

1 Climatic Influence on Sediment Distribution and Transport in the Thar Desert

2 (Sindh and Cholistan, Pakistan)

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13 Abstract

14 The Thar Desert is a major sediment depocenter in southwestern Asia and borders the
15 Indus drainage system to its east. It is unclear where the sediment that built the desert is coming
16 from, and when the desert experienced phases of construction. In particular, we seek to
17 establish the role of the South Asian monsoon in the initial formation and subsequent expansion
18 of the desert. Here we integrate bulk-petrography and heavy-mineral data with U-Pb ages of
19 detrital-zircon, to understand how the desert relates to the major potential sediment sources in
20 the Himalayan orogen and to the major rivers that surround it. Bulk petrography and heavy
21 mineral data from eolian sand in Cholistan (NE Pakistan) show close similarity with that of
22 Himalayan tributaries, whereas eolian sand in Sindh (S Pakistan) contains heavy-mineral suites
23 close to those of Indus sand largely supplied by erosion of the Karakorum and Kohistan.

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24 Kohistan is a particularly rich source of heavy minerals and is thus over-represented in
25 sediment budgets based on that proxy alone. U-Pb ages of detrital-zircon fail to show a sharp
26 difference between dune sands in Sindh and Cholistan, except for revealing somewhat greater
27 supply from the Himalaya in Cholistan and from the Karakorum, Kohistan and Nanga Parbat
28 in Sindh. Zircon ages are similar in Sindh desert sand and in the Indus Delta and are most
29 similar in deltaic sand dated as 7 ka or older. In parallel, the age signature of Cholistan sands
30 resembles more that of older river channels found along the northwestern edge of the desert
31 (e.g., paleo-Ghaggar-Hakra) than that of modern Himalayan tributaries (e.g., Sutlej). Both
32 Cholistan and Sindh sands suggest that sediment supply to the desert was greater in the early
33 Holocene when the monsoon was stronger. The southwesterly summer monsoon turned out to
34 be the most effective agent of eolian transport and recycling of Indus delta sediments entrained
35 towards the central and northern parts of the desert.

36 **Key words:** Provenance analysis; sand petrography; heavy minerals; detrital-zircon
37 geochronology; Climate change; Thar Desert.

39 1. Introduction

40 Source-to-sink studies in major siliciclastic sedimentary systems, especially those
41 supplied by rivers from major mountain belts is complicated by the buffering processes
42 involved with eolian sand seas. Large volumes of sediment may be stored and then recycled as
43 a river system passes across or nearby a sandy desert area on its way to the final depocenter.
44 Changing paleoenvironmental conditions may exercise an important control on how sediment
45 is transported across or recycled within desert areas. Changing amounts and seasonality of
46 precipitation may have an important impact on the growth or reduction of desert sand areas and
47 on their impact on the associated rivers. The fraction of sandy sediments deflated from the river

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48 channel and floodplain and stored in the sand sea is referred to as the net sediment loss from
49 the source system (Trimble, 1983; Petter et al., 2013). These stored sediments may be
50 subsequently eroded and mixed with additional fluvial sediment within the drainage basin.
51 Buffering and recycling of sediment in this way may disrupt the propagation of erosional
52 signals related to climatic change or tectonic events (Whipple, 2001; Alizai et al., 2011a; East
53 et al., 2015; Garzanti et al., 2022).

54 Our ability to resolve the factors that control fluvial-aeolian interactions is affected by
55 mixing of new and recycled sediment within the basin. In this study, we aim to better quantify
56 the processes that control sediment storage and recycling in the western Thar Desert, fed by
57 the Indus River and its Punjab tributaries draining together the Tibetan Plateau and the
58 Karakorum and western Himalayas ranges. Understanding how this sediment-filtering system
59 works represents an essential step to enhance our ability to interpret the more continuous deep-
60 marine depositional record in the Arabian Sea.

61 The efficiency of sediment supply to the Thar Desert on long timescales ($>10^4$ y) is
62 controlled by the interaction between monsoonal climate and rock uplift in the active mountain
63 belts of the Himalaya and Karakorum. This huge sedimentary system is a suitable region for
64 studying this type of source-to-sink problem, characterized by the intensity of the Southwest
65 monsoon and the strong tectonic activity of orogenic sources, which provide clear signals in
66 the sediment carried to the lower reaches of the Indus River. The Himalaya-Karakorum-Tibet
67 orogen is the result of ongoing collision between the Indian and Eurasian continental masses
68 since ~60 Ma (Critelli et al., 1994; Garzanti et al., 1996; Guillot et al., 2003; Hu et al., 2015;
69 Najman et al., 2003, 2017). Ongoing tectonically-driven rock uplift coupled with rapid incision
70 is responsible for the extreme high relief and rapid exhumation characterizing the western
71 Himalayan syntaxis drained by the Indus River (Zeitler et al., 1993; Garzanti et al., 2020a, b).

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72 If we are to understand the tectonics and erosion of this mountain belt over million-year
73 timescales, then this can only be achieved using the Indus River and its associated sedimentary
74 archive, because the record of past exhumation has been removed from the source ranges by
75 erosion. In contrast, the sediment record in the Arabian Sea and Himalayan foreland basin can
76 be used to reconstruct changing patterns and rates of erosion through provenance studies of the
77 detrital sediment flux from the orogenic region (Clift et al., 2001; Clift et al., 2008; Clift and
78 Jonell, 2021). In recent times, all sediment flowing from the various sources into the ocean
79 basin has passed next to or through the Thar Desert, but the processes that operate in this region
80 are not well studied and poorly understood.

81 Studies of sediments in arid and semi-arid regions seldom pay sufficient attention to
82 the interactions that occur during fluvial and aeolian transport (e.g., Thomas and Wiggs, 2008;
83 Belnap et al., 2011; Feder et al. 2018; Liang et al., 2023). These are likely to introduce
84 complexities to the original erosional pulse that might be linked to environmental change, and
85 whose compositional signature closely reflects the eroded source rocks. In the case of the Indus
86 River system, such a change mainly occurs in the Thar Desert, extending across SE Pakistan
87 and NW India (Figs. 1 and 2). In the Thar Desert, the complex nature of sediment transport and
88 recycling has not so far been adequately investigated, even though a better understanding of
89 these processes are essential to constrain the ongoing fluvial-eolian interactions during source-
90 to-sink transport. The sediments stored in the desert reflect the net effects of sediment transport
91 from hillslope to alluvial channels and finally to the dryland area (Ramsey et al., 1999; Prins
92 et al., 2009; Bracken et al., 2015). This transport may be modulated by changes in the
93 environment, as well as in the development of the drainage system that may be particularly
94 influenced by tectonic activity. The arid-semi arid Thar Desert is quite sensitive to fluvial
95 activity and changing climate. The fluvially supplied eolian sands are stored in the dune field
96 and reworked by winds. Subsequently, the sediments may interact further with the river and

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7 97 continue their journey to the ocean (Bullard and Livingstone, 2002; Bullard and McTainsh,
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9 98 2003; Alizai et al., 2011b). Wind is a strong transforming agent that can move sediments
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11 99 hundreds of kilometers from the original fluvial source (Langford, 1989; Garzanti et al., 2017)
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13 100 until they are eventually stored in dune fields such as the Thar Desert.

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15 101 Climate change affects summer-monsoon winds, which move sediment from southwest to
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17 102 northeast across the Thar Desert (Wasson et al., 1983; Singhvi and Kar, 2004; East et al., 2015).
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19 103 This summer transport path is reversed in winter, but this system only plays a marginal role in
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21 104 sand transport, whereas dust storms are associated with early summer monsoon rain in the
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23 105 southwestern Indian subcontinent (Singhvi et al., 2010). The principal aim of the present study
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25 106 is to improve our understanding of the influence that climate change exerts on sediment
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27 107 transport towards and across the Thar Desert by using bulk-petrography, heavy-minerals, and
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29 108 detrital-zircon U-Pb age analyses to trace sediment transport pathways. We reconstruct how
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31 109 sediments were fed into and distributed within the Thar Desert in recent times and evaluate the
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33 110 impact of Holocene environmental and climatic changes on sediment generation in the western
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35 111 part of the Himalayan orogen.

36 37 38 112 39 40 41 113 **2. Thar Desert**

42 43 44 114 **2.1. Geomorphological framework**

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46 115 The term “desert” passed from the Egyptian hieroglyphics *Tésert* to the Latin word
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48 116 *desertum* (place of abandoned wilderness) to describe rainless dune field areas (El-Baz, 1983).
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50 117 El-Baz (1983) argued that a desert is “a vast area of dune fields and arid land where wind
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52 118 *blows dust in the air and moves sand grains across dune profiles*”. Dune sediments are easily
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54 119 eroded and deflated by wind action. Sand derived from mountain sources is thus affected by
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56 120 repeated reworking and recycling during fluvial-eolian interactions.
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121 The Thar Desert, situated in the western Indian subcontinent in an arid environment of
122 sandy rolling hills (locally named “*thul*”), covers ~200,000 km² including the Punjab and Sindh
123 provinces in Pakistan, and the Gujarat and Rajasthan states in India. Sand dunes stretch from
124 north to south for ~800 km and merge in the west into the fertile alluvial plains of the Indus
125 River in Pakistan. The southern side of the desert is bound by the foothills of the NNE to SSW-
126 trending Proterozoic Aravalli ranges in Gujarat (Kar, 2013). The Pakistani part of the Thar
127 Desert includes the Cholistan Desert in the north (Bahawalpur district, Punjab) and the Sindh
128 Desert in the south (Sindh Province). Thar Dunes are remarkable for their constant movement
129 by wind erosion and deposition (Sam et al., 2015), which prevents permanent human settlement
130 (Nordstrom, 2014).

131 The Cholistan Desert, ~480 km in length and between ~23 and ~192 km in width,
132 covers an area of ~26,000 km² (Figs. 1 and 2). Two regions are distinguished based on
133 geomorphological differences, Lesser Cholistan (12,370 km²) and Greater Cholistan (13,960
134 km²) (Akhtar and Arshad, 2006; Ahmed, 2011). Lesser Cholistan at the northern edge of the
135 desert comprises low dune ridges and alluvial flats (Akbar et al., 1996; Mughal, 1997; Mughal
136 et al., 2016) and lies around the Ghaggar-Hakra paleochannel which is a now abandoned
137 Himalayan-sourced Punjabi tributary flowing closest to the desert and that contains remnants
138 of sandy fluvial terraces (Ahmad, 2008) and may have formerly joined the Nara River to the
139 south of Cholistan (Clift et al., 2012; Srivastava et al., 2020). Greater Cholistan in the south is
140 dominated by sand dunes. The Sindh Desert, ~500 km in length and between ~38 km and ~200
141 km in width, covers an area of ~45,790 km² characterized by undulating sand ridges (Chauhan,
142 2003) (Figs. 1 and 2).

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2.2. Thar Desert climate and the Monsoon

The geographic position of the Thar Desert, surrounded by mountain ranges, rivers and alluvial plains, contributes significantly to the weather patterns that shape its distinctive, hot and dry environment. Seasonal winds from the Indian Ocean lose moisture as they approach the Himalayan front where most of the precipitation is released, and a low-pressure system develops adjacent to the mountains in the summer by trapping hot air over the plains (Wang et al., 2005; Boos and Kuang, 2010). The Thar Desert region, which lies on the western edge of this circulation system and is also affected by dry westerly winds from the arid Makran Desert and Afghanistan, has consequently very low precipitation throughout the year (100–200 mm/a; Akbar et al., 1996; Rahaman et al., 2009). Furthermore, the Thar Desert lies in the rain shadow of the Aravalli Range, which contributes to its aridity, with high temperatures that cause strong evaporation (Kumari et al., 2023).

All rain carried by monsoon clouds is absorbed by surrounding regions, so that monsoon winds in the Thar Desert are hot and dry. Nevertheless, the Asian Monsoon remains crucial to the Thar Desert region, delivering 90% of rainfall in Cholistan (100–500 mm/a) during the summer months (May to September, Bookhagen, 2010). The weaker winter monsoon provides little additional precipitation, but some rainfall in the northern part of the desert is associated with winter westerlies. The Thar Desert in Sindh receives ~100 mm of rain in the west and ~500 mm in the east, with significant variations from year to year. The hottest months (temperatures up to 50°C) are May and June, whereas the coldest month is January when frost is frequent (mean minimum temperature 5°–10°C). Dust storms and dust-raising winds with velocities up to 145 km/h are common in May to June.

The Asian Monsoon responds linearly to orbital forcing during obliquity (41 k.y.) and precession (23 k.y.) cycles (Prell, 1984; Clemens et al., 1996; Clemens and Prell, 2003).

169 Monsoon intensity is affected by changes in the volume of Arctic and Antarctic icecaps, and
170 by temperatures in both northern and southern hemisphere resulting from cross-equatorial
171 exchanges of atmospheric heat and pressure (Blinkhorn, 2014). Maximum monsoonal intensity
172 is observed during glacial minima, whereas a steep inter-polar temperature gradient commonly
173 produces a monsoonal minimum ~20 k.y. before the glacial maximum (Zhisheng et al., 2011).
174 Paleoenvironmental data suggest that the Thar Desert may have experienced reduced
175 monsoonal rainfall between ~40 ka and ~20 ka.

2.3. *Paleochannels and sediment flux*

178 The western Himalayan foreland flood plains have been re-incised by the Indus River
179 since ~10 ka, following earlier aggradation during deglaciation (Giosan et al., 2012). Fluvial
180 discharge was mainly controlled by regional monsoonal precipitation, as well as by melting of
181 snow and ice in the Himalayan-Karakorum headwaters (Bookhagen, 2010; Durcan et al.,
182 2010). Several major tributaries supply sediment from orogenic sources to the Indus lower
183 reaches, including the Shyok, Shigar, Gilgit and Kabul tributaries of the upper Indus and the
184 five Punjab (*punj*, five; *ab*, water) tributaries (Jhelum, Chenab, Ravi, Beas, and Sutlej) of the
185 lower Indus (Fig. 1).

186 The mainstream of the Indus River rises from glaciers north of Mt. Kailas and flows for
187 a third of its course along the Indus suture zone before cutting southwards across the Western
188 Himalayan Syntaxis, eventually meeting the Himalayan front where it is now barred by the
189 Tarbela Dam (Fig. 1). The Punjab tributaries cut across the entire Himalayan belt and account
190 for ~44% of total Indus discharge (Alizai et al., 2011a).

191 The Indus River has undergone several drainage changes through time. Western migration
192 of the river courses has occurred in Sindh since the Last Glacial Maximum (LGM) (Kazmi,

193 1984). Other major recent changes include the loss of the Yamuna River, captured by the Ganga
194 River between 10 and 50 ka, and the drying up of the Ghaggar-Hakra tributary (Valdiya, 2002;
195 Saini et al., 2009; Clift et al. 2012; Giosan et al., 2012). The Punjabi tributaries have also
196 experienced repeated changes (Amundson et al., 1986; Wright et al., 2005): the Sutlej River
197 lost contributions from the Yamuna River before 17 ka (Stein, 1942; Sinha et al., 2013; Sinha
198 et al., 2019) and the Ravi drainage moved northwards in the middle Holocene.

199 Tentative estimates indicated that ~11% of the sediment reaching the Arabian Sea in the
200 Holocene was supplied by incision into the northern floodplains, and another ~6% by recycling
201 of Thar Desert sands (Clift and Jonell, 2021). Reworking of Thar Desert sediment occurred
202 after 10 ka, when the Indus and its tributaries incised into the lower reaches of the western
203 Punjabi floodplains (Giosan et al., 2012). Sediment deposited in the Thar Desert date from at
204 least mid-Pleistocene time (Glennie et al., 2002; Singhvi et al., 2010) and 173 ka-old fluvial
205 channels have been identified in the central part of the Thar Desert (Blinkhorn et al., 2020).
206 More recent fluvial channels around the western edge of the desert were abandoned at ~4–5 ka
207 and were buried under parabolic dunes at ~1.4 ka (Clift et al., 2012). Younger intermittent
208 fluvial activity at 2.9–0.7 ka has also been reported from both the upper and lower Ghaggar-
209 Hakra floodplain (Giosan et al., 2012). Evidence from Pb-isotope studies of detrital K-feldspar
210 (Alizai et al., 2011b) and chemistry of detrital garnet (Alizai et al., 2016) in the Nara Valley
211 suggested that the Nara River was either a former course of the river Indus or an ephemeral
212 stream entirely fed with reworked sediments.

214 **2.4. Climatic changes**

215 Because climatic changes have a strong influence on sediment generation and fluvial transport
216 (Haughton, 1991; Gábris, and Nádor, 2007), the study of alluvial deposits offers a key to
217 paleoclimatic reconstructions (Van de kamp and Leake, 1985; Velbel and Saad, 1991). The

1 218 present arid climate of the Thar Desert was established at ~4 ka, after a wet middle Holocene
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3 219 stage preceded by dry conditions in the early Holocene (Dhir and Singhvi, 2012).
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5 220 Luminescence ages indicate relatively inefficient aeolian sediment transport during the LGM
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7 221 (Singhvi and Kar, 2004; Singhvi et al., 2022). In contrast, stronger summer monsoonal winds
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9 222 increased aeolian transport during the Holocene, where sediment supplied during times of
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11 223 heavier rain was blown into the central part of the Thar Desert, where it rapidly accumulated
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13 224 throughout the Holocene (Singhvi et al., 2010; Chatterjee and Ray, 2017). Dunes along the
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15 225 western edge of the desert have advanced over the floodplains only more recently, and the
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17 226 connectivity between the Indus River and Thar dune fields has been documented by
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19 227 stratigraphic studies indicating eolian sediment transport mostly from the southwest (East et
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21 228 al., 2015). This general pattern is consistent with sediment mineralogy, and optically stimulated
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23 229 luminescence (OSL) and thermoluminescence (TL) ages (Singhvi et al., 2010). The Thar
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25 230 Desert expanded after the LGM, during phases of strengthened summer monsoon, and
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27 231 continued to advance westward during weakening of the summer monsoon starting after 8 ka.
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29 232 Pb isotope compositions of detrital K-feldspar grains, along with U-Pb zircon and Nd isotopic
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31 233 data, highlight a similarity between the lower reaches of the Indus River and the northern
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33 234 Cholistan Desert (Alizai et al., 2011a, 2011b).
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44 236 **3. Geological setting of the Thar Desert**

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47 237 Quaternary sediments of the Thar Desert, interspersed with low hills where Cenozoic rocks
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49 238 are locally exposed, are underlain by Archean gneiss, non-conformably covered by Proterozoic
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51 239 sedimentary rocks (~2.5 Ga to 541 Ma) and more recent alluvium (Roy and Smykatz-Kloss,
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53 240 2007; Jain and Banerjee, 2020). Desert sediments of original fluvial origin were repeatedly
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55 241 deflated by wind into sand ridges, even including clay locally. Sand in the northern Cholistan
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57 242 Desert is largely derived from the Himalaya and supplied from the Sutlej River (Ahmed, 2008).
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243 Sand in the southern Sindh Desert, instead, is mainly sourced from the Indus River largely
244 sourced from the Karakorum Range and the Kohistan Arc, as well as the Himalaya and
245 representing the main source for sediments fed into the Arabian Sea (Clift et al., 2002; Garzanti
246 et al., 2020a)

247 The main potential sediment sources to the Indus River and Thar Desert are the Karakorum
248 and Hindukush, the Western Himalayan Syntaxis (WHS, i.e., Nanga Parbat Massif),
249 Transhimalayan arcs (Henderson et al., 2010), the NW Greater and Lesser Himalaya, and
250 subordinately the ranges bounding Pakistan to the west, i.e., Kirthar and Sulaiman ranges
251 (Guillot et al., 2003; Garzanti, 2019a). The Kohistan and Ladakh arcs are bound by the Shyok
252 and Indus Sutures to their north and south (Debon et al., 1987; Garzanti and Haver, 1988;
253 Treloar et al., 1996; Rolland et al., 2000). The Nanga Parbat Massif is cut by the Indus River
254 but supplies a notably lower amount of sediment than the eastern syntaxis despite its rapid
255 exhumation (Clift et al., 2022). The main focus of erosion are the Karakorum domes
256 (Lemennicier et al., 1996; Crawford and Searle, 1992; Searle et al., 1999; Rolland et al., 2001;
257 Searle et al., 2010) cut by the Hushe, Braldu and Hispar rivers (Garzanti et al, 2020b).

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259 **4. Fingerprinting sand sources**

260 Sand in the Thar and Thal deserts is supplied by the Upper Indus that drains the Ladakh
261 and Kohistan arcs, the Karakorum and Hindukush ranges, the Nanga Parbat Massif, and the
262 Punjabi tributaries (Jhelum, Chenab, Ravi, Beas, and Sutlej) draining the Himalaya (Liang et
263 al., 2019; Garzanti et al., 2020b).

264 The Ladakh Arc supplies quartzo-feldspathic to feldspar-rich feldspatho-quartzose,
265 plutoniclastic sand with a rich to very rich transparent heavy mineral (tHM) suite dominated

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266 by amphibole (mostly hornblende) with epidote and minor clinopyroxene (mainly diopside)
267 and titanite.

268 The Kohistan Arc supplies feldspatho-quartzo-lithic to litho-quartzo-feldspathic
269 metamorphiclastic sand with common prasinite and epidote-amphibolite grains ([Garzanti et al.,](#)
270 [2005](#)). The very rich to extremely rich tHM suite is dominated by amphibole (mainly
271 hornblende or pargasite associated with actinolite or hastingsite) with common epidote-group
272 minerals (mostly clinozoisite) and pyroxenes (diopside, pigeonite, augite, and hypersthene).

273 The Karakorum supplies quartzo-feldspatho-lithic (North Karakorum) to quartzo-
274 feldspathic plutoniclastic (Central Karakorum), and litho-feldspatho-quartzose
275 metamorphiclastic sand with marble grains (South Karakorum). Mainly moderately rich tHM
276 suites include common amphibole (mainly hornblende with pargasite, hastingsite, or
277 actinolite), epidote-group minerals (epidote, clinozoisite, and allanite), garnet mostly derived
278 from intermediate-acidic igneous rocks (Bi-type, following [Mange and Morton, 2007](#)), titanite,
279 common diopsidic clinopyroxene, and minor kyanite, staurolite, and sillimanite.

280 The Zanskar River draining the northern side of the Greater Himalaya which contribute
281 litho-feldspatho-quartzose metamorphiclastic sand, with a moderately rich tHM suite including
282 amphibole (pargasite and hornblende with minor hastingsite), mostly Bi-type garnet, fibrolitic
283 sillimanite, kyanite, epidote-group minerals (epidote, clinozoisite, minor allanite), and
284 pyroxenes (diopside, augite, and locally hypersthene). The Nanga Parbat Massif contributes
285 feldspar-rich feldspatho-quartzose sand with a very rich tHM suite dominated by amphibole
286 (mainly hornblende with common tschermakite and minor pargasite) with clinopyroxene
287 (diopside with rare augite), garnet mainly derived from high-grade metabasic rocks (Ci-type)
288 and minor from amphibolite-facies metasediments (Bii-type) ([Mange and Morton, 2007](#)),
289 epidote, clinozoisite, and sillimanite.

290 Cenozoic foreland basin sedimentary strata are incised by the Soan River and shed
291 feldspatho-litho-quartzose sand with a moderately rich, epidote-dominated tHM suite with
292 garnet, hornblende, and tourmaline.

293 5. Methods

294 For this study we collected 50 sand samples from the Sindh province in the southern
295 Thar Desert and 40 sand samples from the Cholistan province in the northern Thar Desert ([Fig. 1](#)). Complete information on sampling sites is given in [Appendix A Table A1](#) and in the Google Earth™ file [Thar.kmz](#). The complete petrographic, heavy-mineral, and geochronological datasets are provided in [Appendix A](#) and [B](#).

299 5.1. Sand petrography

300 A quartered fraction of each sand sample was impregnated with araldite epoxy and cut
301 into a standard thin section stained with alizarine red to distinguish dolomite and calcite.
302 Petrographic analysis was carried out by counting ~450 points under the microscope following
303 the Gazzi-Dickinson method ([Ingersoll et al., 1984](#)). Sand classification was based on the
304 relative abundance of the three main framework components quartz (Q), feldspars (F) and lithic
305 fragments (L), considered if exceeding 10%QFL (e.g., a sand is named feldspatho-litho-
306 quartzose if $Q > L > F > 10\%QFL$ or quartzo-lithic if $L > Q > 10\%QFL > F$; classification
307 scheme after [Garzanti, 2019b](#)). Metamorphic grains were classified according to their protolith
308 and metamorphic rank, expressed by the metamorphic indices MI and MI* ([Garzanti and Vezzoli, 2003](#)). MI varies from 0 (detritus shed by sedimentary and volcanic cover rocks
309 exclusively) to 500 (very-high-rank detritus exclusively shed by high-grade basement rocks).
310 MI* considers only metamorphic rock fragments and thus varies from 100 (very-low-rank
311 detritus shed by very low-grade metamorphic rocks) to 500. Petrographic parameters used in
312 this article include the plagioclase/total feldspar (P/F) ratio; feldspar with cross-hatch twinning
313 is referred to as microcline. Median grain size was determined in thin section by ranking and

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317 visual comparison with in-house standards of sieved $\Phi/4$ classes. Significant detrital
318 components are listed in order of abundance (high to low) throughout the text. Key
319 compositional parameters are summarized in [Table 1](#) and the complete dataset is provided
320 [Appendix A, Table A2](#).

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322 **5.2. Heavy minerals**

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324 From a split aliquot of the 15–500 μm size window obtained by wet sieving, heavy
325 minerals were separated by centrifuging in Na-polytungstate (2.90 g/cm^3), and these were then
326 recovered by partial freezing with liquid nitrogen (Andò, 2020). For each sample, ~250
327 transparent heavy minerals (or all of those present in the grain mount) were point-counted at
328 appropriate regular spacing to minimize overestimation of smaller grains. Transparent heavy-
329 mineral assemblages, called for brevity “tHM suites” throughout the text, do not include
330 alterites, phyllosilicates and carbonates (Garzanti and Andò, 2019). According to the
331 transparent-heavy-mineral concentration in each sample (tHMC), tHM suites are defined as
332 extremely poor ($\text{tHMC} < 0.1$), very poor ($0.1 \leq \text{tHMC} < 0.5$), poor ($0.5 \leq \text{tHMC} < 1$),
333 moderately poor ($1 \leq \text{tHMC} < 2$), moderately rich ($2 \leq \text{tHMC} < 5$), rich ($5 \leq \text{tHMC} < 10$), or
334 very rich ($\text{tHMC} > 10$). The sum of zircon, tourmaline, and rutile over total transparent heavy
335 minerals (ZTR index of Hubert, 1962) measures the relative proportion of durable minerals in
336 the tHM suite and can thus be considered as an index of recycling (Garzanti, 2017). Key heavy-
337 mineral parameters are summarized in [Table 1](#) and the complete dataset is provided [Appendix](#)
338 [A, Table A3](#).

340 **5.3. Detrital-zircon geochronology**

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342 Starting from the heavy-mineral separates of five samples (two from the Cholistan
343 Desert and three from the Sindh Desert), zircon grains were concentrated with standard

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344 magnetic techniques, directly mounted in epoxy resin without any operator selection by hand
345 picking, and identified by automated phase mapping (Vermeesch et al., 2017) under a
346 Renishaw inViaTM Raman microscope. U-Pb zircon ages were determined at the London
347 Geochronology Centre using an Agilent 77003 LA-ICP-MS system, employing a NewWave
348 NWR193 Excimer Laser operated at 10 Hz with a 25-lm spot size and ~ 2.5 J/cm² fluence. No
349 cathodoluminescence imaging was done, and the laser spot was always placed blindly in the
350 middle of zircon grains to treat all samples equally and avoid bias in inter sample comparison
351 ('blind-dating strategy' as discussed in Garzanti et al., 2018). The mass spectrometer data were
352 converted to isotopic ratios using GLITTER 4.4.2 software (Griffin et al., 2008), employing
353 Plešovice zircon (Sláma et al., 2008) as a primary age standard and GJ-1 (Jackson et al., 2004)
354 and 91500 (Wiedenbeck et al., 1995) as a secondary age standard (Appendix B, Figs. S1 and
355 S2). A NIST SRM612 glass was used as a compositional standard for U and Th concentrations.
356 GLITTER files were post-processed using IsoplotR (Vermeesch et al., 2018). Concordia ages
357 were calculated as the maximum likelihood intersection between the concordia line and the
358 error ellipse of ²⁰⁷Pb/²³⁵U and ²⁰⁶Pb/²³⁸U (Ludwig, 1998). The discordance cutoff was set at –
359 5/+15 of the concordia distance (Vermeesch, 2021). Concordant ages are provided in Appendix
360 A Table A4.

362 *5.4. Graphical/statistical tools*

363 To visualize heavy-mineral data we use the compositional biplot (Gabriel, 1971), drawn
364 using CoDaPack software (Comas-Cufí and Thió-Henestrosa, 2011). The biplot allows
365 discrimination between multivariate observations (points) while shedding light on the mutual
366 relationships among variables (rays). The length of each ray is proportional to the variance of
367 the corresponding variable in the dataset. If the angle between two rays is close to 0°, 90° or

180°, then the corresponding variables are correlated, uncorrelated, or inversely correlated, respectively.

Statistical tools applied to U-Pb detrital-zircon age populations include multidimensional scaling (MDS; [Kruskal and Wish, 1978](#); [Vermeesch, 2013](#)). MDS produces a map of points in which the distance between samples is approximately proportional to the Kolmogorov-Smirnov dissimilarity of their compositional or chronological signatures. Closest and second closest neighbors are linked by solid and dashed lines, respectively, and the goodness of fit is evaluated using the “stress” value of the configuration (0.2 = poor; 0.1 = fair; 0.05 = good; [Vermeesch, 2013, 2018](#)).

6. Compositional fingerprints and provenance of Thar Desert sands

6.1. Petrography and Heavy minerals

Cholistan Desert dune sand is feldspatho-litho-quartzose with dominant monocrystalline quartz and subequal plagioclase and K-feldspar ([Figs. 3 and 4](#)). Quartz content decreases southward from ~70% to ~57% ([Table 1](#)). The varied rock-fragment population consists of metasedimentary (paragneiss, schist, slate, calcschist, phyllite, metasandstone), metabasite (prasinite, chloritoschist, amphibolite), carbonate (limestone, dolostone), other sedimentary (shale, siltstone, minor chert), granitoid, felsic to mafic volcanic and metavolcanic, and minor ultramafic (serpentine schist, cellular serpentinite) grains (MI 256–326, MI* 239–314) ([Figs. 3 and 4](#)). Metamorphic and sedimentary lithics and the MI index tend to decrease SE-ward. Muscovite and biotite flakes are observed. The rich tHM suite includes hornblende, subordinate epidote and garnet, and minor clinopyroxene, hypersthene, staurolite, titanite, kyanite, and fibrolitic sillimanite ($ZTR \leq 4$).

Sindh Desert dune sands are litho-felshpatho-quartzose, with plagioclase \geq K-feldspar (P/F 49-60) ([Figs. 3 and 4](#)). Sedimentary rock fragments prevail over metapelite,

393 metapsammite, and metavolcanic grains (MI 263–220). The rich tHM suite consists of
394 blue/green to brown hornblende, subordinate garnet and epidote, minor clinopyroxene and
395 hypersthene, sillimanite, kyanite apatite, and titanite ($ZTR \leq 3$).

396 Sindh sand has less quartz than that from Cholistan (53 ± 2 vs. 60 ± 2), more sedimentary
397 than metamorphic lithics (Lm 39 ± 6 Ls 58 ± 5 vs. Lm 62 ± 2 Ls 37 ± 2), and higher heavy-mineral
398 concentrations with overall similar amphibole-epidote-garnet tHM population typical of
399 orogenic sediments (Garzanti and Andò, 2007; Fig. 4). This indicates a strong affinity to Indus
400 River sand, especially in the southern part of the Sindh Desert, whereas samples from the
401 Cholistan Desert have a greater affinity with sand transported by Punjabi tributaries sourced in
402 the Himalaya.

403

404 **6.2. Detrital-zircon geochronology**

405 Data obtained in this study on dune sands from Cholistan (191 ages from 2 samples) and
406 Sindh (330 ages from 3 samples), together with literature data from another three samples (East
407 et al., 2015), compose a data set of 841 concordant ages overall (Table 2).

408 Zircon grains from Cholistan dunes yielded mainly Paleoproterozoic (26%) and
409 Neoproterozoic (25%) ages (Fig. 5), with minor Neoproterozoic (2%), Mesoproterozoic (5%),
410 Paleozoic (8%), Triassic (1%), Early Cretaceous (10%), Late Cretaceous (5%), Paleocene
411 (5%), Eocene-Oligocene (9%) and Miocene (2%) ages (Appendix A Table A5).

412 Zircon grains from Sindh dunes are mainly Neoproterozoic ages (28%) and
413 Paleoproterozoic (24%) ages, with significant Ordovician, Silurian, and Carboniferous ages
414 (14% overall). Neoproterozoic (1%), Mesoproterozoic (7%), Triassic (1%), Jurassic (1%), Early
415 Cretaceous (9%), Late Cretaceous (3%), Paleocene (6%), Eocene-Oligocene (4%), and
416 Miocene (3%) ages also occur (Appendix A Table A5).

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417 The youngest grains found both in Cholistan (13–21 Ma) and Sindh (17–25 Ma) are
418 only known elsewhere in sand from the Hushe River draining the Baltoro Granite, and thus
419 point to a provenance from the Karakorum Range (MDS, Fig. 6). Cretaceous to Oligocene
420 grains are inferred to be predominantly derived from the Karakorum (characteristic peaks at
421 99–130 Ma and 24–43 Ma) and Transhimalayan arcs (43–96 Ma). Paleozoic and
422 Neoproterozoic grains, most abundant in Sindh, are contributed by both Karakorum and
423 Himalayan sources, whereas Mesoproterozoic to Paleoproterozoic (equally abundant in
424 Cholistan and Sindh) are largely derived from the Greater Himalaya, Lesser Himalaya, and/or
425 the Nanga Parbat Massif (Figs. 6 and 7).

427 7. Provenance budgets

428 The relative contributions from each geological domain to the aeolian sand of the Thar
429 Desert can be estimated by forward mixing models (Garzanti et al., 2012; Resentini et al.,
430 2017). Calculations, however, are nonunique and uncertain, influenced by potentially
431 inaccurate compositional information on end-member sources and unverified assumptions. To
432 ensure robust results, independent tests must be conducted using different criteria (Garzanti et
433 al., 2020).

434 7.1. Petrography and Heavy minerals

435 Forward mixing calculations indicate that Indus River sand upstream of Tarbela Dam is
436 derived ~60% from the Karakorum, ~20% from Transhimalayan arcs, ~13% from the Nanga
437 Parbat massif, and ~6% from the Tethyan and Greater Himalaya (Garzanti et al., 2005; 2020)
438 (Fig. 8a). These proportions change notably in the Lower Indus River, where sand is calculated
439 to be derived 27±3% from the Karakorum, 10±3% from the Hindukush, 3±2% from the Ladakh
440 Arc and Tibetan Plateau, 7±2% from the Kohistan Arc, 6±3% from the Nanga Parbat Massif,

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5 441 and 39±4% from the Himalaya, delivered largely by Punjab tributaries (Garzanti et al., 2005)
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7 442 (Fig. 8b).

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10 443 Based on the integrated petrographic-heavy mineral dataset, dune sand in Cholistan is
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12 444 calculated to be derived ~30% from the Ladakh-Kohistan arcs, ~5% from the Karakorum–
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14 445 Hindukush ranges, ~17% from the Nanga Parbat Massif, and ~48% from Himalayan units (Fig.
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16 446 8d). Instead, dune sand in Sindh is calculated to be derived ~34% from the Ladakh-Kohistan
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18 447 arcs, ~31% from the Karakorum–Hindukush ranges, ~16% from the Nanga Parbat Massif, and
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20 448 ~19% from Himalayan units (Fig. 8c).

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22 449 Calculations based on petrographic data suggest that Cholistan sand dunes are mostly (up
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24 450 to ~90%) derived from Punjabi tributaries and dominantly from the Sutlej River flowing along
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26 451 the northern border of the Thar Desert (Fig. 8e). Calculations based on heavy-mineral data,
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28 452 however, indicate a major contribution (up to two thirds) from the Indus River (Fig. 8f). Some
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30 453 sediment may have been supplied also by the Ghaggar-Hakra River, formerly potentially
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32 454 connected with the Nara River in Sindh and for which we could not determine an end-member
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34 455 composition.

35 36 456 **7.2. Detrital zircon ages**

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39 457 A set of forward mixing calculations based on the age populations illustrated above were
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41 458 performed using the *DZMix* software of Sundell and Saylor (2017). In each case, 10,000
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43 459 attempts were made to match the observed zircon-age spectrum and the best 1% were selected.
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46 460 The best fit was separately assessed for Cholistan and Sindh samples using three statistical
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48 461 approaches (i.e., cross correlation, Kuiper test, and K-S test). End-members for source regions
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51 462 were characterized using detrital-zircon data from modern river sediments rather than bedrock
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53 463 data, because only the former provide a more representative weighed regional average of
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56 464 source-rock compositions. We use data from Zhuang et al. (2018) and Clift et al. (2022) for the
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58 465 Karakorum and Hindukush ranges, from Garzanti et al. (2020b) and Clift et al. (2022) for the

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466 Nanga Parbat Massif, from [Zhuang et al. \(2018\)](#) and [Garzanti et al. \(2020b\)](#) for the Kohistan
467 Arc, from [Garzanti et al. \(2020b\)](#) and [Clift et al. \(2022\)](#) for the Ladakh Arc, from [Jonell et al.](#)
468 [\(2017\)](#) and [Garzanti et al. \(2020\)](#) for the Greater Himalaya, and from [Alizai et al. \(2011b\)](#) for
469 Punjabi tributaries sourced in the Himalaya, separately considering data from the Sutlej and
470 Beas Rivers that flow closest to the Cholistan Desert. As well as looking at modern Thar desert
471 dunes, we also applied the same mixing calculations to the Upper Indus at Attock bridge ([Alizai](#)
472 [et al., 2011b](#); [Garzanti et al., 2020b](#); [Clift et al., 2022](#)), and the modern lowermost Indus River
473 at Thatta ([Clift et al., 2004](#)). We also considered data from post-LGM sediments of the Indus
474 Delta and from older Cholistan sediments cored in the Marot borehole [Clift et al. \(2012\)](#).

475 The results of forward mixing calculations, provided in [Appendix A Table A6](#) and in [Figure](#)
476 [9a-h](#), with the latter showing the preferred contributions based on the V factor of the Kuiper
477 test, indicate that zircons in Cholistan sand are mostly derived from Punjabi tributaries (~69%),
478 with the remaining ~31% supplied by the Upper Indus sources (i.e., Karakorum-Hindukush,
479 Ladakh-Kohistan, and Nanga Parbat). In contrast, zircons in Sindh sand are derived more from
480 Upper Indus sources (~54%; [Fig. 9a](#)) than from Punjabi tributaries (~46%), with notably higher
481 contributions from the Kohistan Arc (~10% vs. only ~2% in Cholistan) and Nanga Parbat
482 Massif (~8% vs. ~3% in Cholistan) ([Table 3](#); [Fig. 9b](#)).

483 [East et al. \(2015\)](#) argued that eolian sand in the southern Thar Desert was largely blown
484 from the Indus Delta. The comparison with detrital zircons from modern Lower Indus River
485 sand ([Clift et al., 2004](#); [Fig. 9c](#)) indicates a greater supply from Punjabi tributaries (~38% vs.
486 ~17% for Sindh Desert sand) at the expense of Karakorum and Kohistan. Detrital zircons from
487 Indus deltaic sand dated as ~7 ka (sample TH-10 in [Clift et al., 2008](#); [Fig. 9d, e](#)), however,
488 appear to be dominantly derived from the Karakoram (~65%), implying drastic provenance
489 changes in the Lower Indus occurred after 7 ka. Zircon ages from sand deposited at ~14 ka
490 (sample KB-40 in [Clift et al., 2008](#); [Fig. 9f](#)) more closely match zircon ages in Sindh sand, but

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5 491 with an inferred greater contribution from Ladakh-Kohistan arcs (~27% vs. ~10%) and a much
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7 492 lower contribution from Punjabi tributaries (~16% vs. ~46%).

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10 493 The comparison between zircon ages in Cholistan Desert and modern Lower Indus River
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12 494 sands indicate a similar contribution from Punjabi tributaries. Cholistan dunes have a similar
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14 495 signature as Marot borehole samples 6 and 12 (Clift et al., 2012; Fig. 9g, h), although a more
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16 496 important zircon supply from Punjabi tributaries is indicated for the Sutlej River for Marot-6
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18 497 dated as ~7 ka, and for the Beas River for Marot-12 dated as > 49 ka (Table 3).

19 499 *7.3. Summary of provenance results*

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22 500 The combination of petrographic, heavy-mineral, and detrital-geochronology methods
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24 501 allow us to put firm constraints on the provenance of Thar Desert sands, indicating that eolian
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26 502 dunes in the northern Cholistan area were largely fed by Punjabi tributaries sourced in the
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28 503 Himalayan belt (primarily the Sutlej-Beas River system, with potential additional contributions
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30 504 from the now extinct Ghaggar-Hakra paleo-river). Instead, eolian dunes in the southern Sindh
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32 505 area have more similar composition as Indus Delta sands and were largely supplied by the
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34 506 Lower Indus River.

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39 507 The concentration of heavy minerals in the sediment varies depending on the supply from
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41 508 contrasting source units, and thus varies between different tributaries with unique bedrock
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43 509 exposures in their headwaters. There are examples where sediments appear to have been
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45 510 partially eroded from source areas that are not present upstream of the sampling location in the
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47 511 adjacent river system. For example, in northern Cholistan it might not be expected to see
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49 512 sediment eroded from Nanga Parbat, Ladakh-Kohistan or the Karakorum. Although the
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51 513 percentage of heavy minerals from those areas is less than other parts of the desert the values
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53 514 are not zero and require recycling via aeolian transport from the south. It should also be
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55 515 remembered that different source areas may have contrasting degrees of fertility in the different
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516 heavy minerals, so that the proportion of sediment derived from each source and calculated
517 from heavy mineral percentages may not be representative of the total mass flux. Hydraulic
518 sorting is a crucial factor for the concentration and variety of heavy minerals from which we
519 derive provenance information. Compared to the modern Upper Indus and Thal Desert sands,
520 Cholistan Desert sand is notably poorer in quartz and sedimentary to low-rank
521 metasedimentary rock fragments. Cholistan sediments are however richer in feldspars,
522 volcanic, metavolcanic and metabasic rock fragments, heavy minerals and especially
523 hypersthene, documenting a significantly greater relative contribution from the Kohistan arc.

524 We can explore the provenance of the desert dunes further using the *DZStats* program from
525 [Saylor and Sundell \(2016\)](#) ([Appendix A Table A7](#)). In this approach, a series of samples are
526 directly compared with one another using a statistical test, in this case the Kuiper Test to
527 quantify their similarity. We compare the Sindh and Cholistan dune averages with potential
528 sources, i.e., the Upper Indus at Attock, the modern delta at Thatta, two post-LGM delta
529 samples (TH-10 at ~7 ka and KB-40 at ~14 ka), together with the modern Sutlej and two older
530 fluvial channel sands from Marot (Marot-6 at ~7 ka and Marot-12 >49 ka). These sands are
531 from the Ghaggar-Hakra. There is no direct measurement of what the Sutlej or other Punjabi
532 tributaries looked like during the Mid Holocene or older times. The results are shown in [Figure](#)
533 [10 and Table 4](#). No strict affinity, however, is observed between the dune fields and either
534 modern Upper Indus or Sutlej River sands. The compositional signatures of Sindh dunes are
535 closest to deltaic sediments dated between ~7 and 14 ka, whereas those of Cholistan dunes are
536 closest to samples 6 and 12 from the Marot borehole, dated as ~7 and > 49 ka ([Clift et al.,](#)
537 [2012](#)). We conclude that eolian sands accumulated in the Thar Desert largely before ~7 ka,
538 most plausibly during the early Holocene when summer monsoon winds were at a maximum
539 ([Gupta et al., 2003](#)).

541 **8. Climatic changes associated with the Asian Monsoon**

542 *8.1. Aeolian sedimentation since the Last Glacial Maximum*

543 A chronology of the evolution of the Thar Desert based on TL dating of aeolian sediments
544 ([Singhvi et al., 2010](#)) revealed multiple phases of dune accretion since 200 ka, interrupted by
545 hiatuses associated with precession-driven climate. In the eastern Thar Desert, the last major
546 phase of dune growth started between 17 ka and 14 ka and lasted until 9 ka, at the onset of the
547 early Holocene wet stage ([Dhir et al., 2010, 2012](#); [Singhvi et al., 2010](#); [Garzanti et al., 2020b](#)).
548 During this time, the region experienced a transitional climate, with southwesterly monsoon
549 winds strengthened after the LGM, characterized by a weak summer monsoon. Mesolithic
550 artifacts dated to the first millennium of the Holocene have been found on top of sand dunes in
551 both eastern Thar and Thal deserts, suggesting similar accretion histories ([Biagi et al., 2019](#)).
552 In contrast, the western Thar Desert (subject of this study) was supplied with sediment since
553 the beginning of the wetter early Holocene and expanded further westward as the climate dried
554 after the mid-Holocene ([East et al., 2015](#)).

555 The incision of moraines and fluvial terraces in the high Karakorum, Kohistan, and
556 Himalayan mountains led to increased sediment fluxes during the deglaciation period that
557 followed the LGM ([Blöthe et al., 2014](#); [Garzanti et al., 2020b](#); [Clift and Jonell, 2021](#)). During
558 the LGM and early deglacial period, summer monsoon rains were weak but meltwater fluxes
559 from shrinking mountain glaciers were high, and sediment was carried to the lower reaches and
560 delta, and accumulated on the northern floodplains ([Saini et al., 2009](#); [Giosan et al., 2012](#)). As
561 monsoon rains strengthened in the Early Holocene, fluvial sediment fluxes increased but
562 sedimentation became focused in the lower reaches and delta through the Holocene, whereas
563 incision prevailed in the northern floodplains after 10 ka ([Giosan et al., 2012](#)). Floodplain and
564 deltaic sediments were deflated by summer monsoonal winds and blown throughout the

1 565 Holocene into the western Thar Desert but mostly only from 17 to 9 ka in the eastern Thar
2 566 Desert (East et al., 2015; Singhvi et al., 2010).
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6 7 568 ***8.2. Paleoclimatic effects of the Asian Monsoon***

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10 569 The intensity of the Asian Monsoon varies over long as well as millennial and even
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12 570 shorter timescales (Clemens and Prell, 2003; Jonell et al., 2018). The Quaternary period
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14 571 experienced significant variations in monsoon intensity, which were influenced by insolation
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16 572 maxima and the export of latent heat from the ocean (Clemens et al., 1991; Clark et al., 2006;
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19 573 Caley et al., 2011). The relationship between monsoon intensity and orbital forcing is strong,
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22 574 but with a delayed response. Thar Desert climate in the Holocene was highly sensitive to
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24 575 changes in Asian Monsoon strength, as evidenced by the presence of an extinct river in the
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27 576 Cholistan Desert and of related fluvial deposition in the Ghaggar-Hakra River upstream
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29 577 (Oldham, 1886; Durcan, 2012; Alok et al., 2023). Relative aridity during the LGM (23–20 ka)
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32 578 was followed by an increase in monsoon intensity in the early Holocene (Durcan, 2012;
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34 579 Fleitmann et al., 2003), and enhanced regional precipitation led to incision of the Sutlej River
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36 580 into the flood plains after 10 ka. Wet conditions associated with strong winds favored sediment
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39 581 supply, deflation, and rapid accumulation in the central Thar Desert through most of the
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41 582 Holocene (Singhvi et al., 2010; Chatterjee and Ray, 2017). Around 4.5 ka, monsoonal rains
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44 583 and consequently fluvial activity in the Ghaggar-Hakra River weakened (Durcan et al., 2019),
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46 584 and along western edge of the desert dunes started to advance over the floodplain (East et al.,
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49 585 2015; Durcan et al., 2019).
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52 53 587 **9. Conclusions**

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56 588 This study employs an integrated provenance approach, combining bulk petrography,
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58 589 heavy minerals, and detrital zircon dating to constrain the origin of sediments in the western
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1 590 Thar Desert. Eolian dunes consist of feldspatho-litho-quartzose metamorphiclastic sand in the
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3 591 northern Cholistan region and of litho-feldspatho-quartzose sedimentaelastic sand in the
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5 592 southern Sindh region. The quartz/feldspar ratio is much higher in Cholistan sand, largely
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7 593 supplied by Punjabi tributaries sourced in the Himalayan belt (chiefly the Sutlej-Beas River
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9 594 system), than for Sindh sand showing notably closer affinity with Lower Indus River sand.
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12 595 This is confirmed by heavy-mineral suites, which contain notably higher percentages of green
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14 596 to brown amphibole in Sindh dunes than in Cholistan dunes. U-Pb age signatures of detrital-
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16 597 zircons are not equally distinct, with slightly more Neoproterozoic to Paleozoic ages recorded
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19 598 in Sindh than in Cholistan. Young zircons yielding Cretaceous to Miocene ages shed by
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22 599 Ladakh-Kohistan arc and Karakorum batholiths occur both in Cholistan and Sindh, indicating
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24 600 extensive mixing of zircon grains originally derived from the Upper Indus but likely
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27 601 transported into the desert from the lower reaches and delta by monsoon winds.

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29 602 Compositional affinities between Thar Desert dunes and sand presently transported by
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31 603 the modern Indus and Sutlej rivers are limited. Conversely, Sindh dunes are closest to Indus
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34 604 Delta sediments dated between ~7 and 14 ka, whereas Cholistan dunes are closest to samples
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36 605 from the Marot borehole in Cholistan dated as ~7 and > 49 ka. We thus infer that eolian sand
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39 606 largely accumulated before ~7 ka, most plausibly during the wet and windy early Holocene
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41 607 stage, when both sediment supply and summer monsoon winds reached maximum and sand
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44 608 transport into the Thar Desert was thus particularly efficient.

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618 **Acknowledgments**

619 The senior author would like to thank the PhD program of the Department of Earth and
620 Environmental Sciences of Milano Bicocca University for supporting his research and the
621 International Association of Sedimentology for receiving a postgraduate grant for detrital-
622 zircon dating. The Charles T. McCord Jr Chair in Petroleum Geology at LSU supported some
623 of this research.

624

625 **SUPPLEMENTARY MATERIALS**

626 Supplementary data associated with this article, to be found in the online version at
627 http://dx.doi._____, include information on sampling sites (Table A1), together with
628 the complete datasets on sand petrography (Table A2), heavy minerals (Table A3), original and
629 literature data on U-Pb zircon-age distributions (Table A4), percentages of U-Pb zircon-age
630 distributions in diverse source-rock domains (Table A5), *DZMix* results summary (Table A6),
631 and *DZStats* input dataset (Table A7). The Google-EarthTM map of sampling sites [Thar.kmz](#) is
632 also provided.

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634

635 FIGURE and TABLE CAPTIONS

636

637 **Figure 1.** Geological map of the western Himalaya syntaxis and other ranges drained by the
638 Indus River system, showing the major tectonic blocks of these regions which form the primary
639 sources of sediment to the Indus River modified after [Garzanti et al. \(2020\)](#). Location of
640 modern sand samples from both Thar Desert and Indus drainage system are shown.

641

642 **Figure 2.** Satellite image of Pakistan and adjacent regions (from Google Earth™). Figure
643 shows sampling locations for desert sand and modern rivers. Note the distribution of the
644 sediments collected during this study in Cholistan and further south in eastern Sindh. Original
645 image from Google Earth. Location of modern sand samples from both Thar Desert and Indus
646 drainage system are shown.

647

648 **Figure 3.** Photomicrographs of representative sediment samples from both Sindh and Cholistan
649 deserts. Grains appear slightly more angular in Sindh sand. Q = Quartz, P = plagioclase, K =
650 K-feldspar, Ls = sedimentary lithic, Lms = metamorphic lithic, c = carbonate, Amp =
651 amphibole, Ep = Epidote.

652

653 **Figure 4.** Petrography of Thar Desert sand compared with sand from potential source regions.
654 A) QFL diagram (compositional fields after [Garzanti, 2019](#)); B) QFP diagram; C) Ls-Lm-Lv
655 diagram ([Ingersoll et al., 1984](#)). Note greater affinity of Cholistan sand with Punjab tributaries
656 sourced in the Himalayan belt and of Sindh sand with Indus River sand. D) Heavy mineral
657 discrimination diagram ([Morton et al., 1999](#)). E) Principal component plot ([Vermeesch et al.,](#)
658 2016). Data from contributing sediment sources from [Garzanti et al. \(2020\)](#). Q = quartz; F =
659 feldspars (P = plagioclase; KF = K-feldspar); L = lithic fragments (Lm = metamorphic; Ls =

660 sedimentary) Amp = amphibole; Ep = epidote; Grt = garnet; Ky = kyanite; Px = pyroxene; Sill
661 = sillimanite; St = staurolite.

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663 **Figure 5.** Detrital-zircon geochronology of five studied samples compared with literature data
664 from potential sources. Left and center panels: KDE diagrams for U-Pb age spectra in modern
665 fluvial and eolian sands from the Indus catchment. Four pie diagrams to the right compare age
666 signatures of Lower Indus and Punjabi tributaries with eolian sands of Cholistan and Sindh.

667
668 **Figure 6.** Multidimensional Scale (MDS) diagram comparing U-Pb age spectra of detrital
669 zircons in Cholistan and Sindh dunes with potential sources.

670
671 **Figure 7.** U-Pb age spectra of detrital zircon in the studied Thar Desert samples plotted as KDE
672 and compared with potential sources throughout the Indus catchment. Data sources cited in
673 text.

674
675 **Figure 8.** Pie diagrams depict the results of provenance budgets obtained by forward mixing
676 calculations based on heavy-mineral assemblages. A) Lower Indus; B) Upper Indus; C) Sindh
677 dunes with tectonic domains as end members; D) Cholistan dunes with tectonic domains as
678 end members; E) Sindh dunes with rivers as end members; F) Cholistan dunes with rivers as
679 end members. Data sources from [Garzanti et al. \(2005, 2020\)](#).

680
681 **Figure 9.** Pie diagrams depict the results of provenance budgets based on detrital zircon U-Pb
682 spectra and estimated using *DZMix* software (Sundell and Saylor, 2017; plots based on V factor
683 of Kuiper test, full results provided in [Appendix Table A5](#)). A) Sindh zircons; B) Cholistan
684 zircons. C) Upper Indus zircons ([Alizai et al., 2011](#); [Clift et al., 2022](#)). D) Lower Indus zircons

685 (Clift et al., 2004). E, F) Indus delta at 7 ka and 14 ka (Clift et al., 2008). G, H) Paleo-Ghaggar-
686 Hakra channel at 7 ka and <49 ka (Marot borehole; Clift et al., 2012). Data on desert sands
687 from East et al. (2015) are included. Note greater zircon contribution from Punjabi tributaries
688 sourced in the Himalayan belt in Cholistan than in Sindh.

689

Figure 10. Heat map (blue = similar; red = dissimilar) comparing U-Pb zircon signatures of
690 Thar Desert sands with selected fluvial samples (as determined by K-S testing using the
691 *DZStats* program of Saylor and Sundell, 2016; details provided in Appendix Table A5). Note
692 affinities between modern Cholistan dunes with Marot channel sands older than 7 ka, and
693 between modern Sindh dunes with Indus Delta sediments older than 7 ka.

695

Table 1. Petrography and heavy minerals composition of Thar (Cholistan and Sindh) Desert
696 sand carried by the Indus River and Punjabi tributaries from the different NW Himalayan
697 sources. Q = quartz; F = feldspars (KF = K-feldspar; P = plagioclase; L = lithic grains (Lvm
698 = volcanic and metavolcanic; Lc = carbonate and metacarbonate; Lh = chert; Lsm = shale,
699 siltstone, slate, and metasiltstone; Lmf = felsic metamorphic; Lmb = metabasite; Lu =
700 ultramafic); HM = heavy minerals; MI* = Metamorphic Index; tHMC = transparent heavy-
701 mineral concentration. ZTR = zircon + tourmaline + rutile; Ttn = titanite; Ep = epidote-group
702 minerals; Grt = garnet; SKS = staurolite + kyanite + sillimanite; Amp = amphibole; Px =
703 pyroxene (Cpx = clinopyroxene; Opx = orthopyroxene, mostly hypersthene); & tHM = other
704 transparent heavy minerals (apatite, chloritoid, Cr-spinel, olivine, prehnite, pumpellyite,
705 brookite, andalusite, baryte).

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708 **Table 2.** An average age structure of the zircon U-Pb ages from Sindh and Cholistan,
709 together with the lower Indus and an average of all the Punjabi tributaries. *Italic:* for standard
710 deviations.
711
712 **Table 3. (a).** The contribution of the heavy minerals into the Indus (Lower and Upper) River
713 and Thar (Cholistan and Sindh) Desert from different NW Himalaya sources. (b). Sediments
714 supply as heavy minerals into the Lower Indus and Thar (Cholistan and Sindh) from the
715 Punjabi tributaries.
716
717 **Table 4.** Detrital Zircon contribution in percentages in the Thar (Sindh and Cholistan) Desert
718 and Indus River with NW Himalayas and Punjabi tributaries
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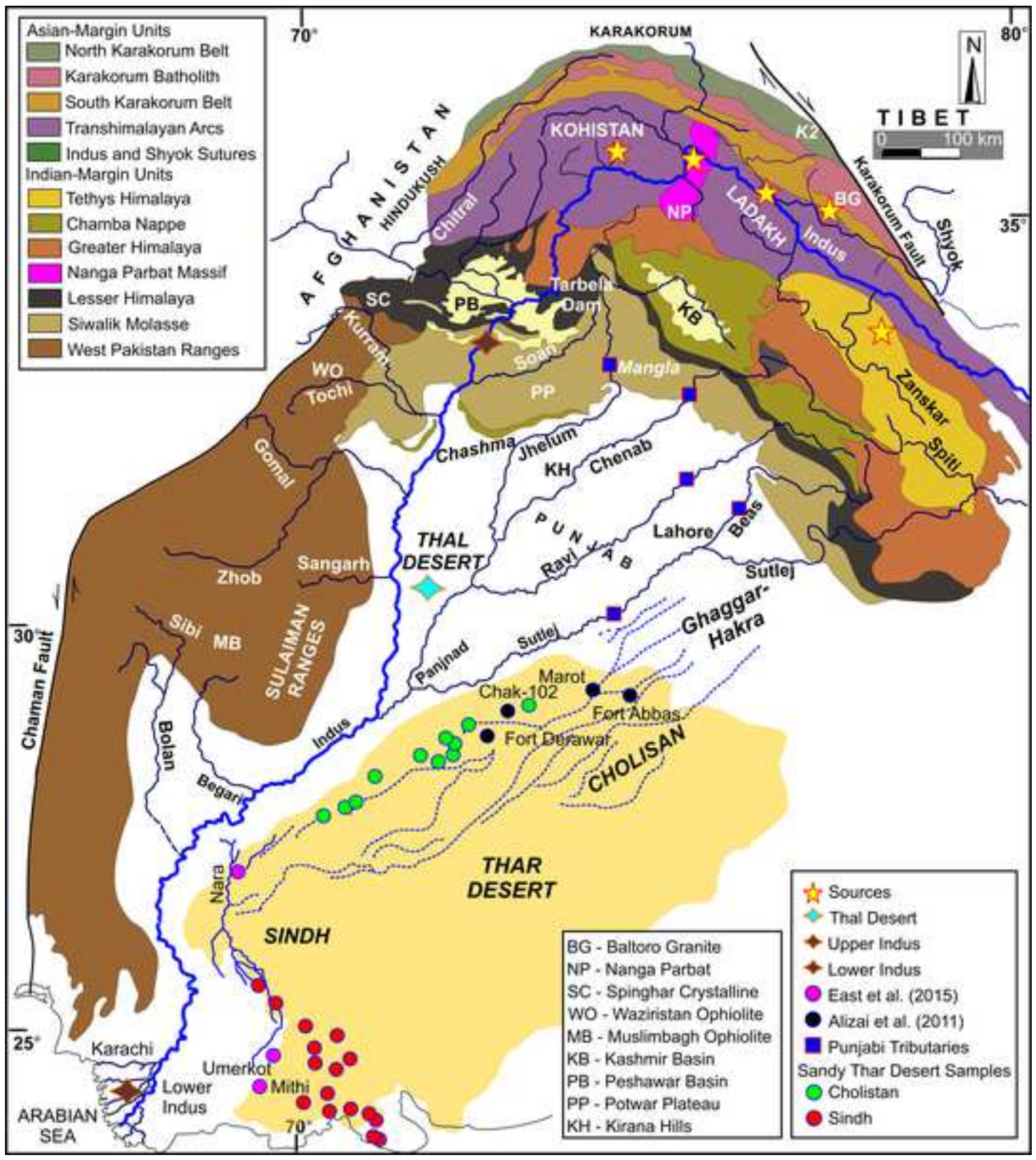
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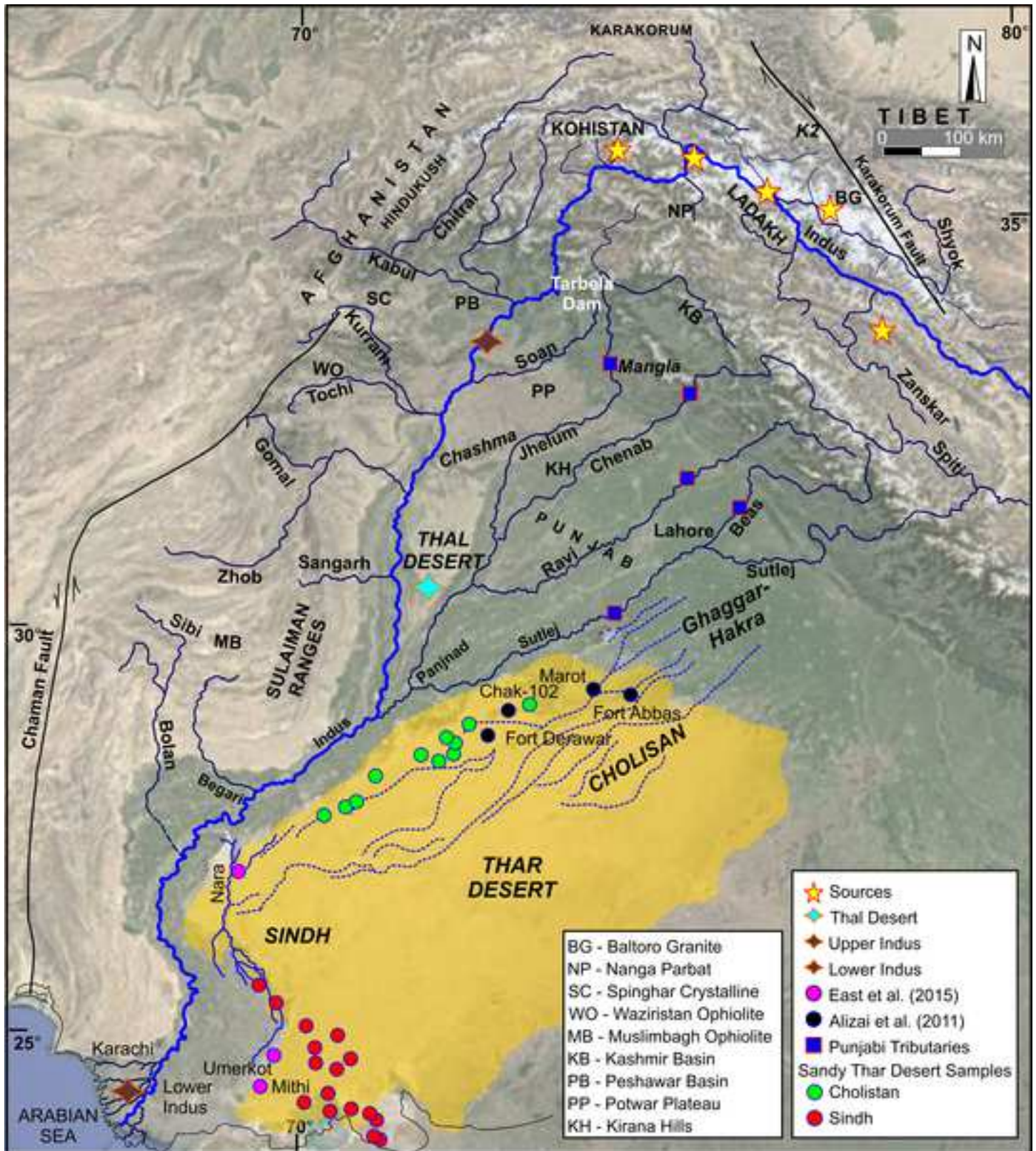
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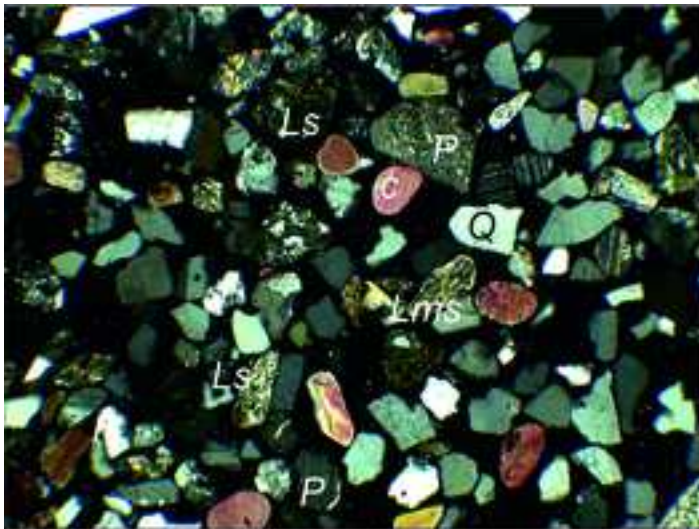
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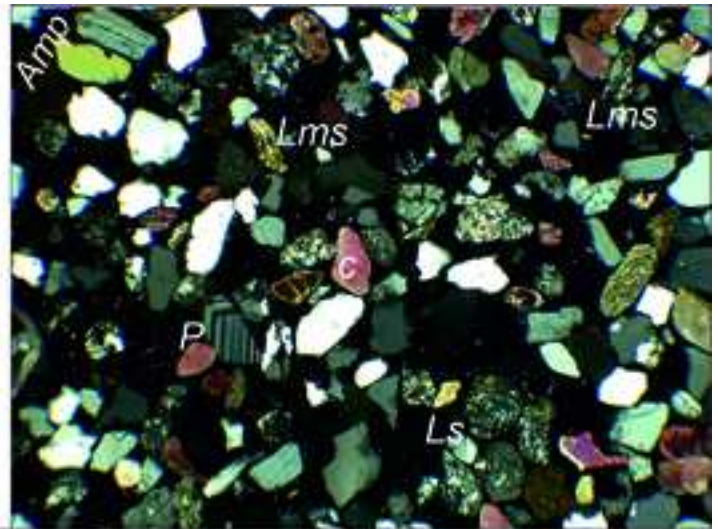
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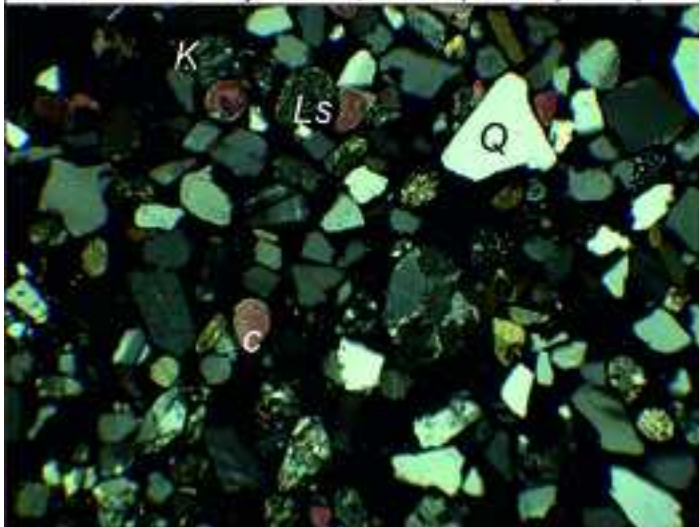




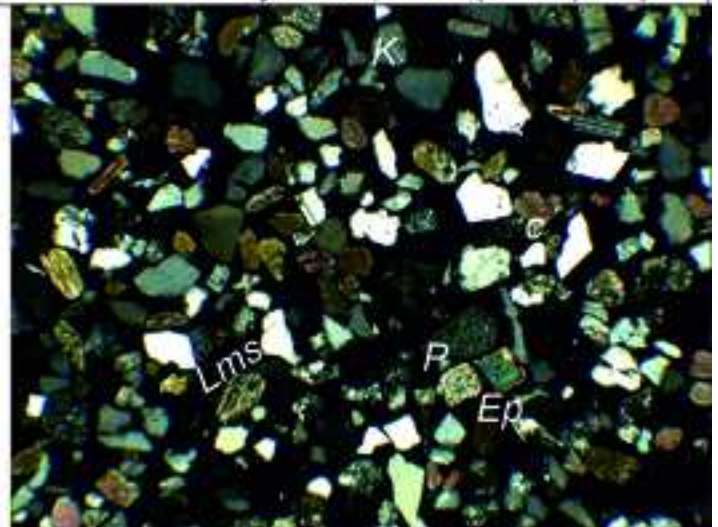
A Thar Sandy Desert, Sindh (S5896) 100μm



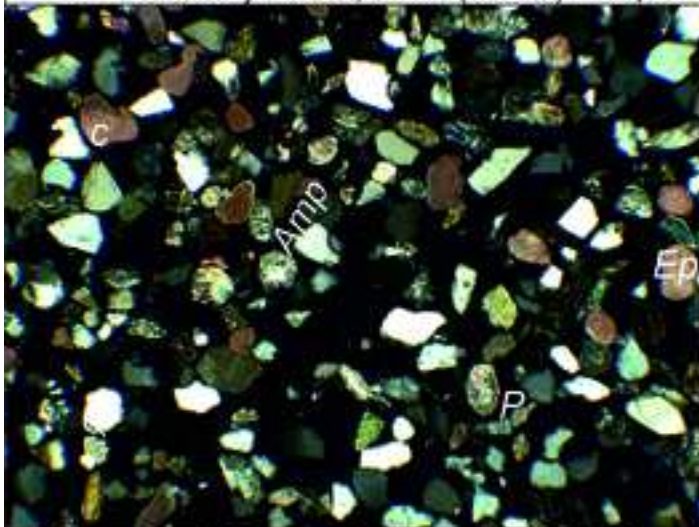
B Thar Sandy Desert, Sindh (S5986) 100μm



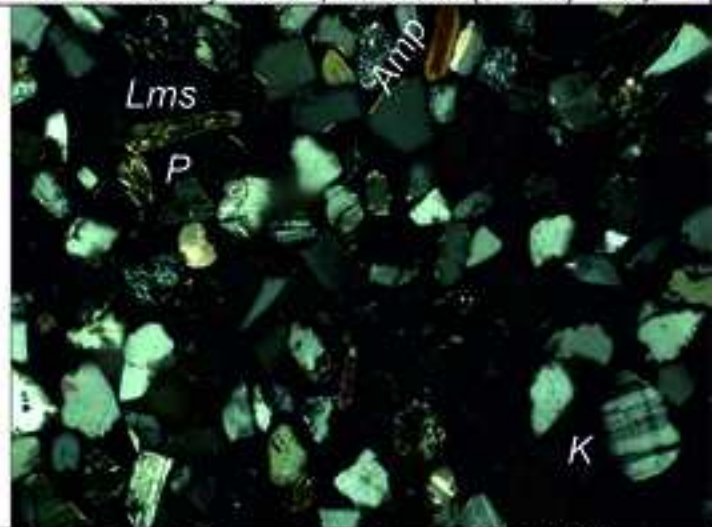
C Thar Sandy Desert, Sindh (S6002) 100μm



D Thar Sandy Desert, Cholistan (S6112) 100μm



E Thar Sandy Desert, Cholistan (S6126) 100μm



F Thar Sandy Desert, Cholistan (S6142) 100μm

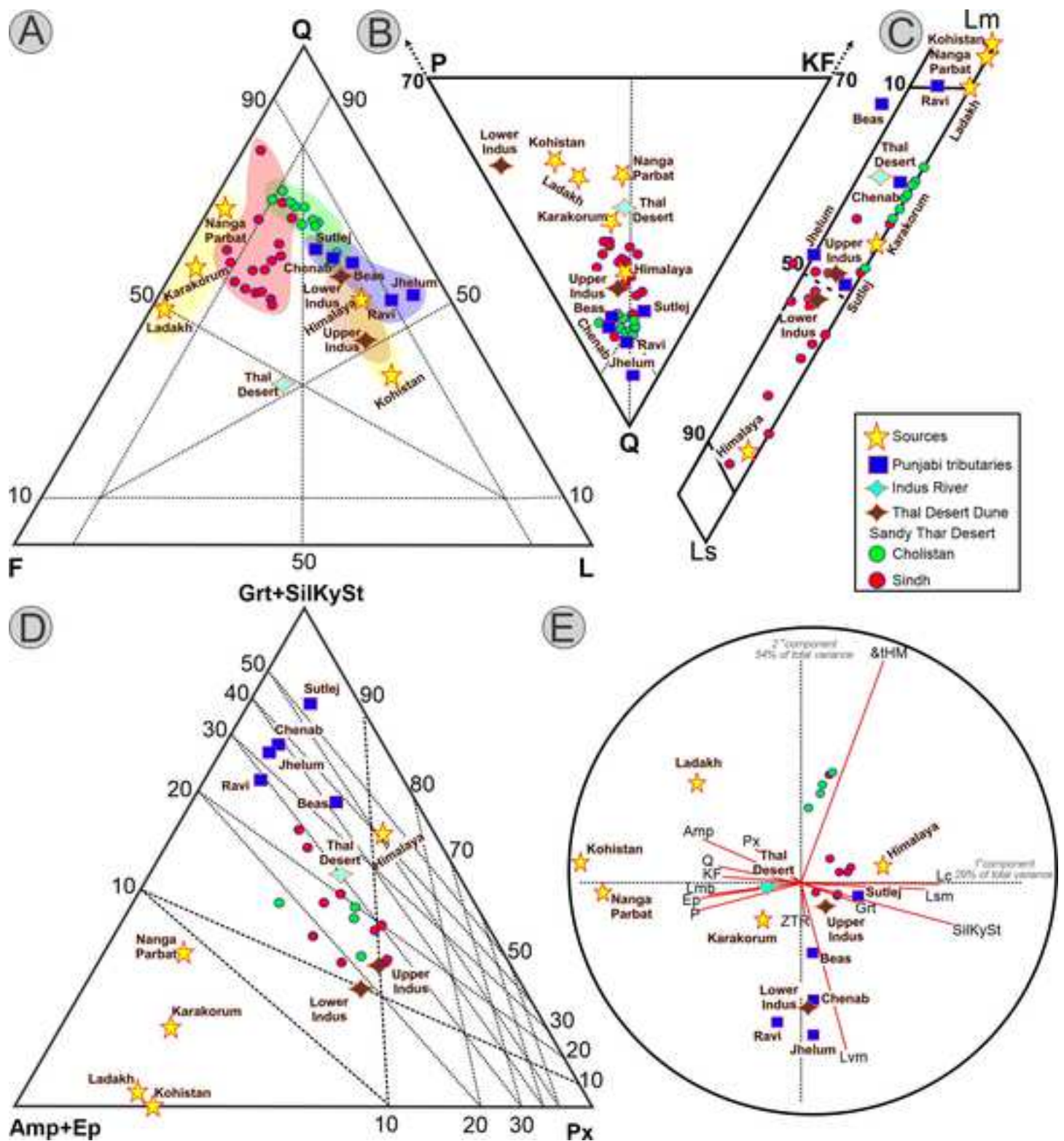


Figure 5

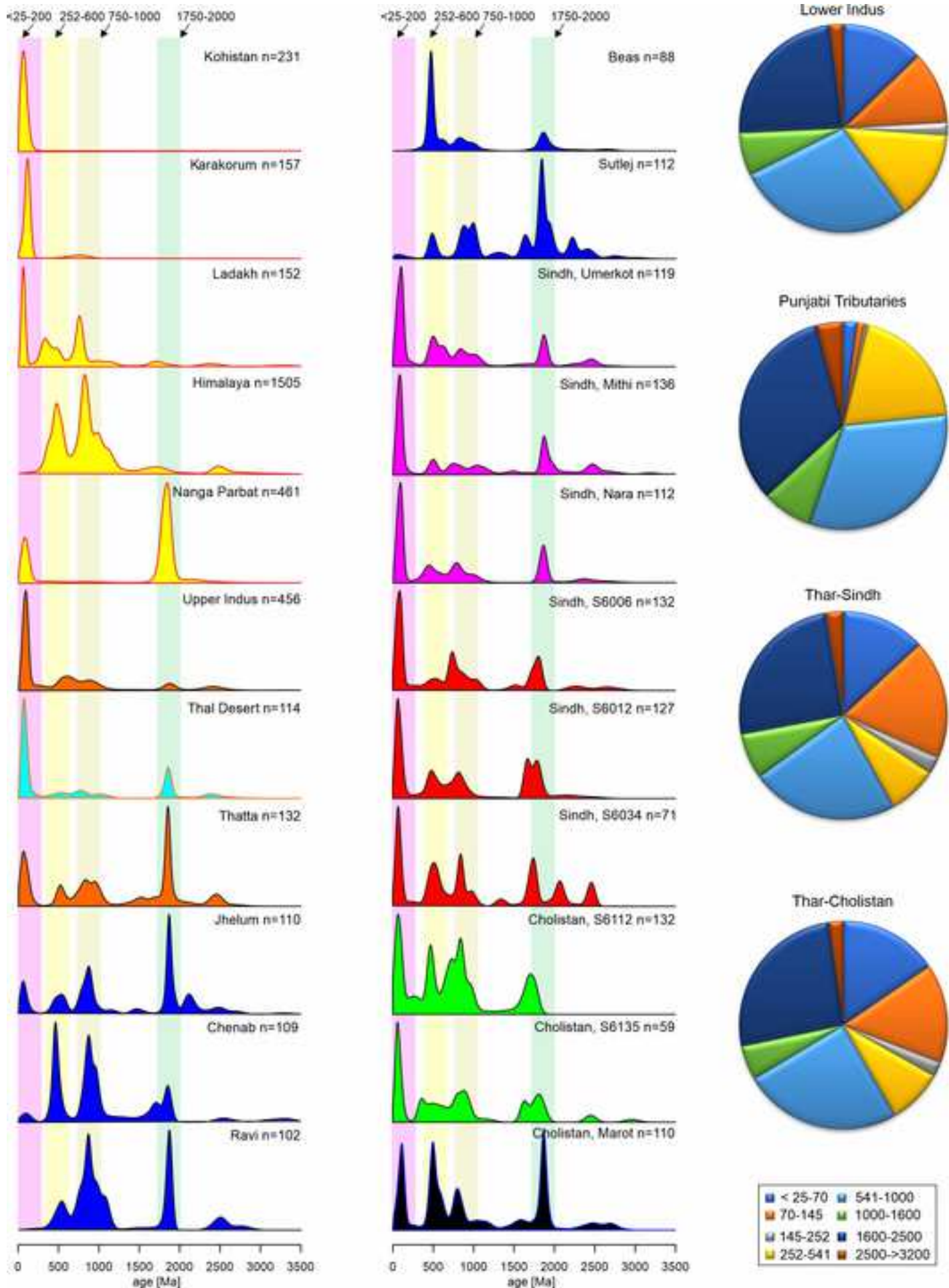


Figure 6

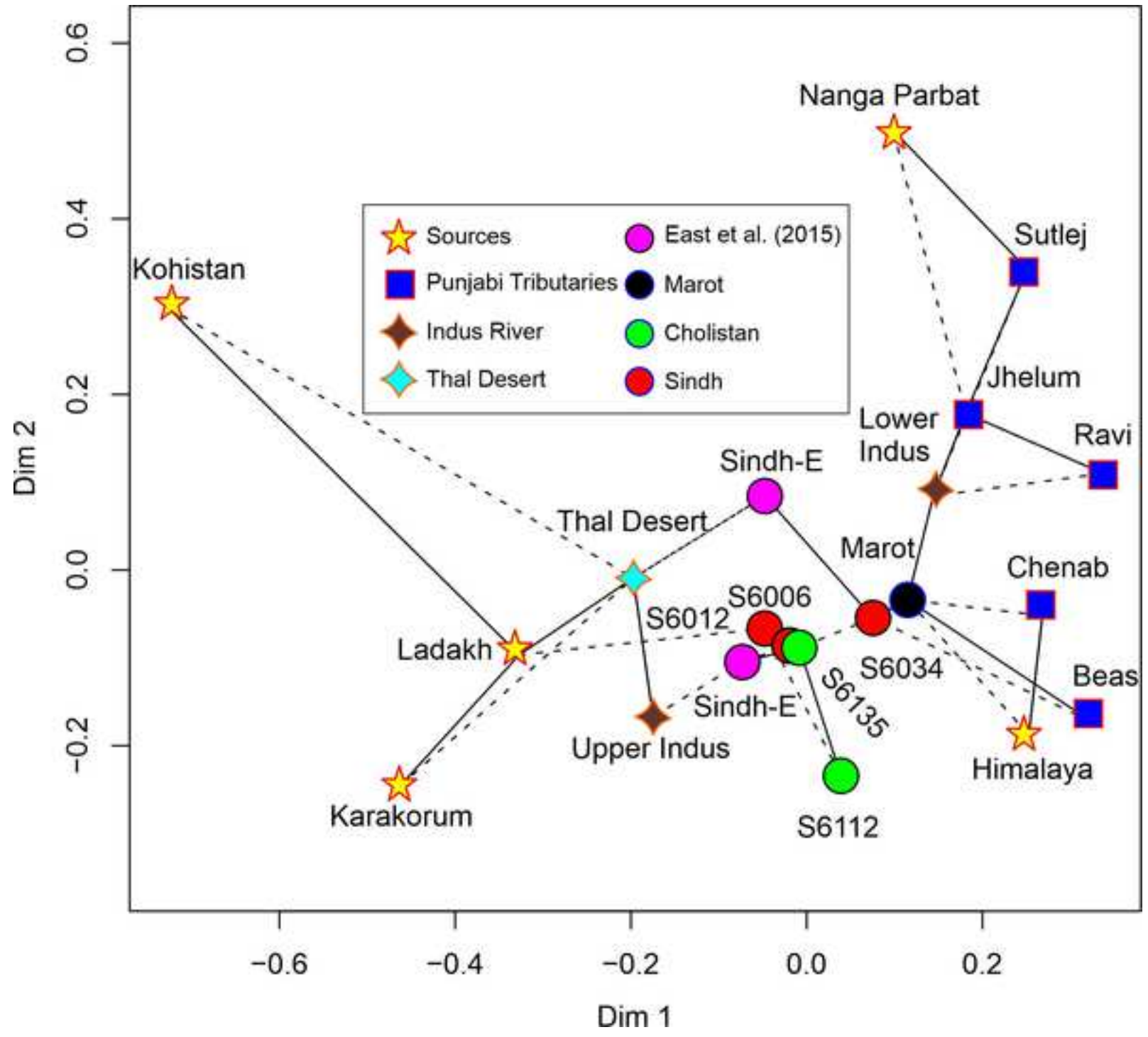
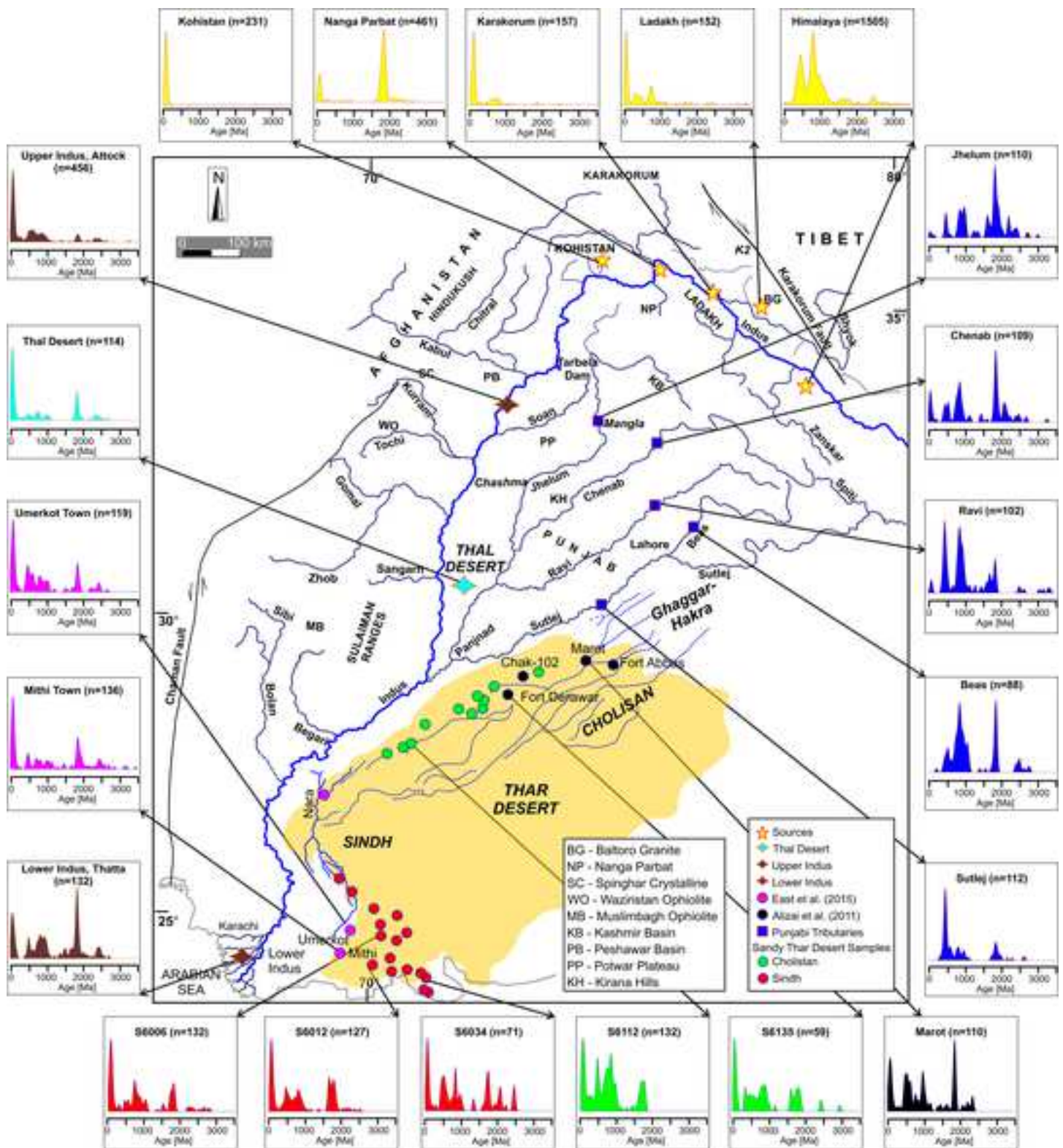
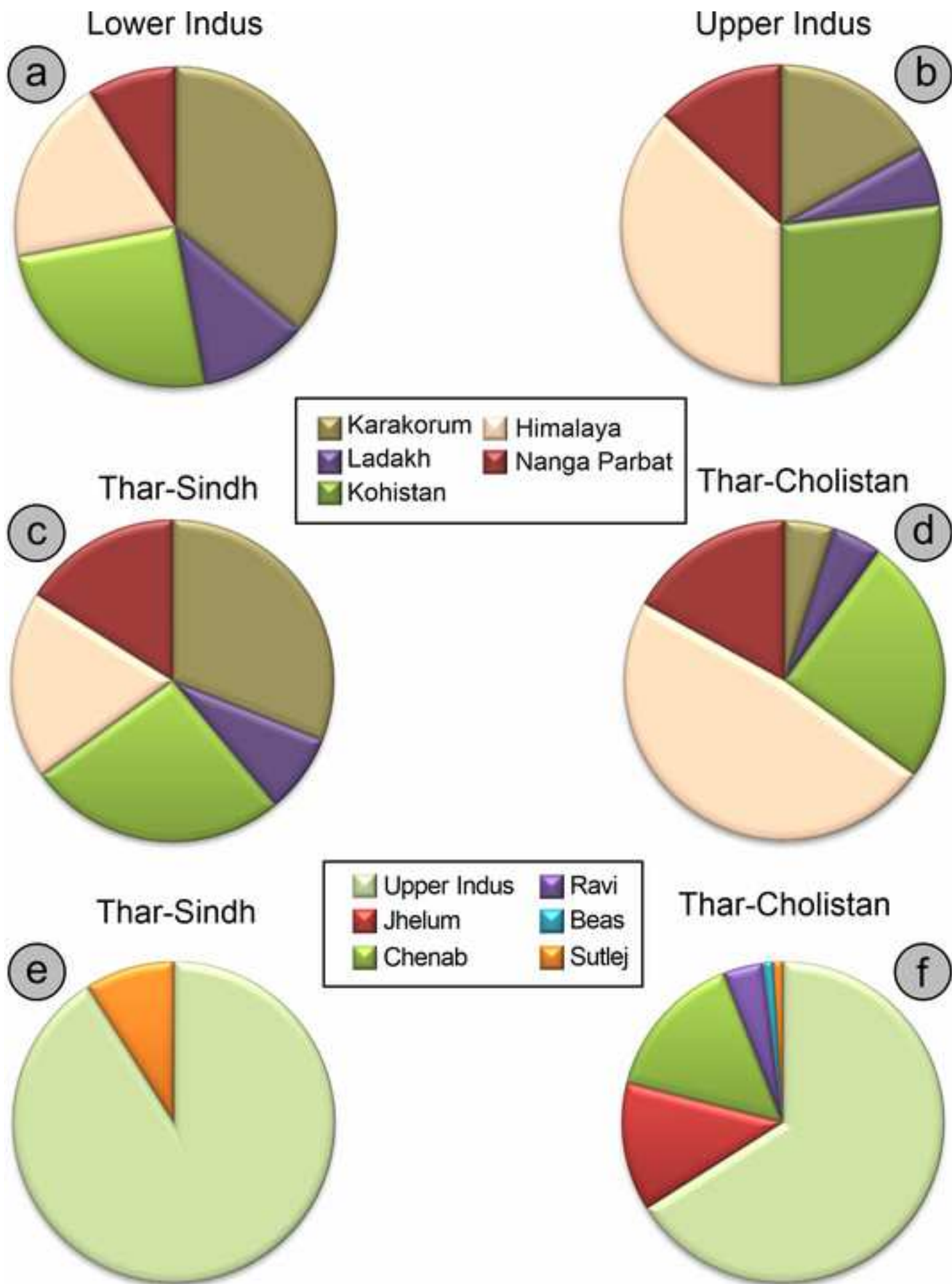


Figure 7

Click here to access/download;Figure;Figure 7. KDEs with Map.jpg





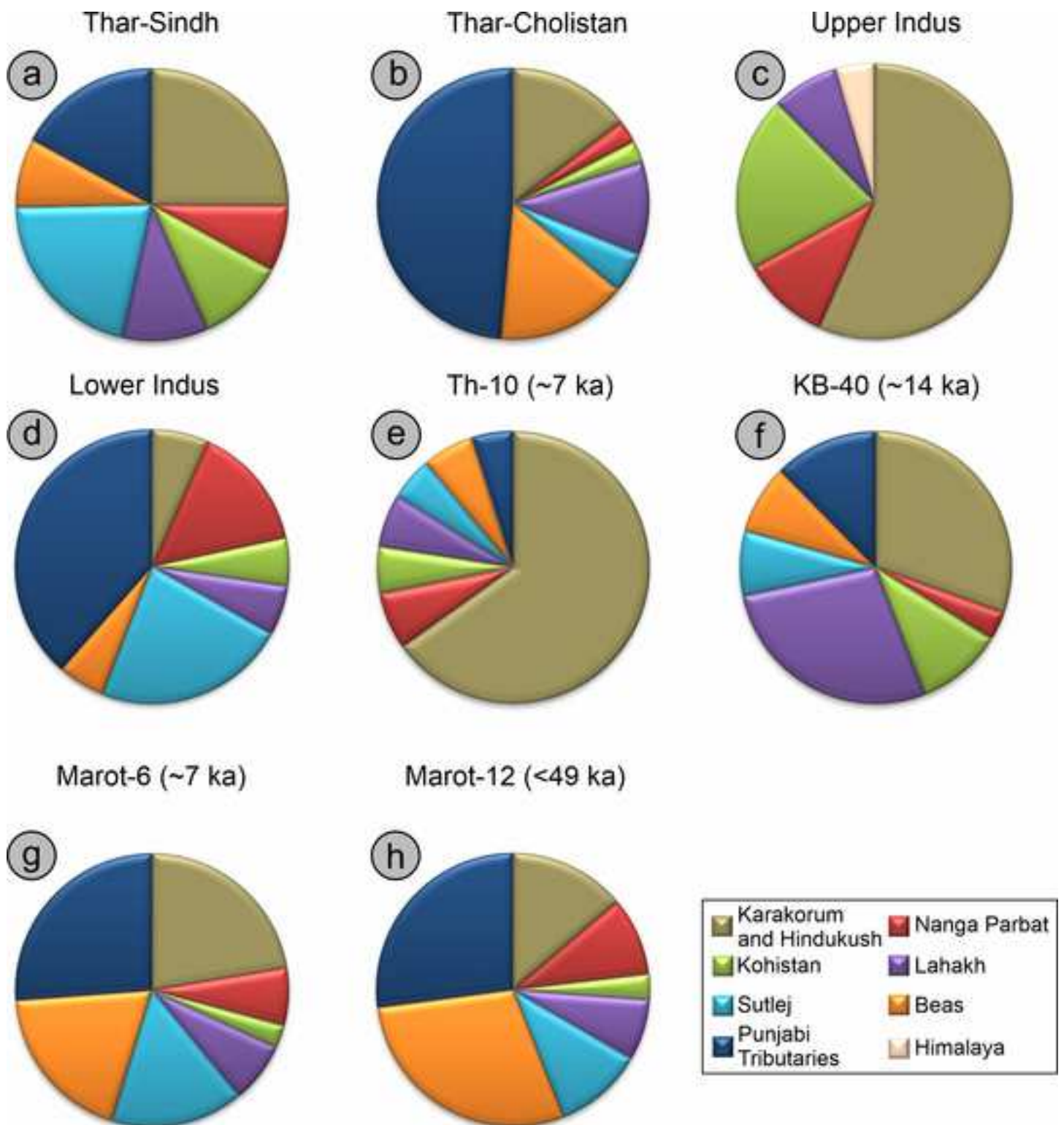


Table 1.		minerals composition of Thar (Cholistan and Sindh) Desert sand carried by the Indus River and Punjabi tributaries from the different																						
River/Desert	Sample	Q	F	Lv	Lc	Lh	Lp	Lm	Lu	total	q/p/Q	P/F	MI*	tHMC	ZTR	Ttn	Ep	Grt	SKS	Amp	Cpx	Opx	&tHM	total
Thar Sindh We	Average	50	30	1	10	0	1	7	0	100.0	14	48	190	6	1	1	16	16	6	54	3	2	1	100.0
	Stand De	0	0	1	1	0	0	1	0		7	2	22	1	1	1	3	0	3	5	2	0	0	
Thar Sindh Cer	Average	58	27	1	6	0	2	6	0	100.0	10	58	161	5	2	0	15	13	2	59	6	2	1	100.0
	Stand De	5	5	1	1	0	1	1	0		4	7	20	1	0	1	3	4	2	5	1	1	1	
Thar Sindh Sou	Average	58	28	0	7	0	2	5	0	100.0	8	54	167	6	3	2	11	13	4	61	4	1	2	100.0
	Stand De	5	5	0	3	0	0	2	0		2	9	17	3	3	2	1	4	1	3	1	1	1	
Thar Cholistan	Average	62	17	0	7	0	1	13	0	100.0	5	50	268	15	3	2	26	9	6	44	4	4	2	100.0
	Stand De	3	1	0	2	0	1	2	0		2	4	20	2	1	1	5	4	0	4	1	1	0	
Thar Cholistan	Average	67	19	0	5	0	0	9	0	100.0	8	58	319	10	4	4	29	10	2	46	3	0	1	100.0
	Stand De	1	2	0	1	0	1	1	0		3	5	13											
Thar Cholistan	Average	67	18	0	6	0	0	9	0	100.0	8	51	309	11	2	2	19	10	6	50	5	2	3	100.0
	Stand De	4	1	0	4	0	0	2	0		1	4	12											
Thal	S1462	37	34	1	7	0	3	18	0	100.0	9	51	292	15	2	1	17	12	2	56	5	4	0	100.0
Sources																								
Karakorum	S1749	55	41	0	2	0	0	3	0	100.0	6	59	405	2	7	10	11	2	1	67	1	0	0	100.0
Ladakh	S4430	47	50	0	0	0	1	2	0	100.0	18	68	403	8	2	2	7	0	0	86	0	1	0	100.0
Kohistan	S1439	32	18	0	2	0	0	48	0	100.0	17	88	320	33	0	0	38	0	0	60	1	0	0	100.0
Swat-Kohistan	S1440	26	47	2	0	0	0	21	3	100.0	14	82	339	28	1	0	8	0	0	67	8	15	0	100.0
Himalaya	S4419	49	15	0	29	0	1	6	0	100.0	3	60	356	5	6	1	12	14	23	31	10	0	4	100.0
Nanga Parbat	S1432	67	30	0	0	0	0	3	0	100.0	18	50	390	17	1	0	19	7	0	71	1	0	0	100.0
Indus																								
Upper Indus	S1447/14	43	21	1	12	2	7	14	1	100.0	12	54	281	10	1	2	22	11	3	52	5	3	1	100.0
Lower Indus	S1489	53	17	1	12	1	5	11	0	100.0	5	54	364	9	2	1	17	6	3	61	6	2	0	100.0
Punjab tributaries																								
Jhelum	S1449	50	6	3	12	3	10	16	0	100.0	17	39	216	9	4	1	55	30	2	7	1	0	0	100.0
Chenab	S1450	58	16	0	5	0	8	13	0	100.0	8	42	261	1	2	0	31	19	16	30	1	0	0	100.0
Ravi	S1451	49	10	1	1	1	17	20	0	100.0	18	50	198	1	14	0	42	20	3	18	1	0	0	100.0
Sutlej	S1467	59	17	0	8	0	6	9	0	100.0	7	38	266	6	10	0	8	36	15	26	1	1	0	100.0
Beas	S2284	55	15	4	3	1	3	20	0	100.0	21	47	164	1	13	1	41	26	4	8	5	0	1	100.0
<p>Q = quartz; F = feldspars (KF = K-feldspar; P = plagioclase; L = lithic grains (Lvm = volcanic and metavolcanic; Lc = carbonate and metacarbonate; Lh = chert; Lsm = shale, siltstone, slate, and metasilstone; Lmf = felsic metamorphic; Lmb = metabasite; Lu = ultramafic); HM = heavy minerals; MI* = Metamorphic Index; tHMC = transparent heavy-mineral concentration. ZTR = zircon + tourmaline + rutile; Ttn = titanite; Ep = epidote-group minerals; Grt = garnet; SKS = staurolite + kyanite + sillimanite; Amp = amphibole; Px = pyroxene (Cpx = clinopyroxene; Opx = orthopyroxene, mostly hypersthene); & tHM = other transparent heavy minerals (apatite, chloritoid, Cr-spinel, olivine, prehnite, pumpellyite, brookite, andalusite).</p>																								

Table 2.		Structure of the zircon U-Pb ages from Sindh and Cholistan, together with the lower Indus and an average of all the												
	Era	Mioc	Oli/Eocer	Paleoc	Up-C	Lw-C	J	Tr	P	Neop	Mesop	Paleop	Arc	
	Ages	<25	25-50	50-70	70-99	99-145	145-201	200-252	252-541	541-1000	1000-1600	1600-2500	2500->3200	
Cholistan	259	1.5%	8.9%	5.0%	5.4%	10.0%	1.2%	1.2%	8.5%	24.7%	5.3%	26.0%	2.3%	
		0.0%	0.3%	1.8%	3.1%	0.8%	0.6%	0.5%	3.5%	2.6%	4.2%	7.8%	2.1%	
Sindh	582	3.2%	3.9%	5.7%	3.4%	7.8%	1.5%	0.6%	13.8%	27.7%	6.7%	23.6%	2.1%	
		0.3%	1.2%	1.8%	1.8%	2.9%	0.4%	0.7%	2.3%	4.8%	1.8%	3.7%	1.7%	
Indus River	588	1.5%	4.8%	6.8%	7.6%	11.0%	1.8%	0.9%	7.6%	23.0%	7.1%	25.2%	2.7%	
		1.1%	2.5%	2.1%	8.5%	7.0%	0.7%	1.2%	2.2%	0.4%	3.9%	17.9%	0.4%	
Punj Trib	521	0.0%	1.3%	0.9%	0.5%	0.5%	0.2%	0.4%	19.5%	31.9%	7.9%	32.6%	4.2%	
		0.0%	2.0%	1.6%	0.8%	0.8%	0.4%	0.5%	17.9%	10.9%	4.1%	16.8%	1.9%	

Table 3a. The contribution of the heavy minerals into the Indus (Lower and Upper) River and					
Sources	Samples	Upper Indus	Lower Indus	Sindh	Cholistan
Karakorum	S1749	17%	36%	31%	5%
Ladakh	S4430	6%	11%	8%	5%
Kohistan	S1440	27%	25%	26%	25%
Himalaya	S4419	37%	19%	19%	48%
Nanga Parbat	S1432	13%	9%	16%	17%

Table 3b. Sediments supply as heavy minerals into the Lower Indus and Thar				
Sources	Samples	Sindh	Cholistan	Lower Indus
Lower Indus	Average	91%	66%	99%
Jhelum	S1449	0%	13%	0%
Chenab	S1450	0%	15%	0%
Ravi	S1451	0%	4%	0%
Beas	S2284	0%	1%	0%
Sutlej	S1467	9%	1%	1%

Sources	Sindh	TH-10 (7 ka)	B-40 (14 ka)	Lower Indus	Cholistan	AROT-6 (7 ka)	AROT-12 (<4 ka)	Sources	Upper Indus
Karakorum	25%	65%	30%	7%	15%	22%	14%	Karakorum	57%
Nanga Parbat	8%	7%	3%	15%	3%	7%	9%	Nanga Parbat	10%
Kohistan	10%	6%	10%	6%	3%	3%	3%	Kohistan	21%
Ladakh	10%	6%	28%	6%	11%	7%	7%	Ladakh	8%
Sutlej	21%	5%	8%	23%	5%	16%	11%	Zaskar	5%
Beas	8%	6%	8%	6%	15%	19%	29%		
Punjab Tributaries	17%	5%	12%	39%	49%	26%	27%		
Total Punjab	47%	16%	28%	67%	69%	61%	67%		