{1}{r_{1}^{2}}\left(\frac{\partial \psi}{\partial r_{1}}\right)^{2}$. Additionally we know that $\psi$ is the restriction on the sphere of a harmonic polynomial, then it is analytic and its Taylor's expansion is a polynomial whose degree 1 terms vanish, since it is a function of the only variables $r_{1}$ and $r_{2}$. Then $\left(\frac{\partial \psi}{\partial r_{1}}\right)^{2} \sim r_{1}^{2}$, which provides the sought integrability.

For what concerns the integrability at infinity, it is sufficient to show that the terms of type $\frac{r_{2}^{2}}{r_{1}^{2}}\left(\frac{\partial u}{\partial r_{1}}\right)^{2}$ are in $L^{1}\left(\mathbb{R}^{N}\right)$. We have

$$
\frac{r_{2}^{2}}{r_{1}^{2}}\left(\frac{\partial u}{\partial r_{1}}\right)^{2} \leq 2\left\{\left(r_{1}^{2}+r_{2}^{2}\right)^{\delta-2} r_{2}^{2} \psi^{2}+\frac{r_{2}^{2}}{r_{1}^{2}}\left(r_{1}^{2}+r_{2}^{2}\right)^{\delta}\left(\frac{\partial \psi}{\partial r_{1}}\right)^{2}\right\}
$$

and exploiting equation (9), the expression of the exponent $\delta$ provides the sought integrability.

In the following we consider the cut-off function defined as $\eta\left(r_{1}, r_{2}\right)=\eta_{1}\left(r_{1}\right) \eta_{2}\left(r_{2}\right)$ where

$$
\eta_{1}\left(r_{1}\right)= \begin{cases}\frac{1}{\log \left(R_{2} / R_{1}\right)} \log r_{1} / R_{1} & \text { for } R_{1} \leq r_{1} \leq R_{2} \\ 1 & \text { for } R_{2} \leq r_{1} \leq R_{3} \\ 1-\frac{1}{\log \left(R_{4} / R_{3}\right)} \log r_{1} / R_{3} & \text { for } R_{3} \leq r_{1} \leq R_{4} \\ 0 & \text { elsewhere }\end{cases}
$$

$\eta_{2}$ being defined similarly. Given the special form of $\eta$, we note $|\nabla \eta|^{2} \leq\left|\nabla \eta_{1}\right|^{2}+\left|\nabla \eta_{2}\right|^{2}$, that is

$$
|\nabla \eta|^{2} \leq \frac{1}{\log ^{2} R_{2} / R_{1}}\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right) \quad \text { for } R_{1} \leq r_{1}, r_{2} \leq R_{2}
$$

and analogously for $R_{3} \leq r_{1}, r_{2} \leq R_{4}$. Thus, we have

$$
\begin{equation*}
|\nabla \eta|^{2} \leq 3\left(\frac{1}{\log ^{2} R_{4} / R_{3}}+\frac{1}{\log ^{2} R_{2} / R_{1}}\right)\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right) \tag{10}
\end{equation*}
$$

Lemma 3.3. There is a suitable choice of the parameters $R_{1}, R_{2}, R_{3}$ and $R_{4}$ such that the quadratic form associated to the operator (4) is negative definite both on $\eta w^{+}$and $\eta w^{-}$.

Proof. Let us fix $\varepsilon>0$ small and choose $R_{1}=\varepsilon^{2}, R_{2}=\varepsilon$ and $R_{3}=\varepsilon^{-1}, R_{4}=\varepsilon^{-2}$. We multiply equation (7) by $\eta^{2} w^{+}$and integrate by parts. We obtain

$$
\begin{align*}
\int_{\mathbb{R}^{N}} & \left|\nabla\left(\eta^{2} w^{+}\right)\right|^{2}-\frac{a}{|x|^{2}}\left(\eta^{2} w^{+}\right)^{2}-f_{y}^{\prime}(|x|, u) \eta^{2}\left(w^{+}\right)^{2} \\
& =\int_{\mathbb{R}^{N}}|\nabla \eta|^{2}\left(w^{+}\right)^{2}-\int_{\mathbb{R}^{N}}\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right) \eta^{2}\left(w^{+}\right)^{2} \tag{11}
\end{align*}
$$

If $\varepsilon$ is small enough, the second term in (11) is far away from zero, or rather, it is quite close to $\int_{\mathbb{R}^{N}}\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right)\left(w^{+}\right)^{2}$, say for instance

$$
\int_{\mathbb{R}^{N}}\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right) \eta^{2}\left(w^{+}\right)^{2}>\frac{1}{2} \int_{\mathbb{R}^{N}}\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right)\left(w^{+}\right)^{2} .
$$

On the other hand, the first term in (11) can be made very small with respect to $\int_{\mathbb{R}^{N}}\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right)\left(w^{+}\right)^{2}$, since from (10)

$$
\int_{\mathbb{R}^{N}}|\nabla \eta|^{2}\left(w^{+}\right)^{2} \leq \frac{6}{\log ^{2} \varepsilon} \int_{\mathbb{R}^{N}}\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right)\left(w^{+}\right)^{2}
$$

so that (11) is seen to be negative.
Repeating the same argument multiplying by $\eta^{2} w^{-}$we reach the same conclusion.
First case: Morse index $m(u)=0$. In this case Lemma 3.3 clearly contradicts the hypothesis $m(u)=0$, unless $w^{+}=w^{-} \equiv 0$, that is the only stable solution to (7) is the trivial one.
Second case: Morse index $m(u)=1$. In this case we infer that $w$ has constant sign, say positive, and therefore $w>0$ for $r_{1}>0$ and $r_{2}>0$ by the Strong Maximum Principle. Now we show a contradiction. Consider a vector field of the form $\alpha X_{2}+\beta X_{3}$. Along this vector field, choosing $\alpha=\cos \gamma$ and $\beta=\sin \gamma$, the derivative of $u$ is
$\nabla u \cdot\left(\alpha X_{2}+\beta X_{3}\right)=\alpha w_{2}+\beta w_{3}=w\left(\alpha \sin \left(\theta_{1}+\theta_{2}\right)-\beta \cos \left(\theta_{1}+\theta_{2}\right)\right)=-w \sin \left(\theta_{1}+\theta_{2}-\gamma\right)$.
Now we turn to the directional derivative of $\theta_{1}+\theta_{2}$ along the vector field $\alpha X_{2}+\beta X_{3}$. Using the polar coordinates (5), it results

$$
\theta_{1}=\arctan \frac{x_{2}}{x_{1}}, \quad \theta_{2}=\arctan \frac{x_{4}}{x_{3}}
$$

so that checking the motion along $X_{2}$ we have

$$
\begin{aligned}
& \nabla \theta_{1} \cdot X_{2}=\frac{x_{3} x_{1}-x_{2} x_{4}}{x_{1}^{2}+x_{2}^{2}}=\frac{r_{2}}{r_{1}}\left(\cos \theta_{1} \cos \theta_{2}-\sin \theta_{1} \sin \theta_{2}\right)=\frac{r_{2}}{r_{1}} \cos \left(\theta_{1}+\theta_{2}\right) \\
& \nabla \theta_{2} \cdot X_{2}=\frac{-x_{3} x_{1}+x_{2} x_{4}}{x_{3}^{2}+x_{4}^{2}}=\frac{r_{1}}{r_{2}}\left(-\cos \theta_{1} \cos \theta_{2}+\sin \theta_{1} \sin \theta_{2}\right)=-\frac{r_{1}}{r_{2}} \cos \left(\theta_{1}+\theta_{2}\right)
\end{aligned}
$$

whereas along $X_{3}$

$$
\begin{aligned}
& \nabla \theta_{1} \cdot X_{3}=\frac{x_{4} x_{1}+x_{2} x_{3}}{x_{1}^{2}+x_{2}^{2}}=\frac{r_{2}}{r_{1}}\left(\cos \theta_{1} \sin \theta_{2}+\sin \theta_{1} \cos \theta_{2}\right)=\frac{r_{2}}{r_{1}} \sin \left(\theta_{1}+\theta_{2}\right) \\
& \nabla \theta_{2} \cdot X_{3}=\frac{-x_{2} x_{3}-x_{1} x_{4}}{x_{3}^{2}+x_{4}^{2}}=\frac{r_{1}}{r_{2}}\left(-\sin \theta_{1} \cos \theta_{2}-\cos \theta_{1} \sin \theta_{2}\right)=-\frac{r_{1}}{r_{2}} \sin \left(\theta_{1}+\theta_{2}\right)
\end{aligned}
$$

and finally we obtain

$$
\begin{aligned}
\nabla\left(\theta_{1}+\theta_{2}\right) \cdot\left(\alpha X_{2}+\beta X_{3}\right)=\left(\frac{r_{2}}{r_{1}}-\frac{r_{1}}{r_{2}}\right)\left(\alpha \cos \left(\theta_{1}+\theta_{2}\right)+\right. & \left.\beta \sin \left(\theta_{1}+\theta_{2}\right)\right) \\
& =\left(\frac{r_{2}}{r_{1}}-\frac{r_{1}}{r_{2}}\right) \cos \left(\theta_{1}+\theta_{2}-\gamma\right)
\end{aligned}
$$

Now we are in good position to conclude. For a given point $\bar{x}$ of the sphere - located by angles $\overline{\theta_{1}}$ and $\overline{\theta_{2}}$, we choose $\gamma=\gamma(\bar{x})=\overline{\theta_{1}}+\overline{\theta_{2}}-\pi / 2$, so that the quantity $\theta_{1}+\theta_{2}$ is at rest for the associated vector field $\cos \gamma X_{2}+\sin \gamma X_{3}$. With this choice the function $u$ is monotone along the flow $\alpha X_{2}+\beta X_{3}$ since $\dot{u}=-w \sin \left(\theta_{1}+\theta_{2}-\gamma\right)=-w$ and the $\operatorname{sign}$ of $w$ is constant by the previous discussion. Since the trajectory of the flow is a circle, we will reach again the initial point in finite time, but with a strictly smaller value of $u$ (if we consider the first eigenfunction $w$ positive). This is clearly a contradiction.
Generalization to higher dimensions. In dimension $N \geq 5$ the argument is very similar. Relabeling we may always assume $u=u\left(\rho_{1}, \rho_{2}\right)$, where we have fixed the notation $\rho_{1}=|\xi|$ and $\rho_{2}=|\zeta|$, while $|x|=\sqrt{|\xi|^{2}+|\zeta|^{2}}$, being $x=(\xi, \zeta) \in \mathbb{R}^{k} \times \mathbb{R}^{N-k}$. Now we repeat the argument performed in the 4 -dimensional space with respect to the variables $x_{k-1}, x_{k}, x_{k+1}$, $x_{k+2}$, considering the vector fields with those same four components as above and the other ones being zero. Hence we define $r_{1}=\sqrt{x_{k-1}^{2}+x_{k}^{2}}$ and $r_{2}=\sqrt{x_{k+1}^{2}+x_{k+2}^{2}}$. When discussing the integrability properties, it can be worthwhile noticing that

$$
\frac{1}{r_{i}} \frac{\partial u}{\partial r_{i}}=\frac{1}{\rho_{i}} \frac{\partial u}{\partial \rho_{i}}
$$

Arguing as above, we can prove that the solution $u$ is actually radial with respect to those four variables. We can imagine to iterate this proceeding for every hyperplane whose rotations the function $u$ is supposed not to be invariant for. Finally, it follows that $u$ is radial in $\mathbb{R}^{N}$.

Proof of Theorem 1.3. Since now $v$ is a function defined over $\mathbb{S}^{N}$, recalling the Laplace operator in polar coordinates

$$
\Delta_{\mathbb{R}^{N+1}}=\partial_{r}^{2}+\frac{N}{r} \partial_{r}+\frac{1}{r^{2}} \Delta_{\mathbb{S}^{N}}
$$

we define $\widetilde{v}(x)=v(y)$ for $x \in(-\varepsilon, \varepsilon) \times \mathbb{S}^{N}$, so that $\Delta_{\mathbb{S}^{N}} v=\Delta_{\mathbb{R}^{N+1}} \widetilde{v}$. At first, let us suppose $N=3$. Obviously, since $v$ is invariant with respect to the group $O(2) \times O(2)$, so is $\widetilde{v}$.

Following the same argument in the proof of Theorem 1.2, we wish to prove the vanishing of $\widetilde{w}=\frac{\partial \widetilde{v}}{\partial r_{1}} r_{2}-\frac{\partial \widetilde{v}}{\partial r_{2}} r_{1}$. On the other hand, being $\widetilde{v}$ homogenous of degree $0, \widetilde{w}$ is homogenuos of degree 0 too (it can be proved by differentiating identity $\widetilde{v}(x)=\widetilde{v}(\lambda x)$ ), then the $w$ associated with $v$ is nothing else that $\widetilde{w}$ restricted on the sphere $\mathbb{S}^{N}$. Therefore $\Delta_{\mathbb{R}^{N+1}} \widetilde{w}=\Delta_{\mathbb{S}^{N}} w$, and following the proof of Proposition 2.3 we see $w$ is a solution to

$$
-\Delta_{\mathbb{S}^{N}} w-f^{\prime}(v) w+\left(\frac{1}{r_{1}^{2}}+\frac{1}{r_{2}^{2}}\right) w=0
$$

analogous to equation (7). The rest of the proof fits also in this case.

## 4. An application to Best Sobolev constants with symmetries

Solutions to the critical exponent equation

$$
\begin{equation*}
-\Delta u=\frac{a}{|x|^{2}} u+|u|^{2^{*}-2} u \tag{12}
\end{equation*}
$$

are related to extremals of Sobolev inequalities (cfr [12]). To our purposes, the functions $u$ will be complex-valued and $a \in\left(-\infty,(N-2)^{2} / 4\right)$. Then, thanks to Hardy inequality, an equivalent norm on $D^{1,2}\left(\mathbb{R}^{N}\right)$ is

$$
\left(\int_{\mathbb{R}^{N}}|\nabla u|^{2}-a \frac{|u|^{2}}{|x|^{2}}\right)^{1 / 2}
$$

hence we can seek solutions to (12) as extremals of the Sobolev quotient associated with this norm on different symmetric spaces .

The whole group of rotations $S O(2) \times S O(N-2)$ induces the following action on $D^{1,2}\left(\mathbb{R}^{N} ; \mathbb{C}\right)$ :

$$
u(\xi, \zeta) \mapsto R^{-m} u(R \xi, T \zeta)
$$

for $m \in \mathbb{Z}$ fixed. We denote, as usual, $D_{\text {rad }}^{1,2}\left(\mathbb{R}^{N}\right)$ and $D_{\text {birad }}^{1,2}\left(\mathbb{R}^{N}\right)$ the subspaces of real or complex radial and biradial functions. Moreover, let $k$ and $m$ be fixed integers; for a given rotation $R \in S O(2)$ of order $k$, we consider the space of symmetric functions

$$
D_{R, k, m}^{1,2}\left(\mathbb{R}^{N} ; \mathbb{C}\right):=\left\{u \in D^{1,2}\left(\mathbb{R}^{N} ; \mathbb{C}\right): u(R \xi, T \zeta)=R^{m} u(\xi, \zeta), \forall T \in S O(N-2)\right\}
$$

This is of course a proper subspace of
$D_{\text {birad }, m}^{1,2}\left(\mathbb{R}^{N} ; \mathbb{C}\right):=\left\{u \in D^{1,2}\left(\mathbb{R}^{N} ; \mathbb{C}\right): u(S \xi, T \zeta)=S^{m} u(\xi, \zeta), \forall(S, T) \in S O(2) \times S O(N-2)\right\}$.
Note this last space coincides with the usual space of biradial solution once $m=0$.
Thanks to its rotational invariance, for any choice of the above spaces $D_{*}^{1,2}\left(\mathbb{R}^{N} ; \mathbb{C}\right)$, solutions to the minimization problem

$$
\begin{equation*}
\inf _{\substack{u \in D_{*}^{1,2}\left(\mathbb{R}^{N} ; \mathbb{C}\right) \\ u \neq 0}} \frac{\int_{\mathbb{R}^{N}}|\nabla u|^{2}-a \frac{|u|^{2}}{|x|^{2}}}{\left(\int_{\mathbb{R}^{N}}|u|^{2^{*}}\right)^{2 / 2^{*}}} \tag{13}
\end{equation*}
$$

are in fact solutions to equation (12).
The minimization of the Sobolev quotient over the space of radial functions follows from a nowadays standard compactness argument; in addition, see for instance [12], we have:

$$
\inf _{\substack{u \in D^{1,2},\left(\mathbb{R}^{N} ; \mathbb{C}\right) \\ \text { rof } \\ u \neq 0}} \frac{\int_{\mathbb{R}^{N}}|\nabla u|^{2}-a \frac{|u|^{2}}{|x|^{2}}}{\left(\int_{\mathbb{R}^{N}}|u|^{2^{*}}\right)^{2 / 2^{*}}}=S\left(1-a \frac{4}{(N-2)^{2}}\right)
$$

where $S$ denote the best constant for the standard Sobolev embedding. Moreover, generalizing the results in [3] in higher dimensions (see also [1]), one can easily prove existence of minimizers of the Sobolev quotient (13) in the spaces $D_{\text {birad }, m}^{1,2}\left(\mathbb{R}^{N} ; \mathbb{C}\right)$, for any choice of the integer $m$.

At first, let us consider the case $m=0$. Then it is easily checked that the minimizers can be chosen to be real valued and that the corresponding solution to (12) have biradial Morse index exactly one. Hence our Theorem 1.2 applies and such biradial solutions are in fact fully
radially symmetric, and therefore the infimum on the biradial space equals that on the radial. Now, let us turn to the case $m \neq 0$. We remark that elements of the space $D_{\text {birad, } m}^{1,2}\left(\mathbb{R}^{N} ; \mathbb{C}\right)$ have the form $u\left((\xi, \zeta)=\rho(|\xi|,|\zeta|) \mathrm{e}^{i m \theta(\xi)}\right.$, where $\theta(\xi)=\arg (\xi)$, so that

$$
|\nabla u|^{2}=|\nabla \rho|^{2}+\rho^{2}|m \nabla \theta|^{2}=|\nabla \rho|^{2}+m^{2} \frac{\rho^{2}}{|\xi|^{2}}
$$

Then the following chain of inequalities holds:

$$
\begin{aligned}
& \min _{\substack{u \in D_{\text {birad, }}^{\begin{subarray}{c}{1,2 \\
u \neq 0} }}}\end{subarray}} \frac{\int_{\left.\mathbb{R}^{N} ; \mathrm{C}\right)}|\nabla u|^{2}-a \frac{|u|^{2}}{|x|^{2}}}{\left(\int_{\mathbb{R}^{N}}|u|^{2^{*}}\right)^{2 / 2^{*}}}=\min _{\substack{\rho \in D_{\text {biridad }}^{1,2}\left(\mathbb{R}^{N}, \mathbb{R}\right) \\
\rho \neq 0}} \frac{\int_{\mathbb{R}^{N}}|\nabla \rho|^{2}+m^{2} \frac{\rho^{2}}{|\xi|^{2}}-a \frac{\rho^{2}}{|x|^{2}}}{\left(\int_{\mathbb{R}^{N}} \rho^{2^{*}}\right)^{2 / 2^{*}}} \\
& >\min _{\substack{\rho \in D_{\text {birad }}^{1,2}\left(\mathbb{R}^{N} ; \mathbb{R}\right) \\
\rho \neq 0}} \frac{\int_{\mathbb{R}^{N}}|\nabla \rho|^{2}+\left(m^{2}-a\right) \frac{\rho^{2}}{|x|^{2}}}{\left(\int_{\mathbb{R}^{N}} \rho^{2^{*}}\right)^{2 / 2^{*}}}=\min _{\substack{\rho \in D_{\begin{subarray}{c}{1,2 d \\
\rho \neq 0} }}^{\left.1, \mathbb{R}^{N} ; \mathbb{R}\right)}}\end{subarray}} \frac{\int_{\mathbb{R}^{N}}|\nabla \rho|^{2}+\left(m^{2}-a\right) \frac{\rho^{2}}{|x|^{2}}}{\left(\int_{\mathbb{R}^{N}} \rho^{2^{*}}\right)^{2 / 2^{*}}} \\
& =S\left(1+\frac{4\left(m^{2}-a\right)}{(N-2)^{2}}\right)
\end{aligned}
$$

where we have used $|\xi| \leq|x|$; the intermediate line follows again from Theorem 1.2 , and the last from [12]. Then, this argument states a very useful lower bound (see [1]) to the minima problems (13). Indeed, it allows us to compare the infimum over the space of $D_{R, k, m}^{1,2}\left(\mathbb{R}^{N} ; \mathbb{C}\right)$ with that on $D_{\text {birad, } m}^{1,2}\left(\mathbb{R}^{N} ; \mathbb{C}\right)$, and to prove the occurrence of symmetry breaking in some circumstances. In fact it has been proven (see [1]) that, for large enough $k$, the first minimum is achieved and less that $k^{2 / N} S$, while the latter increases with $|a|$ and $m$. Symmetry breaking holds whenever it can be shown that $1+\frac{4\left(m^{2}-a\right)}{(N-2)^{2}}>k^{2 / N}$ for appopriate choices of the parameters.

## 5. Optimality with respect to the Morse index

We want to stress our results Theorem 1.2 and 1.3 are sharp with respect to the Morse index. By that, we mean that doubly radial solutions with Morse index greater or equal to 2 , need not to be completely radial.

To prove this, we will take advantage from a result proved by Ding in [4] in such a way which will be clear later. The quoted paper by Ding has to do with solutions to a related equation on $\mathbb{S}^{N}$, for this reason we state first some connections between these two environments.
5.1. Conformally equivariant equations. We recall a general fact cited in [4] about elliptic equations on Riemannian manifolds.

Lemma 5.1. Let $(M, g)$ and $(N, h)$ two Riemannian manifolds of dimensions $N \geq 3$. Suppose there is a conformal diffeomorphism $f: M \rightarrow N$, that is $f^{*} h=\varphi^{2^{*}-2} g$ for some positive $\varphi \in C^{\infty}(M)$. The scalar curvatures of $(M, g)$ and $(N, h)$ are $R_{g}$ and $R_{h}$ respectively. Set the
following corresponding equations:

$$
\begin{align*}
& -\Delta_{g} u+\frac{1}{4} \frac{N-2}{N-1} R_{g}(x) u=F(x, u)  \tag{14}\\
& -\Delta_{h} v+\frac{1}{4} \frac{N-2}{N-1} R_{h}(y) v=\left[\left(\varphi \circ f^{-1}\right)(y)\right]^{-\frac{N+2}{N-2}} F\left(f^{-1}(y),\left(\varphi \circ f^{-1}\right)(y) v\right) \tag{15}
\end{align*}
$$

where $F: M \times \mathbb{R} \rightarrow \mathbb{R}$ is smooth. Suppose $v$ is a solution of (15). Then $u=(v \circ f) \varphi$ is a solution of (14) such that $\int_{M}|u|^{2^{*}} d V_{g}=\int_{N}|v|^{2^{*}} d V_{h}$.

We consider the inverse of the stereographic projection $\pi: \mathbb{S}^{N} \backslash\{p\} \rightarrow \mathbb{R}^{N}$. We denote it by $\Phi=\pi^{-1}: \mathbb{R}^{N} \rightarrow \mathbb{S}^{N} \backslash\{p\}$, moreover $g_{0}$ will denote the standard metric on $\mathbb{S}^{N}$ and $\delta$ the standard one on $\mathbb{R}^{N}$.

The diffeomorphism $\Phi$ is conformal between the two manifolds, since it results

$$
g \doteq \Phi^{*} g_{0}=\mu(x)^{\frac{4}{N-2}} \delta
$$

where

$$
\mu(x)=\left(\frac{2}{1+|x|^{2}}\right)^{\frac{N-2}{2}}
$$

In addition, we point out the manifold $\left(\mathbb{R}^{N}, g\right)$ is the same as $\left(\mathbb{S}^{N}, g_{0}\right)$, in terms of diffeomorphic manifolds.

We recall the following
Definition 5.2. We define the conformal Laplacian on a differentiable closed manifold ( $M, g$ ) of dimension $N$ the operator

$$
L_{g}=-\Delta_{g}+\frac{N-2}{4(N-1)} R_{g}
$$

where $\Delta_{g}$ denotes the standard Laplace-Beltrami operator on $M$ and $R_{g}$ the scalar curvature of the manifold.

Moreover, this operator has a simple transformation law under a conformal change of metric, that is

$$
\text { if } \quad \widetilde{g}=\mu(x)^{\frac{4}{N-2}} g \quad \text { then } \quad L_{\widetilde{g}} \cdot=\mu(x)^{-\frac{N+2}{N-2}} L_{g}(\mu(x) \cdot)
$$

In our case we are dealing with the same manifold $\mathbb{R}^{N}$ endowed with the two metrics $\delta$, the standard one, and $g=\Phi^{*} g_{0}$. Thus in our case we have

$$
L_{\delta}=-\Delta \quad L_{g}=-\Delta_{g}+\frac{1}{4} N(N-2)
$$

so it is quite easy to check directly the correspondence between the equations stated in Lemma 5.1 by calculations.
5.2. Proof of the optimality of Theorem 1.2 with respect to the Morse index. In this section we discuss the optimality of Theorems 1.3 with respect to the solutions' Morse index. First of all, we consider the the equation on the sphere $\mathbb{S}^{N}$ related to (2) through the weighted composition with the stereographic projection $\pi$ as conformal diffeomorphism from $\mathbb{S}^{N} \backslash\{p\}$ onto $\mathbb{R}^{N}$ : it is immediate to check that it is

$$
-\Delta_{\mathbb{S}^{N}} v(y)+\frac{1}{4} N(N-2) v(y)=f(v(y)) \quad y \in \mathbb{S}^{N}
$$

In his paper [4], Ding states the following result:
Lemma 5.3. There exists a sequence $\left\{v_{k}\right\}$ of biradial solutions to the equation

$$
\begin{equation*}
-\Delta_{\mathbb{S}_{N}} v+\frac{1}{4} N(N-2) v=|v|^{\frac{4}{N-2}} v \quad v \in C^{2}\left(\mathbb{S}^{N}\right) \tag{16}
\end{equation*}
$$

such that $\int_{\mathbb{S}^{N}}\left|v_{k}\right|^{\frac{2 N}{N-2}} d V \rightarrow \infty$ as $k \rightarrow \infty$.
The choice of working in a space of biradial is motivated by the compact embedding of the space of $H^{1}$-biradial functions on the sphere into $L^{2 N /(N-2)}$. In this way one can overcome the lack of compactness due to the presence of the critical exponent and prove the result as an application of the Ambrosetti-Rabinowitz symmetric Mountain Pass Theorem. We are interested in classifying the solutions according to their Morse index. We can state the following

Lemma 5.4. Among the solutions $\left\{v_{k}\right\}$ in Lemma 5.3 there is also a constant one, which is unique and corresponds to the minimum of Sobolev quotient. All the other biradial solutions have biradial Morse index at least 2, and there is at least one non constant biradial solution having Morse index exactly 2.

Proof. We can check directly there exists a unique constant solution:

$$
\frac{1}{4} N(N-2) c=c^{\frac{N+2}{N-2}} \quad \Longrightarrow \quad c=\left(\frac{1}{4} N(N-2)\right)^{\frac{N-2}{4}}
$$

which corresponds to the Talenti functions on the sphere ([11]). We mean it is the image of the function $w(x)=\frac{(N(N-2))^{\frac{N-2}{4}}}{\left(1+|x|^{2}\right)^{\frac{N-2}{2}}}=\mu(x) c$ through the diffeomorphism $\pi^{-1}$ and

$$
L_{g} c=\mu(x)^{-\frac{N+2}{N-2}} \Delta(\mu(x) c)
$$

Then it reaches the minimum of Sobolev quotient $\inf _{v \neq 0} \frac{\int_{\mathbb{S}^{N}|\nabla v|^{2}}}{\left(\int_{\mathbb{S}^{N}}|v|^{\frac{2 N}{N-2}}\right)^{2 / 2^{*}}}$, and therefore it is quite simple to prove it is the mountain pass solution, i.e. its (plain, radial, biradial) Morse index is $m(c)=1$. Now, thanks to Theorem 1.3, every other biradial solution having biradial Morse index at most 1 is constant, hence all the other solutions have biradial Morse index at least 2. Now, it is well known that Talenti's solutions are unique among positive solutions of equation (2) on $\mathbb{R}^{N}$, so we can assert that the only biradial positive solutions of (16) are constant. On the other hand, it can be proven for example using Morse Theory in ordered Banach spaces (see [2]), that the equation admits a biradial sign-changing solution having biradial Morse index at most 2. Hence there is a biradial solution of (16) with Morse index exactly 2 which is not constant.

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